

**UNIVERSIDADE DE BRASÍLIA
INSTITUTO DE GEOCIÊNCIAS**

DISSERTAÇÃO DE MESTRADO

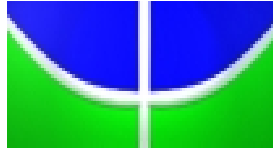
Nº 235

**“ESTUDO PETROGRÁFICO E GEOQUÍMICO DAS ROCHAS
VULCÂNICAS AFLORANTES NO EMBASAMENTO
(PROTEROZÓICO SUPERIOR-PALEOZÓICO INFERIOR) DA
CORDILHEIRA ORIENTAL, NOROESTE ARGENTINO.”**

**Autora:
NATALIA HAUSER**

**Orientador:
Prof. Dr. Márcio Martins Pimentel**

**BRASILIA-DF
DEZEMBRO DE 2007**



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**BRASILIA-DF
DEZEMBRO DE 2007**

*Toda porção de verdade deve encaixar em qualquer outra porção,
porque todas elas formam parte do todo.*

Atkinson Thomas Willian

*Quando se sabe uma coisa, sustentar que se sabe; e, quando não se
sabe, admitir que não sabe. Este é o verdadeiro conhecimento.*

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RESUMO

O embasamento da Cordilheira Andina no noroeste de Argentina registra pelo menos duas orogênias paleozóicas: a Orogênese Pampeana (~520 Ma) e a Orogênese Famatiniana (~460 Ma).

A orogênese Pampeana foi registrada mais extensivamente na parte noroeste das Sierras Pampeanas, na Cordilheira Oriental e no altiplano da Puna. Nas Sierras Pampeanas, a orogênese foi caracterizada pela formação de um arco magmático relacionado à subducção de litosfera oceânica durante o Cambriano Inferior, seguida de colisão continental (535-520 Ma) de um terreno semiautoctone (bloco Pampia ou o Craton Arequipa-Belen-Antofalla). O Ciclo orogênético Famatiniano, ocorrido desde 490 até 400 Ma, foi documentado no Sistema de Famatina, na Cordilheira Oriental e no altiplano da Puna. Registra a subducção de litosfera oceânica durante o Paleozóico Inferior com a acreção de um terreno exótico, a Precordilheira, aos 460 Ma aproximadamente.

No noroeste argentino a orogênese Pampeana é documentada na Cordilheira Oriental e no setor leste do Altiplano da Puna como uma seqüência turbidítica de margem passiva denominada Formação Puncoviscana. Rochas plutônicas, com idades de ca. 530 Ma, intrudem a Formação já deformada, marcando o final da orogênese. Entanto a Orogênese Famatiniana no noroeste argentino é representada na Cordilheira Oriental por rochas siliciclasticas de plataforma e marinhas rasas e na Puna por rochas marinhas profundas. Durante este ciclo no Altiplano da Puna são gerados dois arcos magmáticos, denominados “*Faja Eruptiva de la Puna Occidental*” e “*Faja Eruptiva de la Puna Oriental*”.

Rochas vulcânicas e sub-vulcânicas intercaladas na Formação Puncoviscana foram reconhecidas em diferentes localidades da Cordilheira Oriental e na borde leste do Altiplano da Puna: Vale do Río Grande, Serra de Niño Muerto e no Vale do Río Blanco. O presente trabalho reúne um conjunto de dados petrográficos, de química mineral, geoquímicos, e geocronológicos (idades U-Pb e K-Ar e relações isotópicas Sr-Nd) dessas rochas vulcânicas com o objetivo de classificá-las, conhecer o seu ambiente tectônico e vinculá-las por meio de suas idades à evolução da margem oeste da América do Sul.

No Vale do Río Grande, as rochas do embasamento Neoproterozoico-Paleozoico Inferior são intrudidas por diques basálticos. Os dados petrográficos, de química mineral assim como os dados geoquímicos e isotópicos permitiram classificá-las como “ocellar-analcima monchiquitos”, uma variedade de lamprofiro alcalino sem feldspato. Dados K-Ar indicaram a idade de 163 Ma, as rochas exibem baixas a médias relações iniciais isotópicas $^{87}\text{Sr}/^{86}\text{Sr}$ (0.70377-0.70781) e relativamente altas relações iniciais de $^{143}\text{Nd}/^{144}\text{Nd}$ (0.512506-0.512716), correspondendo a valores positivos de $\epsilon_{\text{Nd}}(t)$ (+1.5-+5.6) e idades modelos entre

0.25-0.64 Ga. Os novos dados permitem sugerir que os lamprofiros de Río Grande, foram gerados durante o estágio de pre-rift do Rift Intracontinental de Salta (Jurássico Superior-Eoceno Inferior). Os dados isotópicos e as evidências petrológicas e geoquímicas, junto com os dados da literatura referentes à composição do manto neste setor dos Andes Centrais, permitiram inferir que os magmas derivam de um manto metassomatizado, modificado provavelmente pela Orogênese Famatiniana durante o Paleozóico Inferior.

A área de El Niño Muerto, na borda leste do Altiplano da Puna, consiste em uma serra de direção nordeste-sudoeste, formada totalmente pelas rochas metassedimentares da Formação Puncoviscana. Rochas calci-alcálinas de alto K e composição dacítica intrudem na formação como um sill. Novos dados U-Pb por LA-ICPMS em zircões, bem como de isótopos de Sr-Nd indicam que os metadacitos de El Niño Muerto cristalizaram há 495 ± 4 Ma com altas relações iniciais de $^{87}\text{Sr}/^{86}\text{Sr}$ (0.71107-0.71180) e valores de $\epsilon_{\text{Nd}}(t)$ negativos entre -9.7 e -5.9. O caráter peraluminoso apresentado por estas rochas sugerem que o magma do qual derivam os dacitos teve forte participação de material crustal seja por contaminação ou por fusão parcial de uma crosta continental com importante componente sedimentar. A idade e as características geoquímicas e isotópicas permitem localizar estas rochas no Arco Famatiniano desenvolvido durante o Ordoviciano no noroeste da Argentina.

No vale do Río Blanco, na Cordilheira Oriental, foram reconhecidas rochas subvulcânicas, lavas almofadadas e depósitos peperíticos que se intercalam às rochas sedimentares hemipelágicas e turbidíticas reconhecidas como parte da Formação Puncoviscana. A geoquímica de elementos traço e terras raras, permitiu classificar estas rochas como basaltos e determinar que elas têm semelhanças com E-MORBs, com razões Zr/Nb desde 6.8 até 9.5. Os novos dados LA-ICPMS U-Pb obtidos em zircões indicam que os basaltos cristalizaram há 501 ± 9 Ma a partir de um magma primitivo, com valores positivos de $\epsilon_{\text{Nd}}(T)$. Eles são caracterizados por relações moderadamente altas $^{87}\text{Sr}/^{86}\text{Sr}$ (0.70713-0.70942), valores de $\epsilon_{\text{Nd}}(T)$ positivos entre +2.47 e +4.46 e idades modelo entre 0.84 e 1.12 Ga. O ambiente tectônico de formação destas rochas permanece pouco claro e as características geoquímicas indicam que eles podem ser comparados com os basaltos tipo E-MORB ou com os basaltos anorogênicos de intraplaca.

Tendo em conta que os basaltos do Vale de Río Blanco e os metadacitos de arco de El Niño Muerto foram gerados no mesmo momento, sugere-se nesta dissertação que os basaltos podem representar os remanescentes de uma bacia de retroarco, em relação ao Arco Famatiniano desenvolvido mais para o oeste durante o Cambriano Superior-Ordoviciano Inferior no noroeste de Argentina.

ABSTRACT

The basement of the Andean Cordillera in NW Argentina records at least two orogenic episodes during the Paleozoic: the Pampean Orogeny, (~520 Ma) and the Famatinian Orogeny (~ 460 Ma)

The Pampean Orogeny was registered extensively in the northwestern Sierras Pampeanas and in the Eastern Cordillera and Puna Plateau. In northwestern Sierras Pampeanas this cycle was characterized by an early subduction-related magmatic arc followed by an Early Cambrian (535-520 Ma) continental collision of an semiautoctone terrane identified as the Pampia block or the Arequipa-Belen-Antofalla Craton. The Famatinian Orogenetic Cycle, documented in Famatina System, Eastern Cordillera and Puna Plateau, took place between 490 and 400 Ma. It records the subduction of oceanic lithosphere during the Early Paleozoic, followed by the an exotic terrane accretion, the so-called Precordillera terrane at ca. 460 Ma.

In NW Argentina the Pampean Orogeny is documented in Eastern Cordillera and eastern border of Puna Plateau as a passive margin turbidite sequence, named Puncoviscana Formation. Plutonic rocks intrude into the folded sedimentary sequence, with age ca. 530 Ma, marking the end of this cycle. The Famatinian Orogeny, in NW Argentina is represented by platformal siliciclastic, shallow-marine sedimentary succession in the Eastern Cordillera and by deeper marine sediments to the west, in the Puna Plateau. In the Puna Plateau, during this cycle, two magmatic arcs so-called “*Faja Eruptiva de la Puna Occidental*” and “*Faja Eruptiva de la Puna Oriental*” are generated.

Volcanic and subvolcanic rocks of uncertain age intercalated with detrital sediments of the Puncoviscana Formation have been recognized in several localities of the Eastern Cordillera and easternmost Puna Plateau: in the Río Grande Valley, in the El Niño Muerto Hill and in the Río Blanco Valley. The present study presents petrographic, geochemical and geochronological (U-Pb and K-Ar ages and Sr-Nd isotopic ratios) data for these volcanic/subvolcanic rocks. This study attempts to characterizations these rocks, identify their original tectonic setting and link them with the western margin evolution of South America.

In the Río Grande Valley, the Neoproterozoic to lower Paleozoic basement rock units are intruded by dikes of basaltic compositions. Petrochemical, mineral chemistry, geochemical and isotopic data for these dykes permitted to classify them as ocellar-analcime monchiquite, a feldspar-free variety of alkaline lamprophyre. Biotite K-Ar data indicate the age of 164 ± 8 Ma. They have low to medium $^{87}\text{Sr}/^{86}\text{Sr}$ initial ratios (0.70377-0.70781), and Nd isotopic compositions corresponding to positive $\epsilon_{\text{Nd}}(t)$ values (+1.5-+5.6) and T_{DM} model ages between 0.25-0.64 Ga. Taking in account the new data, we suggest that the Río Grande

lamprophyres are associated with pre-rift stage of Upper Jurassic to Lower Eocene Salta Intracontinental Rift. The isotopic date and petrochemical characteristics suggest together with the mantle compositions bibliography for this sector, that they derive from an old metasomatized lithospheric mantle, which has probably been chemically transformed during the Ordovician Famatinian Orogeny.

The El Niño Muerto area, localized in the easternmost part of Puna Plateau, consists of a NE-SW hill formed entirely by the Puncoviscana metasedimentary sequence. High-K calc-alkaline rocks of dacitic compositions constitute a sill which intruded into the sequence. New LA-ICPMC U-Pb zircon data and Sr-Nd isotopic ratios indicate that the metadacites crystallized at 495 ± 4 Ma with high initial $^{87}\text{Sr}/^{86}\text{Sr}$ (0.71107-0.71180) and negative $\epsilon_{\text{Nd}}(t)$ values, between -9.7 to -5.9. Their peraluminous character suggests that the original magma was strongly contaminated with continental crust or that it derived from the partial melting of a crustal material with an important sedimentary component. The geochemical and isotopic data presented in this study indicate that the El Niño Muerto metadacites formed in the Famatinian calc-alkaline magmatic arc developed in Ordovician in NW Argentina.

In the Río Blanco Valley, in the eastern Cordillera, have been recognized subvolcanic rocks, pillow lavas and peperite deposits interbedded with metasedimentary rock units which are considered part of the Puncoviscana Formation. The geochemical data allows us to classification the Río Blanco rocks as basalts with geochemical characteristics similar to E-MORBs, with Zr/Nb ratios from 6.8 to 9.5. New LA-ICPMC U-Pb zircon data and Sr-Nd isotopic ratios indicate that the basalts crystallized at 501 ± 9 Ma from a primitive magma with positive $\epsilon_{\text{Nd}}(t)$ values. The Río Blanco basalts have moderately high $^{87}\text{Sr}/^{86}\text{Sr}$ initial ratios (0.70713-0.70942), positive $\epsilon_{\text{Nd}}(t)$ values (+2.47 to +4.46) and T_{DM} model ages between 0.84 to 1.12 Ga. The tectonic setting of formation of the Río Blanco basalts remains unclear and their geochemical characteristics indicate that they correspond either to an E-MORB or to within plate anorogenic basalts.

Taking in account that the Río Blanco basalts and the El Niño Muerto Metadacites had been generated in the same moment, we suggest that the Río Blanco basalts would be interpreted as a remnant of a back-arc basin, in relation to the Famatinian Arc developed more to west during the Late Cambrian /Early Ordovician in NW Argentina.

CAPÍTULO I

1.1. Introdução

O embasamento do noroeste da Argentina, entre 22° e 33° S, registra ao menos duas orogenias durante tempos paleozóicos, as Orogenias Pampeana e Famatiniana (Aceñolaza and Toselli 1973).

O ciclo orogenético Pampeano, aproximadamente contemporâneo com o episódio Panafricano-Brasiliano (Miller and Sollner, 2005), ocorre entre 550 e 515 Ma (Sollner *et al.*, 2000b). Ele é documentado nas províncias morfoestruturais da Cordillera Oriental (CO), Puna e Serras Pampeanas Orientais (Aceñolaza *et al.*, 1990). Esta orogenia compreende etapas de subducção (arco magmático) e colisão de um terreno semiautoctone ao final do Cambriano, o terreno de Pampia (Rapela *et al.*, 2001 a, b) ou o cráton de Arequipa-Belén-Antofalla (Ramos 1988). Depois da acreção ortogonal do terreno, a protomargem de Gondwana se tornou uma margem passiva até os 499 Ma, quando se inicia o Ciclo Famatiniano.

O ciclo Famatiniano ocorre entre 499 e 400 Ma e é um ciclo completo que compreende dois arcos, um externo e o outro interno associados às etapas de subducção e colisão de um terreno exótico a Gondwana durante tempos silurianos, o terreno Precordillera.

Em contraposição à teoria que postula duas orogêneses (Aceñolaza *et al.*, 1990; Rapela *et al.*, 2001 a, b; Ramos 1988), outros autores (Lucassen *et al.*, 2000, Lucassen *et al.*, 2006) acreditam na existência de uma única orogênese desenvolvida entre 600 e 400 Ma e propõem uma evolução em ambiente de cinturão móvel para o embasamento dos Andes Centrais (Lucassen *et al.*, 2000).

O Ciclo Pampeano está representado entre 22° e 27°S por depósitos turbidíticos da Formação Puncoviscana, de idade proterozóica superior - cambriana inferior, e por granitóides cambrianos alojados durante esta orogenia (Bachmann *et al.*, 1987, Omarini *et al.*, 1999).

A Formação Puncoviscana aflora na província morfoestructural da Cordilheira Oriental (Turner 1960) e no limite entre esta e a província da Puna. É constituída por seqüências clásticas turbidíticas (pelitos e grauvacas) e inclui conglomerados e calcários subordinados (Omarini y Baldis, 1984, Jezek, 1990). A complexidade tectônica (Mon y Hongn, 1991) limita as avaliações estratigráficas e torna difíceis as correlações entre os distintos afloramentos.

O Ciclo Famatiniano é caracterizado no noroeste da Argentina por rochas sedimentares e magmáticas. Na Cordilheira Oriental as rochas sedimentares são seqüências marinhas pouco

profundas enquanto na Puna predominam as seqüências marinhas profundas. O magmatismo famatiniano desenvolvido na Puna compreende rochas básicas e ácidas associadas a importantes depósitos turbidíticos vulcanoclásticos. Durante este ciclo se desenvolveram dois arcos magmáticos, denominados pela posição geográfica “Faixa Eruptiva da Puna Oriental” (Mendez *et al.*, 1973) e “Faixa Eruptiva da Puna Ocidental” (Palma *et al.*, 1986).

Rochas vulcânicas de natureza calci-alcalina e alcalina (Toselli e Aceñolaza, 1984, Coira *et al.*, 1990; Omarini *et al.*, 1999) de idades incertas se intercalam na Formação Puncoviscana. Os afloramentos de rochas vulcânicas estão expostos em três setores: um setor norte, o Vale de Río Grande; um setor central, a Serra de El Niño Muerto e um setor sul, o Vale de Río Blanco (Fig. 1).

Estas rochas vulcânicas não foram estudadas em detalhe e pouco se sabe a respeito de sua relação com a Formação Puncoviscana e a margem de Gondwana. Os estudos propostos neste trabalho oferecem a possibilidade de definir a idade e o ambiente de alojamento das mesmas e a sua relação com a evolução da Plataforma Sulamericana.

1.2. Objetivos

Objetivos gerais

Os objetivos gerais desta dissertação são gerar novos dados petrográficos, geoquímicos e isotópicos das rochas vulcânicas que se intercalam na seqüência sedimentar da Formação Puncoviscana, no noroeste da Argentina. O estudo detalhado dessas rochas aportará melhor conhecimento dos ambientes de alojamento e permitirá vinculá-las com os orógenos da Plataforma Sul-americana.

Objetivos específicos

Os objetivos específicos desta dissertação são:

- a) caracterizar petrograficamente as litologias vulcânicas com a finalidade de classificá-las;
- b) caracterizar geoquimicamente as rochas vulcânicas mediante análises de elementos maiores, traço e Terras Raras com a finalidade de determinar o ambiente tectônico e os possíveis processos magmáticos envolvidos na sua geração e evolução;
- c) determinar idades através de análises U-Pb em zircões por *laser ablation inductively coupled plasma mass spectrometry*, (LA-ICPMS) e K-Ar em biotitas, além de estudos de isótopos de Sr-Nd. Estes dados em conjunto têm como finalidade determinar a fonte magmática das rochas vulcânicas e associá-las a um quadro tectônico apropriado.

1.3. Localização e vias de acesso

A área de trabalho está inserida nas unidades morfoestruturais da Cordilheira Oriental e no limite desta com o platô da Puna e coincide com afloramentos da Formação Puncoviscana. As áreas estudadas podem ser divididas geograficamente em três setores.

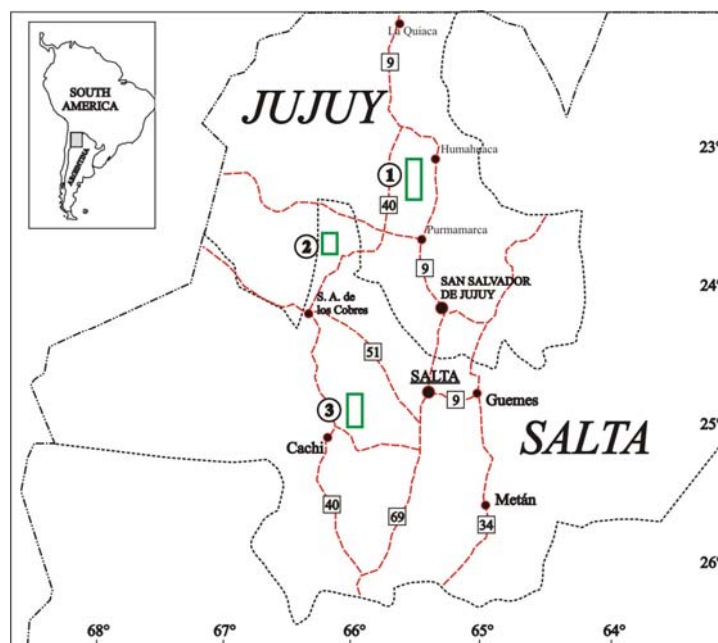


Figura 1: Mapa de localização e vias de acesso às áreas estudadas. 1, Vale de Río Grande. 2, Serra de El Niño Muerto. 3, Vale de Río Blanco.

1) Setor norte: está localizado no departamento de Humahuaca, na província de Jujuy. A área de estudo compreende a localidade de Coraya e as Quebradas de Yacoraite e Huichaira, as quais são delimitadas pelos paralelos $23^{\circ} 35' 00.01''$ e $23^{\circ} 09' 13.00''$ S e meridianos $65^{\circ} 27' 27.00''$ e $65^{\circ} 22' 00.01''$ W (Fig. 1). Essa área se encontra 150 km a norte da cidade de Salta (capital da província de Salta) e 30 km a norte da cidade de San Salvador de Jujuy (capital da província de Jujuy), sendo a principal via de acesso para a área estudada a Rodovia Nacional N° 9. Neste setor foram reconhecidos diferentes diques alcalinos que intrudem a Formação Puncoviscana.

Tanto Coraya quanto as duas quebradas fazem parte de vales secundários relacionados ao Vale de Río Grande ou Quebrada de Humahuaca. Este vale foi um importante centro de assentamento de povoados pré-históricos.

2) Setor Central: o acampamento de El Niño Muerto está no setor sul do Salar de Salinas Grandes, no limite entre as províncias morfoestruturais da Cordilheira Oriental e da Puna. Localiza-se no departamento de Susques, na província de Jujuy e é limitado pelos

paralelos 23° 46' 34.73" e 23° 44' 07.68" S e meridianos 66° 14' 50.52" e 66° 12' 48.88" W (Fig. 1). A principal via de acesso é a Rodovia Nacional N° 51, que parte da cidade de Salta e leva até a localidade de San Antonio de los Cobres, de onde se segue por 30 Km para norte por estradas não asfaltadas até o acampamento de El Niño Muerto. Neste setor foi identificado um sill de composição dacítica que instrude turbiditos finos da Formação Puncoviscana.

3) Setor sul: no setor sul foram agrupadas as localidades de Las Cuevas e Vale do Río Blanco.

O Vale de Río Blanco se localiza no setor leste dos Vales Calchaquíes, no departamento de Cachi (província de Salta), a aproximadamente 160 km a oeste da cidade de Salta-capital. A área de estudo é limitada pelos paralelos 24° 58' 08.73" e 24° 57' 05.43" S e meridianos 65° 58' 03.41" e 65° 57' 35.71" W (Fig. 1). Partindo da cidade de Salta pela Rodovia Nacional N° 78 se chega até a localidade de Guachipas, onde se encontra a Rodovia Nacional N° 40, principal via de acesso aos Vales Calchaquíes para leste. Após a Cuesta del Obispo, na entrada dos povoados de Payogasta e Cachi, se encontra uma rodovia não asfaltada que, para leste, conduz à Chácara de Colomé. Segue-se aproximadamente 15 km por esta estrada até a entrada do Vale do Río Blanco. Neste setor foram reconhecidas rochas vulcânicas submarinas contemporâneas com a deposição de seqüências hemipelágicas e turbiditos, reconhecidos como parte da Formação Puncoviscana.

1.4. Marco geológico regional do noroeste Argentino

Os Andes Centrais entre 22° e 26° S (Fig. 2) estão caracterizados por diferentes províncias morfoestruturais que de oeste para leste são: Cordilheira de la Costa, Precordilheira Chilena, Valle Longitudinal, Depressão Preandina, Cordilheira Occidental, Altiplano/Puna, Cordilheira Oriental, Serras Subandinas e Planície Chaqueña.

Tanto os afloramentos da Formação Puncoviscana quanto as rochas vulcânicas estudadas estão expostas na Cordilheira Oriental e no limite entre esta e o Altiplano-Puna, portanto estas duas províncias morfoestruturais serão brevemente descritas a seguir:

Altiplano-Puna

O Platô Altiplano-Puna foi chamado inicialmente Puna de Atacama por Brachkebusch (1883). Os investigadores que o seguiram a chamaram simplesmente de Puna (Bonarelli, 1913, 1915; Keidel, 1927). É caracterizado por altitudes acima de 3.700 m.s.n.m. Apresenta limites bem definidos com a Cordilheira Oriental (Fig. 2) a leste e com a Cordilheira

Occidental a oeste, já em território chileno. Para sul é limitado pela Cordilheira Frontal e Sistema de Famatina. Suas características foram bem definidas por Turner em 1970.

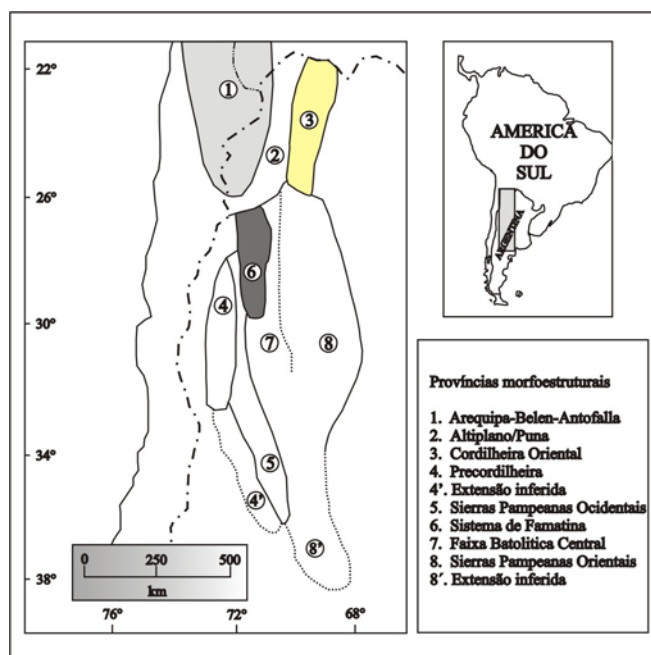


Figura 2: localização e distribuição das províncias morfoestruturais no centro-norte da Argentina. (modificado de Pankhurst and Rapela, 1998)

Sua estrutura está caracterizada por grandes falhas de empurrão de idade paleógena a miocênica que estruturaram esta altiplanície. Alguns autores (Alonso *et al.*, 1984) propuseram dividir a Puna em dois setores com base em suas características geológicas regionais: Puna Setentrional ou Jujeña, e a Puna Austral ou Saltocatarqueña. O limite entre essas províncias corresponde ao lineamento Calama-Olacapato-El Toro (COT) (Salfity *et al.*, 1975, 1984).

A Puna setentrional ou Jujeña é caracterizada pela presença de rochas ordovicianas enquanto a Puna Austral é caracterizada por afloramentos de rochas de baixo grau metamórfico pré-cambrianas, uma faixa eruptiva ordoviciana com granodioritos, granitos e extenso magmatismo basáltico pleistoceno associado a um regime extensional.

As rochas mais antigas que afloram são sedimentos e rochas ordovicianas de baixo grau metamórfico sem base exposta, as quais no setor ocidental se encontram fortemente deformadas. Elas se associam a dois arcos magmáticos, a Faixa Eruptiva da Puna Oriental (Mendez *et al.*, 1973) e a Faixa Eruptiva da Puna Ocidental. (Palma *et al.*, 1986). No limite com o Chile, rochas marinhas silurianas e devonianas encontram-se sobre os sedimentos ordovicianos, que são sobrepostos por depósitos continentais de idade neopaleózoica.

Os depósitos mesozóicos a eoterciários correspondem à bacia do rift do Grupo Salta, que no setor norte da Puna apresentam um importante depocentro de sedimentação. Os depósitos que preenchem o rift são em sua maioria continentais, porém há uma breve transgressão marinha que corresponde à Formação Yacoraite, de idade campaniana-maestrichtiana (Salfity y Marquillas, 1981, 1994).

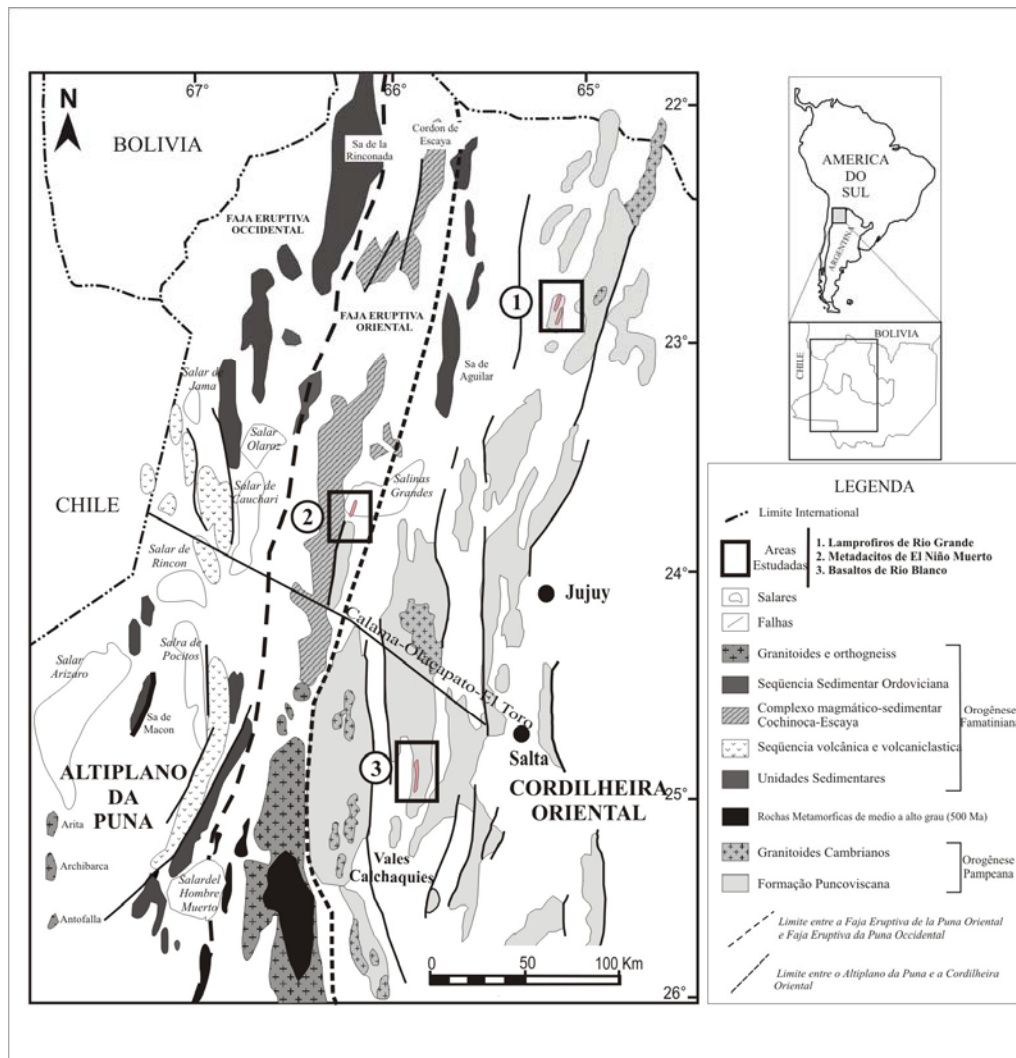


Figure 3: afloramentos do embasamento Neoproterozóico - Paleozóico-Inferior do noroeste da Argentina (modificado de Do Campo and Guevara, 2005 and Viramonte et al., 2007)

O Platô Altiplano-Puna apresenta como característica principal o vulcanismo orogénico cenozóico, representado por estrato vulcões de composição andesítica-dacítica, caldeiras vulcânicas associadas a fluxos ignimbríticos e cones basálticos monogénéticos. Eles se intercalam com depósitos terciários continentais desenvolvidos em bacias endorréicas que culminam com depósitos evaporíticos. Estes últimos constituem os grandes salares da Puna, outra características marcante.

Cordilheira Oriental

A Cordilheira Oriental foi definida por Brackebusch (1892) para o território argentino, entretanto Keidel (1925) e Groeber (1938) restringiram o uso do termo para as províncias de Salta e Jujuy. É limitada a oeste pelo Platô da Puna, que corresponde ao limite da antiga plataforma eo-paleozoica. Para sul é limitada pelas Serras Pampeanas Orientais (Fig. 2). Para leste o limite com as Serras Subandinas é tectônico, representado pela falha de empurrão Andina Principal (*main Andean Thrust*, Roeder, 1988). Esta falha foi responsável pelo soerguimento do embasamento neoproterozóico e das seqüências cambro-ordovicianas

O soerguimento da Cordilheira Oriental ocorreu durante o Mioceno superior com importantes movimentos neotectônicos que prosseguem até os dias atuais. Sua estrutura está caracterizada por imbricações de escamas de direção norte-noroeste que envolvem o embasamento granítico e sedimentar. As escamas têm uma estrutura complexa, resultado das orogenias Pampeana (limite Neoproterozóico-Cambriano), Oclóyica (Ordoviciano médio-superior) e Chánica (final do Devoniano).

A característica principal da Cordilheira Oriental é o desenvolvimento de grandes lascas de empurrão descritas por Keidel (1943). Estas lascas são compostas por um bloco neoproterozóico caracterizado por sedimentos pouco metamorfisados da Formação Puncoviscana. Rochas plutônicas do Cambriano, como os Granitos de Tastil, Cañani, Nevado de Cachi e Nevado de Palermo alojam-se nesta unidade e sedimentos clásticos de ambiente de plataforma dos Grupos Mesón (Cambriano) e Santa Victoria (Ordoviciano Inferior a Médio) depositam-se em discordância angular sobre ela (Fig. 3). Depósitos marinhos silurianos e devonianos estão em suave discordância angular sobre os depósitos anteriores. Essa discordância é resultante dos movimentos da fase Oclóyica, que ocasionou a ascensão da protoPuna (Bonarelli, 1915) ou, alternativamente, da Protocordilheira Oriental (Ramos y Palma, 1996).

As diferentes unidades da bacia do rift do Grupo Salta, desenvolvido entre o Cretáceo e o Eoceno, se depositaram em discordância sobre os depósitos paleozóicos e mais antigos. Escasso vulcanismo cenozóico se aloja no Abra de Pives e Diego de Almagro.

No extremo austral da Cordilheira Oriental (Turner 1970) está a subprovíncia de Cumbres Calchaquíes (Baldis *et al.*, 1976), que compreende o extremo norte das Serras Pampeanas Norocidentales (Caminos, 1979). Esta não apresenta variações estratigráficas tão marcantes para separá-la da Cordilheira Oriental. A subprovíncia de Cumbres Calchaquíes é formada por um embasamento neoproterozoico de baixo grau metamórfico, a Formação Puncoviscana, e seu equivalente de maior grau metamórfico, a Formação La Paya. O embasamento também está constituído por granitóides eopaleozoicos como a Formação Cachi

e ordovicianos como os Granitos Angostura, Pucará, Complexo Eruptivo Oire. Sobre estes se depositam os sedimentos continentais do Subgrupo Pirgua (Cretáceo superior) e os depósitos continentais com registro de transgressões marinhas, os Subgrupos Balbuena e Santa Bárbara (Paleógeno). Os depósitos do Grupo Salta são característicos da região oriental. Sobre eles estão os depósitos da Formação Luracatao e do Grupo Payogastilla. Durante este período (Neógeno) se registrou um episódio vulcânico (Ignimbritas Pucarilla) que representaria uma conexão com o vulcanismo da Puna.

1.5. Metodologia geral do trabalho

O trabalho de pesquisa foi dividido em 8 etapas:

(i) *Etapas de campo*: foi dividida em três partes e consiste no reconhecimento e amostragem das rochas vulcânicas, objeto deste estudo nas diferentes localidades. Posteriormente as amostras foram enviadas para o Brasil, para continuar com as etapas seguintes.

(ii) *Estudo da Petrografia das rochas vulcânicas*: as lâminas correspondentes à primeira campanha foram realizadas no *Laboratório de Petrotomia de la Universidad Nacional de Salta-Argentina*. As lâminas das amostras recolhidas nas duas etapas subseqüentes foram confeccionadas no Laboratório de Laminação do Instituto de Geociências da Universidade de Brasília-UnB. Foram descritas 110 lâminas em microscópios petrográficos de luz transmitida do Laboratório de Microscopia do Instituto de Geociências da UnB. A partir do estudo petrográfico, foram selecionadas amostras para as análises geoquímicas realizadas no Laboratório Acme, no Canadá.

(iii) *Análises químicas de minerais - Microsonda Eletrônica (ANEXO II)*: a composição química do de olivina, clinopiroxênio, anfibólio, mica, analcima, apatita e carbonato foi obtida no Laboratório de Microsonda Eletrônica do Instituto de Geociências da UnB. O modelo da microsonda utilizada é CAMECABAX SX 50. As condições analíticas foram 10 nÅ de transmissão de corrente e 15 KeV de transmissão de energia. O tempo de aquisição foi de 10 segundos e o diâmetro do feixe foi de 1 µm.

(iv) *Preparação das amostras para análises químicas de rochas e análises isotópicas Sr-Nd*: Superfícies de alteração e veios de quartzo e carbonato foram cuidadosamente retirados das amostras. Posteriormente as rochas não alteradas foram pulverizadas a uma granulometria

inferior a 200 mesh utilizando-se um moinho de bola de carbetto de tungstênio no Laboratório de preparação de amostras do Laboratório de Geocronologia do Instituto de Geociências - UnB.

(v) Separação de minerais: a separação dos minerais pesados (zircão) foi realizada no Laboratório de Concentração de amostras do Laboratório de Geocronologia - UnB. Utiliza-se como método de pré-concentração o método da batéia e posteriormente o separador magnético Frantz. A última etapa de concentração foi realizada por seleção no microscópio binocular.

(vi) Montagem dos grãos de zircão: foram selecionados aproximadamente 30 zircões euédricos e limpos para cada amostra. Posteriormente cada grupo foi montado sobre uma lâmina de vidro e colocado sobre um molde cilíndrico onde se colocou duas partes de resina (Struer, Epo Fix Resin) e uma parte de endurecedor (Struer, Epo Fix Hardener). Cada amostra foi deixada em repouso aproximadamente 12 horas para permitir que a resina e o endurecedor se solidificassem. Em seguida, retira-se o molde cilíndrico e obtém-se os “mount” de zircões. Cada mount foi polido com pasta diamantada e lavado com HNO₃ 3% junto com zircões padrões durante 5 minutos para remover da superfície possíveis contaminações com Pb. Este último passo foi realizado imediatamente antes de serem efetuadas as análises no LA ICP-MS.

(vii) Análises químicas de rochas: as análises químicas de elementos maiores, traços e terras raras foram realizadas em 24 amostras. A preparação das amostras para análises dos elementos maiores e parte dos traços foi feita por fusão com LiBO₂ e para a obtenção dos metais base e preciosos por digestão com água régia. As análises dos elementos maiores, Ni e Sc (ppm) foram efetuadas por espectrometria de emissão com ICP-OES e dos elementos traços por espectrometria de massa com ICP-MS.

(viii) Difratometria de raios-x (ANEXO I): Foram realizadas duas análises por difração de Raios-X. As análises foram feitas no Laboratório de Difração de Raios-X do IG-UnB. O difratômetro usado é um Rigaku-D/MAX-2 A/C com fonte de cobre e filtro de Ni, operando com 40 Kv e 20 mA. A interpretação dos difratogramas foi feita pela professora Edi Mendes Guimarães, usando o Software JADE 3-0 da MDI, com banco de dados PC-PDF (Power Diffraction File-PDF para PC) produzido pelo International Center for Diffraction.

(ix) Geoquímica isotópica Sr-Nd: Análises de Sm-Nd e Rb-Sr foram realizadas em amostras de rocha total (25 exemplares e 11 exemplares, respectivamente) no Laboratório de Geocronologia da UnB. As amostras selecionadas foram: dacitos, lamprófiros e basaltos. As análises Sm-Nd foram feitas segundo o método descrito por Gioia & Pimentel (2000). Entre 70 e 100mg de pó de amostra foram misturados com uma solução traçadora mista (spike) de ^{149}Sm - ^{150}Nd e dissolvidos em cápsulas Savillex. A extração dos elementos lantanídeos foi feita através de técnicas convencionais de troca iônica em colunas de quartzo, usando resina BIO-RAD AG-50W-X8. As extrações de Sm e Nd foram realizadas em colunas de teflon empacotadas com resina LN-Spec (resina líquida HDEHP-ácido di-ethylhexil fosfórico impregnada em pó de teflon). As frações de Sr, Sm e Nd foram depositadas em arranjos duplos de filamentos de rênio. As razões isotópicas foram determinadas em espectrômetro de massa multi-coletor Finnigan MAT 262 em modo estático. As incertezas para $^{87}\text{Sr}/^{86}\text{Sr}$ são menores do que 0.01% (2σ) e para as razões $^{147}\text{Sm}/^{144}\text{Nd}$ e $^{143}\text{Nd}/^{144}\text{Nd}$ são melhores do que $\pm 0.2\%$ (1σ) e $\pm 0.005\%$ (1σ) respectivamente, baseadas em análises repetidas dos padrões BHVO-1 e BCR-1. As razões $^{143}\text{Nd}/^{144}\text{Nd}$ são normalizadas para o valor de $^{146}\text{Nd}/^{144}\text{Nd}$ de 0.7219 e a constante de desintegração (λ) usada é $6.54 \times 10^{-12} \text{ a}^{-1}$. Os resultados foram processados utilizando-se o programa ISOPLOT/Ex 3 (Ludwig 2007). Os valores de T_{DM} foram calculados utilizando o modelo DePaolo (1981).

(viii) Geocronologia U-Pb, K-Ar: as determinações isotópicas U-Pb sobre zircões foram realizadas no Laboratório de Geocronologia da Universidade de Brasília-UnB. O equipamento utilizado foi o Laser Ablation (LA)-ICP MS multi-colector. O Laser Ablation usa um laser de corrente concentrada para vaporizar pequenas quantidades de amostra de zircão contidas numa cela fechada. O material evaporado é transportado em gás Hélio (gás transportador), desde a cela até o ICP-MS para a quantificação isotópica. O instrumento tem incorporado um microscópio que permite obter uma imagem de alta resolução, o que facilita a localização dos diferentes grãos de zircão da amostra.

As condições analíticas foram as seguintes:

Padrão usado: **GJ 1**

Método de varredura: linear (2x)/ pré-ablation

Diâmetro do laser: 30 μm .

Energia (%): 32%

Energia J/cm^3): 0.15

Frequência: 12 Hz.

Tempo de aquisição: 10 segundos

Fluxo de Argônio: 0.903

Fluxo de Helio: 0.4

(ix) Confeção da dissertação de mestrado e dos artigos: esta etapa consistiu na redação do texto da dissertação, elaboração dos artigos, preparação, apresentação de defesa de dissertação.

1.6. Escopo da dissertação

Os resultados obtidos nesta dissertação são referentes às rochas vulcânicas do embasamento sedimentar do noroeste argentino. Os dados são apresentados em forma de dois artigos.

No *Capítulo 2*, são apresentados os dados obtidos no Vale de Río Grande. A razão da escolha foi a idade jurássica obtida para os diques da região o que permitiria ligá-los à abertura do Rift de Salta, desenvolvido durante o Cretáceo no noroeste da Argentina. O título do artigo é: **“K-Ar AGE AND Sr-Nd ISOTOPIC GEOCHEMISTRY OF ALKALINE LAMPROPHYRE DYKES, FROM THE RIO GRANDE VALLEY, EASTERN CORDILLERA, NW ARGENTINA: ORIGIN OF PARENTAL MAGMAS FROM A METASOMATIZED MANTLE.”**

O artigo se inicia com uma introdução e contexto geológico sobre o Jurássico nos Andes Centrais e sobre o desenvolvimento do rift Cretáceo tanto na Argentina como no sul da Bolívia. Seguem, nos capítulos de geologia de campo e Petrografia, a descrição de campo, tipos de contato com o embasamento e caracterização petrográfica das rochas. A caracterização química dos minerais destas rochas é apresentada no capítulo química mineral. Posteriormente há uma breve descrição dos métodos analíticos usados para a obtenção dos dados de elementos maiores, traços e terras raras e para os dados isotópicos Sr-Nd. Nos capítulos de geoquímica e geocronologia as rochas são caracterizadas e classificadas. Finalmente os dados são tratados no capítulo de Discussões

No *Capítulo 3*, são apresentados os dados obtidos na Serra de El Niño Muerto e no Vale do Río Blanco. Os motivos da escolha foram as idades semelhantes obtidas para estes dois grupos de rochas que permitiram relacioná-los com a margem de Gondwana durante o Paleozóico Inferior.

O título do artigo é: **“LA-ICPMS U-PB GEOCHRONOLOGY AND PETROLOGY OF VOLCANIC ROCKS OF THE NEOPROTEROZOIC**

TO LOWER PALEOZOIC BASEMENT OF THE CENTRAL ANDES, NW ARGENTINA: INSIGHTS ON THE EVOLUTION OF WESTERN GONDWANA.”

O artigo começa com uma introdução e contexto geológico sobre as orogenias registradas no embasamento do noroeste, desde o Neoproterozoico até o Paleozóico Médio-Superior. Seguem, nos capítulos de Relações de campo e Petrografia, a descrição de campo, tipo de contatos com o embasamento e caracterização petrográfica das rochas. Posteriormente segue uma breve descrição dos métodos analíticos usados para a obtenção dos dados de elementos maiores, traços e terras raras, além de dados isotópicos Sr-Nd e idades U-Pb sobre zircões. Nos capítulos de geoquímica e geocronologia as rochas são caracterizadas e classificadas. Finalmente os dados são tratados no capítulo de Discussões.

CAPÍTULO II

Este Capítulo apresenta os resultados petrológicos, geoquímicos e isotópicos, referentes às rochas subvolcânicas do Vale de Río Grande. As informações serão apresentadas na forma de artigo completo, que será submetido ao periódico *Journal of South American Earth Sciences*, com o título:

“K-Ar AGE AND Sr-Nd ISOTOPIC GEOCHEMISTRY OF ALKALINE LAMPROPHYRE DYKES, FROM THE RIO GRANDE VALLEY, EASTERN CORDILLERA, NW ARGENTINA: ORIGIN OF PARENTAL MAGMAS FROM A METASOMATIZED MANTLE.”

K-Ar AGE AND Sr-Nd ISOTOPIC GEOCHEMISTRY OF ALKALINE LAMPROPHYRE DYKES, FROM THE RIO GRANDE VALLEY, EASTERN CORDILLERA, NW ARGENTINA: ORIGIN OF PARENTAL MAGMAS FROM A METASOMATIZED MANTLE

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Abstract

In the Río Grande Valley, NW Argentina, several dykes intrude the Neoproterozoic to lower Paleozoic basement. In this study new petrographic, geochemical and isotopic data allowed to classify these rocks as ocellar-analcime monchiquite, a feldspar-free variety of alkaline lamprophyre composed mainly of Ti-rich-diopside or augite, Ti-rich biotite/phlogopite, forsteritic olivine, titanian-pargasite and analcime with abundant ocelli filled with analcime and carbonate. High LOI, Cr, Ni, Sc, Co and V and variable contents in Ba, Sr, Rb and Nb are observed. They show distinct enrichment in LILE and LREE compared to HFSE and HREE, and the lack of Eu anomaly indicates the absence of plagioclase fractionation. Trace element contents suggest that they were formed in a within plate setting. K-Ar analyses of biotite indicates an age of 164 ± 8 Ma. They present variable $^{87}\text{Sr}/^{86}\text{Sr}$ initial ratios, ranging from 0.70377 to 0.70781, and initial $^{143}\text{Nd}/^{144}\text{Nd}$ ratio between 0.512506 and 0.512716), corresponding to positive $\epsilon_{\text{Nd}}(t)$ values (+1.5 to +5.6) and T_{DM} model ages between 0.25-0.64 Ga. Taking into account their age, petrography, geochemical and isotopic data, we suggest that the Río Grande lamprophyres originated during the Jurassic by partial melting of old metasomatized lithospheric mantle, during the pre-rift stage in NW Argentina.

1. Introduction

Lamprophyres comprise an important group of rocks since they are considered to represent primary mantle-derived magmas generated at considerable depths (100-150 km) (*e.g.* Rock, 1991, Bernard-Griffiths et al., 1991, Bédard, 1994). Generally they carry mantle and crustal xenoliths, yielding crucial information on the mantle source and also on geodynamic processes involved in the origin of these magmas. Similarly to other alkaline rocks, lamprophyres are abundant in continental rifts, failed arms of triple junctions (Brooks & Platt 1975) as well as in some oceanic islands (Almeida 1955, 1961, Mitchell and Thom, 1976).

However they also occur in orogenic zones and their peripheries (Richey 1939, Barrab 1952, Viterbo and Zanettin 1959, Rock 1982, Knill 1982), in island arcs (Yagi et al., 1975, Nixon et al., 1980), in passive to destructive continental margins (Snaveley and Wagner 1961), in anomalous uplifted fragments of oceanic crust (Aumento and Ade-Hall 1973, Aumento et al., 1974, Cornen 1981, 1982,) and near major transcurrent faults (Cooper 1979). Therefore, lamprophyres are associated not only with intra plate settings, but also with divergent, passive and even convergent plate margin settings (Rock, 1987).

In the Río Grande Valley in northwestern Argentina (Eastern Cordillera) subvolcanic alkaline rocks, which are the subject of the present study, are emplaced into Upper Neoproterozoic to Lower Paleozoic basement rock units of uncertain age. Originally these rocks have been classified as ankaramites, piroxenitic hornblendites, picrites and basalts with alkaline affinities (Toselli and Aceñolaza, 1984, Manca et al., 1987 and Chayle et al., 1987) (Fig. 1). Toselli and Aceñolaza (1984) and Omarini et al. (1999) suggested a Neoproterozoic age for these dykes based on field evidence, whereas other authors proposed an uncertain post- Ordovician age (Manca et al., 1987 and Chayle et al., 1987, Coira et al., 1990). Due to the lack of detailed field and petrological data as well as of accurate geochronology, little is known about their tectonic setting of emplacement.

In this paper new field, petrographic, geochemical and isotopic data for the subvolcanic rocks of the Río Grande Valley are discussed. The new data show that these rocks are alkaline lamprophyres and they also allowed us to place constraints on their petrogenesis. We show and discuss, for the first time, their mineral chemistry, geochemical data and a K-Ar radiometric age on biotite. Additionally, a new set of whole rock Sr-Nd isotopic data is presented and helped to better understand the magmatic evolution, nature of the mantle source and tectonic setting of emplacement of these lamprophyres. Finally, we attempt to clarify the geodynamic significance of these dyke swarm and integrate the new results into the geodynamic evolution of NW Argentina.

2. Geological Setting

Since the Early Mesozoic, the geodynamic evolution of the Andean margin was characterized by a complex interaction of the Aluk-Phoenix, Farallon and Pacific oceanic plates with the South American Continental plate (Zonenshayn et al., 1984, Jaillard et al., 1990, 2000, Scheuber et al., 1994). During the Lower Jurassic (Sinemurian-Pleinsbachian times) the subduction of the Aluk (Phoenix) plate developed a large magmatic arc along the western margin of South America. This magmatic activity is represented by the emplacement of subalkaline plutonic and volcanic rocks exposed, for instance, in northern Chile (Mpodoziz

and Ramos, 1990, Romeuf et al., 1995, Jaillard et al., 2000). At that time, the whole region, including southeastern Bolivia and northwestern Argentina, was affected by extensive tectonic and magmatic activity, which produced abundant alkaline-subalkaline plutonic and volcanic rocks (e.g. Galliski and Viramonte, 1988, Rubiolo, 1992, Viramonte et al., 1999, Menegatti, 2001, Cristiani, 2004, Comin-Chiaramonte et al., 2005, Cristiani et al., 2005). To the east, in northwestern Argentina and eastern Bolivia the magmatism was clearly associated with the formation and evolution of a continental rift system (Fig. 1) (Salfity and Marquillas, 1981, Galliski and Viramonte, 1988, Avila Salinas, 1986 and Omarini et al., 1989) in a back-arc position with respect to the Jurassic-Cretaceous volcanic arc. This structure was filled with a sedimentary sequence represented by the Salta Group in Argentina (Moreno, 1970, Reyes and Salfity, 1973, Salfity and Marquillas, 1981, Salfity, 1982, Gomez Omil et al., 1989) and the Puca Group in Bolivia (Russo and Rodrigo, 1965, Reyes and Salfity, 1973, Cherroni, 1977). The Salta Group comprises three units: the Pirgua, Balbuena and Santa Barbara subgroups in which three major magmatic phases have been recognized and are briefly described below (Salfity and Marquillas, 1981, Avila Salinas, 1986, Galliski and Viramonte, 1988, and Omarini et al., 1989).

During the Early Jurassic and Cretaceous the *pre-rift* phase (Fig. 1) was dominated by regional doming produced by the rise of plume-related asthenospheric masses, with associated crustal thinning and an intricate pattern of extensional faults. This stage was characterized by



Figure 1: Map of the Jurassic-Cretaceous Rift of NW Argentina. The location of pre-rift plutonic rocks and the study area are shown. Modified from Salfity and Marquillas, (1981) and Viramonte et al., (1999).

the emplacement of anorogenic igneous complexes along the major extensive tectonic structures striking NNE (e.g. Cobres Salinas Grandes fault). The Agua Del Desierto, Rangel, Tusaquillas, Abra Laite, Aguilar, Fundiciones, Hornillos and Mecoya plutonic complexes characterize this ***first magmatic phase (at ca. 150-100 Ma)*** in NW Argentina, consisting of plutonic rocks which intruded into Late Precambrian/Lower Paleozoic metasedimentary sequences (Lanfranco, 1972, Rubiolo, 1992, Menegatti, 2001, Cristiani, 2004).

The *syn-rift phase* was marked by thermal subsidence, with continental red beds infilling the intracontinental rift, with minor marine transgressions registered in the Pirgua Subgroup. The Balbuena Subgroup is the main representative of the syn-rift sequence and is the most characteristic lithostratigraphic unit of the Cretaceous-Paleocene basins in central-western South America. During this stage the two other magmatic phases took place. ***The second magmatic phase (80-75 Ma)*** is marked by alkaline volcanism (basanites and mugearites) developed along the axis of the main tectonic structures. The ***third and late igneous phase (65-60 Ma)*** comprises lamproitic sills intruded into the Blanquitos Formation of the Pirgua Subgroup (Omarini et al., 1989), basanites interbedded in the Balbuena and Santa Barbara Subgroups in Argentina as well as basanites, phonotephrites (Vargas Gil, 1965, Galliski and Viramonte, 1988) and lava flows (Castaño and Rodrigo, 1978) associated with equivalent sedimentary successions in Bolivia. This late igneous phase marks the onset of the thermal subsidence and collapse of the basin during the *post-rift stage*, just before the closure of the Cretaceous-Paleocene rift basin due to the dramatic change in the geodynamic style of the region caused by the Incaic compressive phase during the Upper Paleocene.

The origin of this important extensional geological structure, which played a major role in the Central Andean Cordillera evolution, still remains controversial. Galliski and Viramonte (1988) related this magmatism to the back-arc zone in relation to the Jurassic subduction of the Aluk plate under South America, whereas Cristiani et al. (2000) associated the Upper Jurassic-Cretaceous magmatism in NW Argentina with the break-up of Gondwana. We are in agreement with the second hypothesis.

The pre-rift magmatic phase was characterized by the emplacement of intrusive rocks located in central and southeastern Bolivia (Velasco and Ayopaya region) and in NW Argentina (Puna Plateau and Eastern Cordillera). Close to the Argentina-Bolivia border alkaline and ultramafic lamprophyres of uncertain affinity have been described by Rubiolo et al. (1997). These authors recognized aillikite dykes with K-Ar age of 224 ± 08 Ma intruding into the Puncoviscana Formation. Towards the south, Rubiolo et al. (1997) identified sannaites associated with the cretaceous alkaline complex of Hornillos (59 ± 04 and 71 ± 09 K-Ar age) intruding into the Neoproterozoic-Lower Paleozoic basement rock units. They have

been interpreted as representative of igneous activity associated with extensional processes related with the formation of the intracontinental rift (Galliski and Viramonte, 1988, Rubiolo, 1992, 1997, Rubiolo et al., 1996). Two lamprophyre dykes of this complex were dated with the K-Ar method at 71 ± 9 and 59 ± 4 Ma (Rubiolo et al., 2003). Approximately 100 km east of the Río Grande Valley, in Sierras Subandinas, in the Los Alisos locality, ultramafic dykes dated at 303 ± 10 Ma K-Ar age have been recognized (Mendez and Villar, 1979, Barbieri et al., 1997), apparently not related to the Salta rift development.

In the Río Grande Valley several alkaline dykes, object of the present study, outcrop in the Coraya, Yacoraite and Huichaira secondary valleys (Manca et al., 1987 and Chayle et al., 1987, Coira et al., 1990) between $23^{\circ} 35' 00.01''$ and $23^{\circ} 09' 13.00''$ S from $65^{\circ} 27' 27.00''$ to $65^{\circ} 22' 00.01''$ W (Fig. 1). These dykes intrude into the Neoproterozoic-Lower Paleozoic basement of the Eastern Cordillera, formed by the Puncoviscana Formation, together with the Meson and Santa Victoria Groups.

2.1. Local Geology

The study area is characterized by Neoproterozoic to Lower Paleozoic sedimentary basement rock units. The lower unit is the Puncoviscana Formation which, according to available geochronological (e.g. Do Campo, 1999b, Do Campo et al., 1999), and paleontological data (Buatois and Mangano, 2003), has been interpreted as Vendian to Nemakitian-Daldynian age. It is represented by a turbidite sequence consisting of centimetric to decimetric beds of grey and grayish-red to green greywacke and pelite which have been intensely folded. Their tectonic settings of deposition remain a subject of debate. Models vary from passive margin (Ramos, 1988, Jezek, 1990, Sims et al., 1998, Rapela et al., 1998) and foreland basin settings (Kraemer et al., 1995, Keppie and Bahlburg, 1999) to intracontinental rift which might have evolved into an active plate margin (Omarini et al., 1999a, b). This unit was affected by polyphase deformation during the evolution of western South America, however the main deformational event took place during the Tilcaric Orogeny in late Neoproterozoic to early Cambrian times (Ramos and Basei, 1997, Do Campo, 1999b). The sequence is unconformably overlain by pink sandstones of the Middle to Upper Cambrian Meson Group which is interpreted as a shallow-marine passive margin system (Aceñolaza et al., 1982, Kumpa and Sanchez, 1988). This is overlain by the greywacke and pelites of Santa Victoria Group (Turner, 1960, Salfity et al., 1984) which represent the Famatinian Cycle in the Eastern Cordillera.

Table 1: Summarized lithological and petrographic characteristics of the Río Grande sills and dykes.

<i>Locality</i>	<i>Sample N°</i>	<i>Lithology</i>	<i>Host rock</i>	<i>Texture</i>	<i>Phenocrysts</i>	<i>Microphenocrysts and groundmass</i>	<i>Ocelli</i>	<i>Mantle xenoliths</i>
<i>Coraya</i>	BCO 49	dyke, 30 cm thick	Turbidites of Puncoviscana Formation	porphyritic- strongly panidiomorphic	Ol, Cpx (altered to Cc, Serp)	Abundant Fe-ti oxides, Cpx, Bt	Scarce. Carbonate compositions	Not observed
	BCO 51	sill, 40 cm thick						
	BCO 52	dyke, 50 cm thick						
<i>Yacoraite</i>	P8	sill	Turbidites of Puncoviscana Formation	porphyritic; panidiomorphic, flow differentiation	Amph, Bt, Cpx, Ap	Amph, Cpx, Phlog, Ap	Scarce. Carbonate compositions	Abundant, 0.5-5 cm large. Carbonatitic and Horblendite compositions
	BYA 122	sill		porphyritic	Cpx completely altered to carbonate	Carbonate, Fe-Ti oxides	Not observed	Not observed
	BYA 123	sill		porphyritic, panidiomorphic, crystals rich	Cpx, scarce Ol altered to talc	Cpx, Bt, Fe-Ti oxides	Scarce. Carbonate-analcime compositions	Not observed
	BYA 124a	sill-multiple intrusions		porphyritic, panidiomorphic, crystals rich	Strongly zoned titanian Cpx, Bt, scarce Ol	Cpx, Bt	Scarce. Carbonate compositions	Not observed
	BYA 125	sill		porphyritic	Not observed. Completely altered to Fe-Ti oxides	Not observed. Completely altered to Fe-Ti oxides	Not observed	Not observed
	BYA 126	dyke		porphyritic, panidiomorphic	Amph, Cpx, Bt	Amph, Bt	Not observed	Abundant. Horblendite compositions
	BYA 127	dyke		porphyritic, panidiomorphic, cpx glomeroporphyritic	Ol, Cpx	Cpx, Bt	Abundant. Carbonate compositions	Not observed
<i>Huichaira</i>	BHU 131	sill	Turbidites of Puncoviscana Formation	porphyritic-panidiomorphic-glomeroporphyritic	Cpx, Ol (altered to carbonate and talc), Bt	Bt, Fe-Ti oxides	Abundant. Carbonate-analcime, carbonate and cpx compositions. Segregations	Not observed

The Río Grande Dykes and Sills

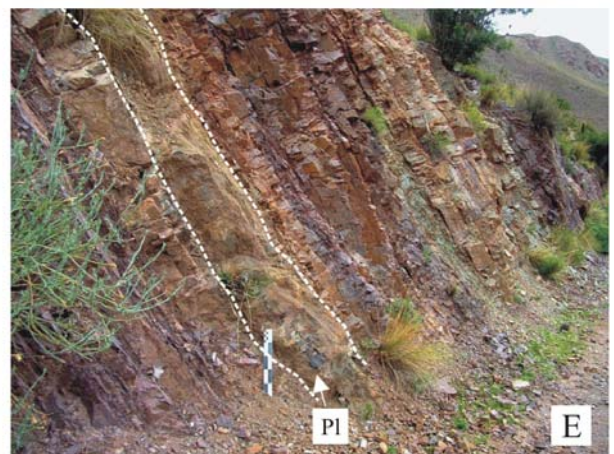
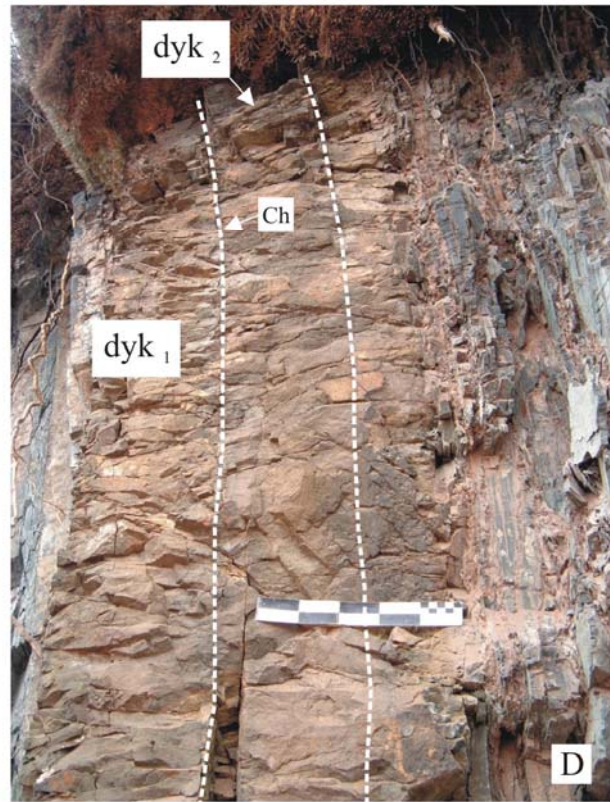
The rocks investigated here are 30-40 cm-thick concordant to discordant in relation with the host Puncoviscana rocks and are folded (Fig. 2) according to the regional structures. Detailed lithological and petrographic characteristics are described in Table 1.

The best exposures of sills and dykes are in the *Yacoraite Valley*, where two textural varieties have been identified: (i) dykes with aphanitic to porphyritic textures with clinopyroxene phenocrysts (Fig. 2 A) and (ii) dykes with mica flakes and amphibole macrocrysts up to 1 cm together with abundant xenoliths (Fig. 2 B, C) concentrated along bands, (e.g. sample P8), indicating flow differentiation (Currie and Ferguson, 1970, Rock, 1987, Rock, 1991). Almost all dykes display a well developed 5-10 cm thick chilled margin and produce thermal metamorphism and bleaching of the host pelites (Fig. 2 A). This suggests that these rocks were emplaced into a relatively cool basement at upper- to medium- crustal levels (Delor and Rock, 1991). The presence of chilled margin inside the BYA 124 dyke, indicate multiple intrusions (Fig. 2 D) or flow differentiation. The multiple intrusions are a typical feature associated with distensive settings.

In the *Coraya Valley* the length/width ratios of dykes are smaller than that of the *Yacoraite valley*. Some dykes display chilled margins and contain fragments of thermometamorphosed pelite (Fig. 2 E). The dykes are strongly altered and their mineralogy is only recognized by crystals habit; the most common rock type is formed by euhedral olivine and pyroxene phenocrysts up to 2 mm in size in a fine-grained green groundmass.

In the *Huichaira Valley*, the sedimentary sequence is completely fractured making it difficult to follow up and characterize individual dykes. Columnar joints and the spotted appearance of rocks, due to red glomeroporphyritic aggregates or individual crystals of altered olivine, are very common.

Figure 2: typical field features of the Río Grande dykes and sills. A) Typical aspect of the Río Grande dykes. They display chilled margin (Ch mg) and the host rocks appear blanked. B) Flow differentiation of mica, amphibole and xenoliths. C) Example of subvolcanic rocks from the Yacoraite Valley. They are partially concordant with the Puncoviscana Formation. D) Zoned internal structures by multiple intrusions. The dyk_2 occur inside the dyk_1 , and the chilled margin (Ch mg) in the border of dyk_2 ; indicate that this was generated in a second time. E) Concordant dykes with thermometamorphic pelite (Pl) inside, from the Coraya locality. The scale has 25 cm.



3. Petrography

3.1. The Río Grande Lamprophyres

The *Coraya Valley* rocks are characterized by porphyritic, panidiomorphic texture with phenocrysts of olivine and pyroxene, almost completely replaced by calcite and serpentine, with pseudo-hexagonal and pseudo-octagonal outlines. The phenocrysts represent approximately 25% of the rock. The groundmass is altered to carbonate and is characterized by the presence of abundant biotite and minor Fe-Ti oxide crystals. These latter also occur as poikilitic inclusions in the rims of olivine and pyroxene phenocrysts.

The *Yacoraite and Huichaira Valleys* rocks display porphyritic to seriate, panidiomorphic textures, globular structures and crustal-mantle xenoliths. Two groups of rocks are recognized.

The first group, represented by samples BYA 123, 124, 127 and BHU 131 (Table 1), is characterized by the presence of zoned Ti-rich clinopyroxene and olivine phenocrysts (Fig. 3 A). Samples BYA 127 and BHU 131 show abundant olivine phenocrysts (Fig. 3 B) whereas samples BYA 123, 124 exhibit predominant clinopyroxene phenocrysts, in both cases they are immersed in a groundmass made of small prismatic crystals of diopside and abundant phlogopite (Fig. 3 B). Globular structures made of carbonate and analcime are common (Fig. 3 I, J). In this group poikilitic phenocrysts of *clinopyroxene with olivine inclusions* are observed (Fig. 3G).

The second group is represented by samples P8 and BYA 126 (Table 1), from the Yacoraite Valley. This group is characterized by zoned amphibole, clinopyroxene and phlogopite phenocrysts together with apatite microphenocrysts (Fig. 3 C, D). Abundant xenoliths also occur. The groundmass is made of carbonate enclosing abundant microliths of biotite, amphibole and apatite (Fig. 3 D). In sample P8, two types of clinopyroxene are present: the first is euhedral to subhedral, has green cores, pale pink rims and is often surrounded by analcime; the second type is anhedral, colourless. Similarly, two distinct varieties of amphibole phenocrysts are recognized (Fig. 3 D). One shows zonating from brown core to pale brown rim and the other displays zonating from green core to pale brown rim. Phlogopite phenocrysts exhibit strong zonating with dark brown core and pale brown rims.

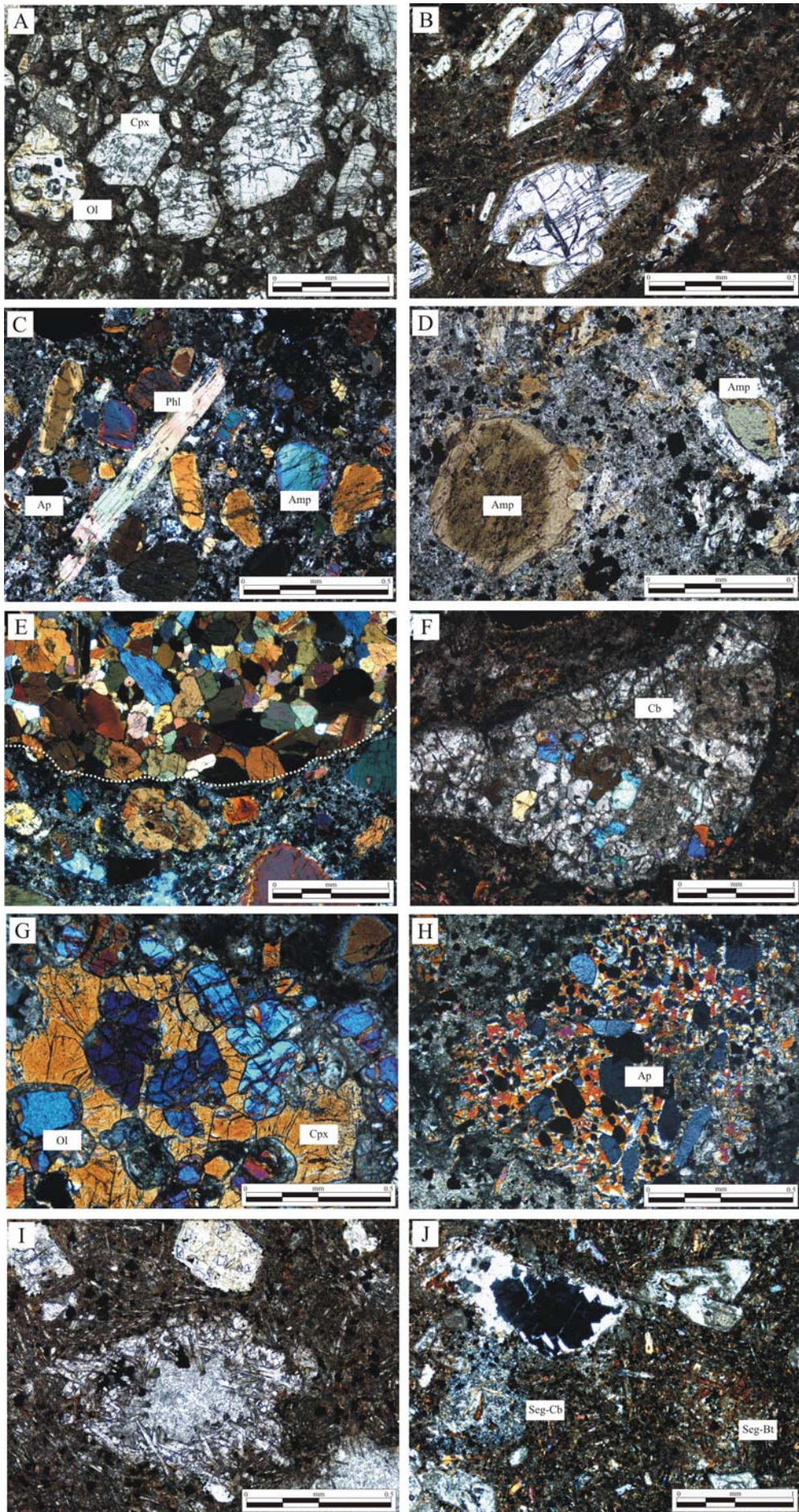


Figure 3: Photomicrographs of several features exhibited by the Río Grande subvolcanic rocks. A) Euhedral and strongly zoned clinopyroxene crystals (Cpx) and minor amounts of euhedral olivine crystals (Ol). Note the high crystals/groundmass proportions. B) Euhedral olivine phenocrysts in a groundmass. These commonly show unstable features as embayment. The groundmass is composed by abundant clinopyroxene and biotite crystals. In this photomicrograph is observed also the Type-III ocelli, completely formed by carbonate. C) Abundant zoned amphibole (Amp)/biotite (Bt) phenocrysts and apatite (Ap) as microphenocrysts characterized the dyke of sample p8. D) The two commonly amphibole types presented. Note the high carbonate and mica concentrations in the groundmass. E) Hornblendite-xenolith with slight mosaic texture. F) Carbonatite-like xenolith. G) Phenocrysts of diopside (Cpx) with inclusion of olivine crystals (Ol). H) Phenocrystal of augite enclosing poikilitically apatite (Ap) crystals. I) Characteristics of **Type-I ocelli**, concentric zones of carbonate surrounded by radial diopside crystals. J) The **Type-II ocelli** and segregations. This ocelli type is marked by analcime forming the central pool surrounded by carbonate. The segregations are characterized by high concentrations of biotite microphenocrysts (Seg-bt) or by high concentrations of carbonate (Seg-Cb).

This group also contains large amounts of varied types of macrocrysts and minor carbonatite-like and hornblendite xenoliths. The *hornblendite xenoliths* (Fig. 3 E) are of variable size (2 mm to 6 cm), rounded to planar. They consist mainly of brown unzoned amphibole and minor brown biotite and clinopyroxene. They display (crystalloblastic polygonal) textures. The *clinopyroxene-bearing carbonate xenoliths* (Fig. 3 F) are small (5 mm in diameter) and dolomite-rich. Minor clinopyroxene and Fe-Ti oxide crystals are also present. Triple junctions between crystals of clinopyroxene and carbonate are observed. The *clinopyroxene phenocrysts with apatite inclusions* (Fig. 3 H) range from 0.5 to 1 mm.

3.2 Globular structures

One characteristic feature of the Río Grande subvolcanic rocks is the presence of ocelli and segregations (Rock, 1991, Fareeddudin, 2001, Azbej, 2006). In the sense of Rock (1991), ocelli are sub-rounded; leucocratic structures and whenever they show no clear limits and grade into their host they may be called “segregations” (Rock 1979, 1991). The ocelli, show globular shapes, display rounded to elongated forms and are approximately 1.0 mm in diameter. Three types of ocelli have been recognized according to their mineral compositions. (i) *Type-I ocelli* have globular to elongated carbonate cores surrounded by radial clinopyroxene and some acicular apatite crystals. The contact between this and their host is invariably sharp Fig. 3 I. (ii) *Type-II* (Fig. 3 J) *ocelli* is more abundant and characterized by analcime surrounded by carbonate crystals and occasionally biotite crystals occur in the rim or inside of ocelli. The contact between this type of ocelli and the host groundmass is marked by carbonate-rich and biotite-poor segregations. The *Type-III ocelli* (Fig. 3 B) are very common

and made of carbonate occasionally surrounded by biotite or Ti-Fe oxide. Besides these three ocelli types, two varieties of segregations have been identified. One type is characterized by the red colour and conspicuous presence of Ti-rich biotite (Seg-Bt) (Fig. 3 J). The other type is richer in carbonate (Seg-Cb) (Fig. 3 J) and also includes biotite crystals. This type is commonly in contact with the ocelli.

Ocelli and segregations have been described in a variety of rock type including basalt (Upton and Wadsworth, 1971), spilite (Smith, 1967), picrite (Drever, 1960) and tholeiitic dolerite (Nakamura and Coombs, 1973), but they are most commonly found in lamprophyres (e.g. Ferguson and Currie, 1971, Cooper, 1979; Phillips, 1968; Philpotts, 1972, 1976, Strong and Harris, 1974; Foley, 1984; Rock, 1991). The large amount of ocelli in the studied rocks suggests high volatile content of the original magma. In the three types of ocelli identified, the predominance of carbonate minerals over biotite or amphibole in some samples, suggests that CO₂ dominate over H₂O in the gas phase (Nedli and Tóth, 2007).

4. Mineral chemistry

The chemical composition of minerals of the R o Grande lamprophyres was obtained with an electron microprobe CAMECABAX SX 50 at the Electron Probe Laboratory of the Universidade de Brasilia. Analytical conditions include beam current of 10 nA, beam energy of 15 KeV and a spot size of 1  m. Acquisition time was 10s. Each element was standardized using either synthetic or natural minerals. Mineral chemistry studies were carried out for olivine, clinopyroxene, amphibole, mica, analcime, apatite and carbonate minerals.

Olivine

Representative chemical analyses of olivine crystals of samples 124a and 127 from the Yacoraite Valley are given in Table 2. The analyses were carried out both on phenocrysts and xenoliths. All analyzed olivine crystals contain 76 to 86 % of forsterite.

Phenocrysts in sample BYA 127 show a slight decrease in forsterite content from cores (86 %) to rims (84 %) however; in general, all phenocrysts, microphenocrysts and xenolith crystals have rather homogeneous compositions.

Clinopyroxene

The analyses of clinopyroxene crystals from the R o Grande rocks are shown in Tables 3 A, B. The analyzed crystals are phenocrysts-microphenocrysts, groundmass crystals, as well as microinclusions in amphibole. Crystals present in the clinopyroxene-bearing carbonate

Table 2: Representative electron microprobe analyses of olivine minerals from the Rio Grande subvolcanic rocks

	<i>Phenocrysts</i>									<i>Microph</i>	<i>in poikilitic clinopyroxene phenocryst</i>		
	Rim			Intermediate			Core						
	1	2	3	1	2	3	1	2	3			4	5
<i>SiO₂</i>	40.32	39.47	39.52	40.21	39.59	38.87	40.22	39.66	39.38	39.49	38.06	40.42	39.60
<i>TiO₂</i>	0.05	0.07	0.00	0.01	0.01	0.03	0.02	0.01	0.05	0.02	0.03	0.03	0.03
<i>Al₂O₃</i>	0.08	0.01	0.02	0.02	0.03	0.03	0.05	0.05	0.06	0.01	0.02	0.03	0.07
<i>FeO_(t)</i>	15.00	15.08	15.03	13.05	14.11	18.94	13.20	13.10	19.14	17.04	21.96	14.97	14.28
<i>MnO</i>	0.22	0.25	0.26	0.15	0.12	0.24	0.18	0.12	0.24	0.22	0.31	0.18	0.19
<i>MgO</i>	44.02	44.12	43.70	45.56	44.79	41.14	45.21	45.68	41.16	42.86	39.05	44.36	44.50
<i>CaO</i>	0.42	0.40	0.38	0.21	0.26	0.15	0.24	0.23	0.11	0.31	0.13	0.33	0.21
<i>Cr₂O₃</i>	0.08	0.09	0.05	0.04	0.04	0.03	0.03	0.07	0.01	0.01	0.01	0.00	0.03
<i>NiO</i>	0.19	0.15	0.20	0.32	0.21	0.21	0.31	0.20	0.16	0.29	0.18	0.24	0.30
<i>K₂O</i>	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.00	0.01	0.00	0.01	0.00
<i>Total</i>	100.38	99.63	99.16	99.58	99.17	99.64	99.46	99.13	100.30	100.25	99.75	100.56	99.20
<i>FO</i>	84.24	84.14	84.03	86.27	85.18	79.44	86.04	86.30	79.22	81.92	75.89	84.30	84.82
<i>FA</i>	16.11	16.13	16.21	13.86	15.05	20.52	14.10	13.88	20.67	18.28	23.94	15.95	15.27
<i>mg#</i>	0.75	0.75	0.74	0.78	0.76	0.68	0.77	0.78	0.68	0.72	0.64	0.75	0.76

xenolith have been also analyzed. According to the nomenclature proposed by Morimoto et al., (1988), all the analyzed crystals belong to the calcic series.

Most of them fall in the Fe-diopside field but some macrocrysts and those in clinopyroxene-bearing carbonate xenoliths plot in the Mg-augite field (Fig. 4 A). The diopside grains are characterized by wollastonite contents of 45-52%, whereas the augite grains have wollastonite contents of 39-45%.

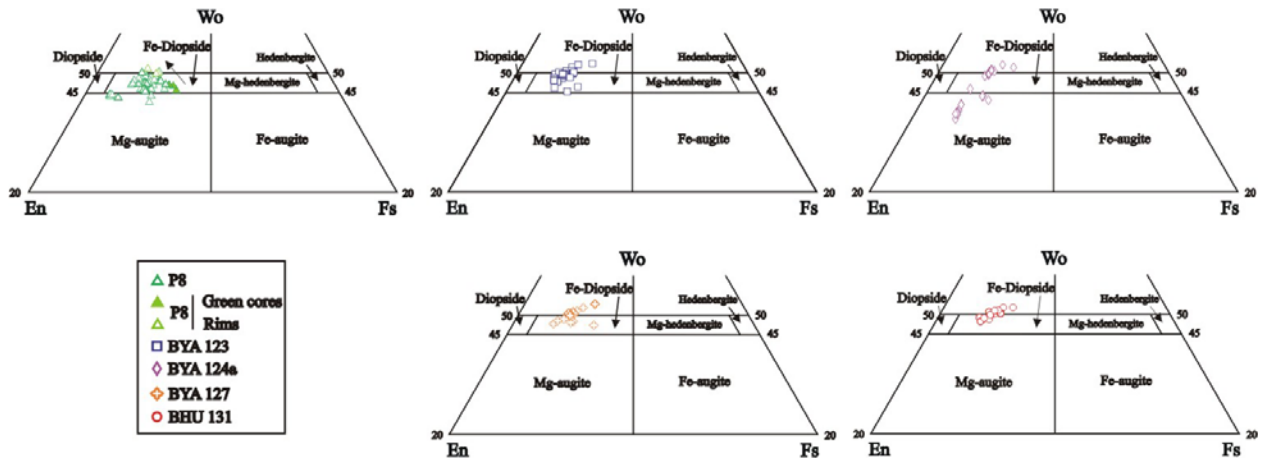


Figure 4: Representative pyroxene and olivine compositions of the Río Grande subvolcanic rocks. Wollastonite (Wo)-Enstatite (En)-Ferrosilite (Fs) classification diagram after Morimoto et al., (1988). They are calcic clinopyroxene, almost all of them fall in the field of Fe-diopside but samples P8 and BYA 124a also fall in the Mg-augite field.

Zoned phenocrysts with green cores and white rims have been recognized only in sample P8. They are characterized by lowest Ti and highest Na contents compared with the diopside crystals from other samples. The cores (green) of these crystals are characterized by the highest Na₂O contents (~2.5 wt. %), and the lowest mg# (~67) showing reverse zoning (increasing mg#, decreasing Na from core to rim). The cores of white diopside are less sodic (<2.5 wt. %), present higher mg# (~78) (Table 3 B) and show normal zoning (decreasing mg#, increasing Na from core to rim). Both groups show similar rim compositions with intermediate Na₂O and mg# (~74) values. In the rest of samples, clinopyroxene have increasing Ti content from the core to the rim and some crystal rims are classified as Titanian-

Table 3B: Representative electron microprobe analyses of clinopyroxene minerals from sample P8, Río Grande subvolcanic rocks.

	<i>Phenocryst</i>								<i>Microph</i>		<i>Incl</i>	<i>Xenoliths**</i>			<i>Poikilitic phenocryst***</i>
	<i>Green cpx</i>				<i>White cpx</i>				5	6	7	8	9	10	11
	Rim		Core		Rim		Core								
	1	2	1	2	3	4	3	4							
<i>SiO₂</i>	48.00	46.44	51.91	51.34	48.26	46.85	52.10	50.81	53.27	51.41	53.11	53.64	53.50	53.03	52.18
<i>TiO₂</i>	1.52	3.06	0.54	0.80	1.60	2.34	0.87	1.09	0.25	0.42	0.37	0.15	0.48	0.19	0.23
<i>Al₂O₃</i>	5.50	6.84	3.27	3.21	4.80	6.73	3.96	5.68	1.36	4.54	1.30	4.00	4.23	0.95	1.88
<i>FeO_(t)</i>	7.67	6.37	9.53	10.24	7.79	7.82	6.80	6.73	7.63	7.89	7.35	3.31	4.38	7.37	8.89
<i>MnO</i>	0.14	0.06	0.19	0.22	0.16	0.20	0.15	0.12	0.21	0.16	0.31	0.06	0.14	0.19	0.23
<i>MgO</i>	12.46	13.13	11.21	11.12	12.73	11.95	14.15	13.53	13.43	12.65	13.78	16.07	15.50	14.01	14.08
<i>CaO</i>	22.81	23.92	20.06	20.30	22.64	22.66	20.94	20.71	20.90	21.60	21.06	19.33	19.20	20.95	19.47
<i>Cr₂O₃</i>	0.02	0.00	0.00	0.03	0.01	0.05	0.55	0.30	0.20	0.00	0.07	1.11	0.84	0.15	0.01
<i>NiO</i>	0.00	0.03	0.08	0.00	0.00	0.05	0.00	0.00	0.00	0.03	0.05	0.06	0.03	0.01	0.02
<i>Na₂O</i>	0.82	0.44	2.51	2.35	0.53	0.89	1.24	1.37	2.06	1.15	1.65	1.71	2.09	1.50	1.72
<i>K₂O</i>	0.00	0.01	0.00	0.02	0.00	0.02	0.00	0.02	0.02	0.00	0.02	0.01	0.02	0.02	0.01
<i>Total</i>	99.41	100.77	99.80	100.19	98.83	100.07	100.95	100.58	99.81	100.02	99.40	99.52	100.64	98.68	99.21
<i>FeO</i>	3.40	2.14	5.08	5.11	4.85	3.34	5.13	4.59	3.38	6.10	4.25	2.86	2.25	4.57	4.32
<i>Fe₂O₃</i>	4.75	4.70	4.97	5.70	3.27	4.98	1.85	2.38	4.73	1.95	3.45	0.50	2.37	3.11	5.08
Structural formula based on 6 oxygens															
Si	1.80	1.72	1.93	1.91	1.82	1.75	1.90	1.86	1.97	1.90	1.97	1.95	1.93	1.98	1.94
Al^{IV}	0.20	0.28	0.07	0.09	0.18	0.25	0.10	0.14	0.03	0.10	0.03	0.05	0.07	0.02	0.06
Fe³⁺	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Sitio T	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00
Al^{VI}	0.04	0.01	0.08	0.05	0.03	0.05	0.07	0.10	0.03	0.10	0.03	0.12	0.11	0.02	0.03
Ti	0.04	0.09	0.02	0.02	0.05	0.07	0.02	0.03	0.01	0.01	0.01	0.00	0.01	0.01	0.01
Fe³⁺	0.13	0.13	0.14	0.16	0.09	0.14	0.05	0.07	0.13	0.05	0.10	0.01	0.06	0.09	0.14
Fe²⁺	0.11	0.07	0.16	0.16	0.15	0.11	0.16	0.14	0.10	0.19	0.13	0.09	0.07	0.14	0.13
Mn	0.00	0.00	0.01	0.01	0.01	0.01	0.00	0.00	0.01	0.00	0.01	0.00	0.00	0.01	0.01
Mg	0.70	0.72	0.62	0.62	0.72	0.67	0.77	0.74	0.74	0.70	0.76	0.87	0.83	0.78	0.78
Cr	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.01	0.01	0.00	0.00	0.03	0.02	0.00	0.00
Sitio O	1.03	1.02	1.02	1.02	1.05	1.03	1.09	1.09	1.02	1.06	1.04	1.13	1.11	1.05	1.10
Ca	0.92	0.95	0.80	0.81	0.92	0.91	0.82	0.81	0.83	0.86	0.84	0.75	0.74	0.84	0.78
Na	0.06	0.03	0.18	0.17	0.04	0.06	0.09	0.10	0.15	0.08	0.12	0.12	0.15	0.11	0.12
K	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Sitio C	0.97	0.98	0.98	0.98	0.95	0.97	0.91	0.91	0.98	0.94	0.96	0.87	0.89	0.95	0.90
Wo	49.44	50.72	46.54	46.38	48.76	49.92	45.58	46.23	45.89	47.65	45.81	43.66	43.45	45.35	42.32
En	37.59	38.74	36.19	35.37	38.15	36.63	42.86	42.04	41.04	38.82	41.70	50.50	48.82	42.20	42.60
Fs	12.97	10.55	17.27	18.26	13.09	13.45	11.56	11.72	13.08	13.53	12.49	5.84	7.74	12.45	15.08
mg#*	74.35	78.60	67.70	65.95	74.45	73.14	78.76	78.20	75.83	74.15	76.96	89.64	86.32	77.22	73.86

*mg# = 100*Mg/(Mg+Fetot)

Xenoliths: 8 and 9, from cpx-bearing carbonate xenolith and 10, amphibole-cpx xenolith and * Poikilitic phenocrystal with apatite inclusions.

diopside (> 0.10 Ti apfu). In the clinopyroxene-bearing carbonate xenolith, augite crystals are classified as chromian-sodian augite, they have the highest Cr_2O_3 content (up 1.1 wt. %) and mg# (~ 90) (Table 3B). In the $\text{Al}_2\text{O}_3/\text{mg}\#$ vs. TiO_2 diagram (Fig. 5 A) clinopyroxene crystals from samples BYA127 and BHU131 display higher TiO_2 contents, compared to pyroxenes in sample P8. Clinopyroxenes of samples BYA123 and 124 seem to be intermediate in composition between the two trends above.

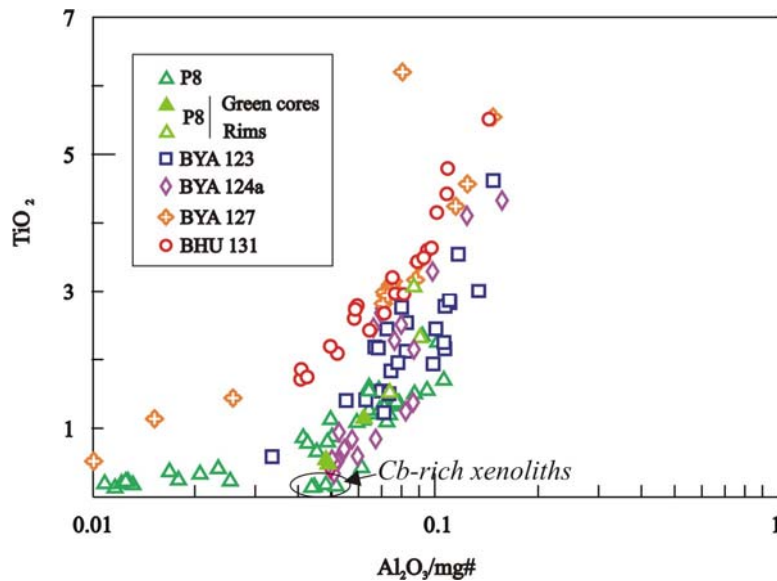


Figure 5: $\text{Al}_2\text{O}_3/\text{mg}\#$ vs. TiO_2 diagram showing the compositions of clinopyroxene crystals. In this diagram are observed two groups of clinopyroxens, differentiated for the TiO_2 contents.

Amphibole

Representative analyses of amphibole crystals are presented in Table 4. This mineral was recognized only in sample P8. The analyses were carried out on brown and green amphibole phenocrysts and xenolith crystals. According to the classification proposed by

Table 4: Representative electron microprobe analyses of amphibole minerals from the Rio Grande subvolcanic rocks

	Phenocrysts										Xenoliths	
	Brown amphibole				Green amphibole						5	6
	Rim		Core		Rim		Core		Inter			
1	2	1	2	3	4	3	4	4	5	6		
SiO ₂	41.22	40.30	40.75	40.99	40.13	41.49	41.01	42.77	41.30	43.53	44.24	
TiO ₂	2.84	3.90	3.99	4.15	3.96	3.46	1.24	1.40	1.22	2.97	2.33	
Al ₂ O ₃	12.07	13.16	12.62	12.45	12.97	11.28	14.45	13.38	14.12	11.57	10.11	
Fe ₂ O ₃	1.33	0.06	0.47	0.82	0.56	0.31	3.07	2.52	3.12	0.19	2.04	
FeO	8.84	10.38	11.74	12.61	10.09	13.37	10.70	9.46	8.92	8.47	10.45	
Cr ₂ O ₃	0.00	0.03	0.07	0.00	0.01	0.07	0.00	0.00	0.01	0.47	0.21	
MnO	0.17	0.14	0.13	0.17	0.14	0.24	0.25	0.30	0.23	0.10	0.20	
MgO	14.57	13.41	12.19	11.67	13.69	11.48	11.13	12.80	12.54	15.27	13.79	
CaO	11.90	11.99	10.64	10.58	11.94	11.24	9.63	9.76	10.14	11.39	10.10	
Na ₂ O	2.74	2.29	2.80	2.70	2.50	3.13	4.02	3.98	3.74	3.29	4.20	
K ₂ O	1.42	1.99	2.11	2.00	1.90	1.86	1.06	1.33	1.29	1.25	1.15	
Total	97.08	97.65	97.50	98.15	97.89	97.94	96.55	97.69	96.60	98.51	98.82	
Structural formula based on 23 oxygens												
Si	6.11	5.98	6.09	6.10	5.95	6.22	6.14	6.28	6.15	6.30	6.46	
Al ^{IV}	1.89	2.02	1.91	1.90	2.05	1.78	1.86	1.72	1.85	1.70	1.54	
Sum T	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	
Al ^{VI}	0.22	0.28	0.31	0.29	0.21	0.22	0.69	0.60	0.62	0.28	0.20	
Fe ³⁺	0.15	0.01	0.05	0.09	0.06	0.04	0.35	0.28	0.35	0.02	0.22	
Ti	0.32	0.44	0.45	0.46	0.44	0.39	0.14	0.15	0.14	0.32	0.26	
Cr	0.00	0.00	0.01	0.00	0.00	0.01	0.00	0.00	0.00	0.05	0.02	
Mg	3.22	2.97	2.72	2.59	3.03	2.57	2.49	2.80	2.78	3.29	3.00	
Fe ²⁺	1.09	1.29	1.47	1.57	1.25	1.68	1.33	1.16	1.11	1.03	1.28	
Zn	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	
Mn	0.00	0.02	0.00	0.00	0.01	0.03	0.00	0.00	0.00	0.01	0.02	
Sum C	5.00	5.00	5.00	5.00	5.00	4.93	5.00	5.00	5.00	5.00	5.00	
Mg	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	
Fe ²⁺	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.00	
Zn	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	
Mn	0.02	0.00	0.01	0.02	0.01	0.00	0.03	0.04	0.03	0.01	0.01	
Ca	1.89	1.91	1.70	1.69	1.90	1.81	1.54	1.54	1.62	1.77	1.58	
Na	0.09	0.09	0.28	0.29	0.09	0.19	0.42	0.43	0.36	0.23	0.41	
Sum B	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	
Na	0.70	0.57	0.53	0.49	0.63	0.72	0.75	0.71	0.72	0.70	0.78	
K	0.27	0.38	0.40	0.38	0.36	0.36	0.20	0.25	0.24	0.23	0.21	
Sum A	0.97	0.95	0.93	0.87	0.99	1.07	0.95	0.96	0.97	0.93	0.99	
(Ca+Na) _B	1.98	2.00	1.99	1.97	1.99	2.00	1.96	1.96	1.97	1.99	1.99	
Na _B	0.09	0.09	0.28	0.29	0.09	0.19	0.42	0.43	0.36	0.23	0.41	
Ca _B	1.89	1.91	1.70	1.69	1.90	1.81	1.54	1.54	1.62	1.77	1.58	
(Na+K) _A	0.97	0.95	0.93	0.87	0.99	1.07	0.95	0.96	0.97	0.93	0.99	
Si	6.11	5.98	6.09	6.10	5.95	6.22	6.14	6.28	6.15	6.30	6.46	
Mg/(Mg+Fe²⁺)	0.75	0.70	0.65	0.62	0.71	0.60	0.65	0.71	0.71	0.76	0.70	

Leake et al., (1997), the amphibole compositions of sample P8 belong to the calcic group (Fig. 6) with $(Ca+Na)_B > 1.00$ and $Na_B < 0.50$ (fig). All crystals have pargasitic compositions ($Al^{IV} > Fe^{+3}$) and due to their Ti apfu content ($0.25 < Ti < 0.49$) these amphiboles are classified as “*titanian pargasite*”. Only one macrophenocryst is edenite in composition (Si in formula > 6.5).

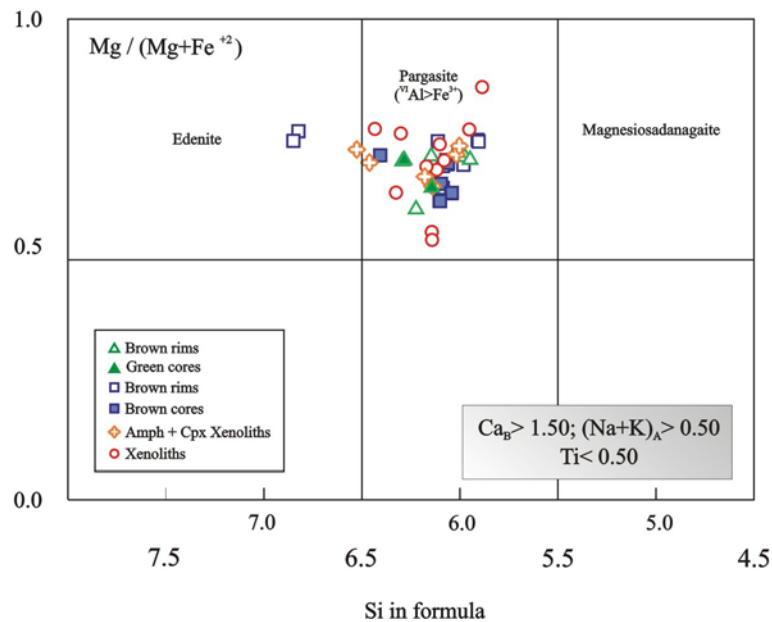


Figure 6: Classification diagram for amphiboles after Leake et al., 1997. All crystals are plotted in the calcic amphiboles field and are classified as Titanian-Pargasite.

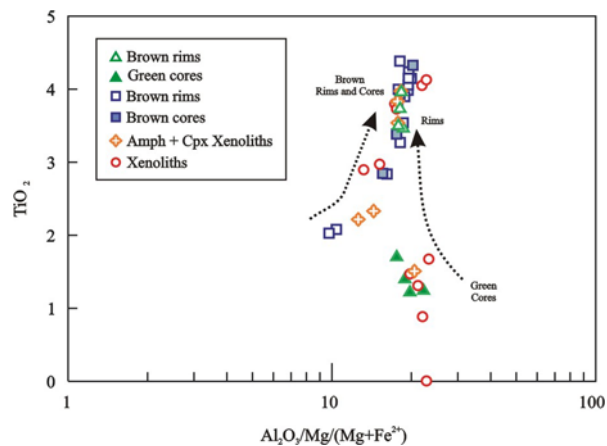


Figure 7: $Al_2O_3/mg\#$ vs. TiO_2 diagram shows the lower TiO_2 compositions for some green cores and xenoliths crystals. Observed that the rims of crystals with green cores have similar compositions with the brown amphiboles.

Green and brown amphibole show compositional differences in their cores. Green amphibole cores have $mg\#$ around 0.68, TiO_2 contents of approximately 2 % and show higher Al_2O_3 (13.0 wt. %), whereas cores of brown amphibole have $mg\#$ between 0.63-0.70, TiO_2 of 4

% and are less rich in Al₂O₃ (12.5 wt. %). Rims of green and brown amphiboles have light brown colour and show higher mg# (0.70) compared to their cores and are more homogeneous in TiO₂ and Al₂O₃ contents (3.5% and 12.0% respectively). These differences are evident in Al₂O₃/mg# vs. TiO₂ diagram (Fig. 7A).

Amphibole compositions are plotted in the Al₂O₃/TiO₂ vs. CaO/Na₂O diagram of Rock, 1987 (Fig. 7B). All titanian-pargasite fall in the intersection of Ultramafic lamprophyres-Alkaline lamprophyres with Calcio-alkaline lamprophyres (UML-AL from CAL lamprophyres) fields.

Mica

Representative analyses of mica crystals from the Río Grande rocks appear in Table 5. The analyzed crystals are represented by phenocrysts, mesophenocrysts and groundmass crystals. Microinclusions in amphibole and clinopyroxene as well as their presence in the globular structures (ocelli and segregations) have also been analyzed. Mica compositions of phenocrysts, microphenocrysts, groundmass crystals and inclusions belong to the phlogopite field (Fig. 8), whereas biotite composition is observed in microphenocrysts inside and around the globular structures (ocelli and segregations).

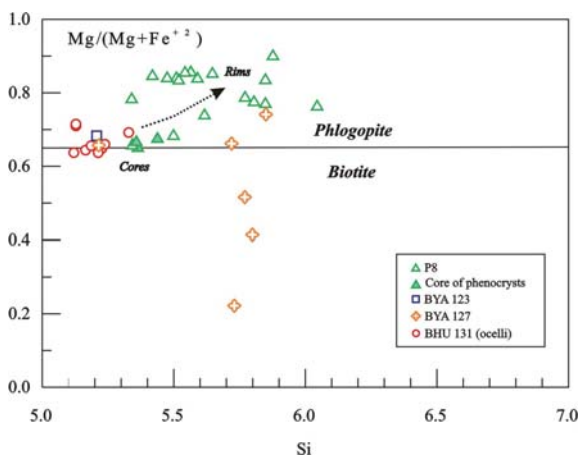


Figure 8: Classification diagram for mica compositions after Rieder et al., 1998. Note the ore-rim trend followed by singular grains.

The phlogopite crystals have Mg/Mg+Fe⁺² ratios from 0.66 to 0.89; whereas biotite shows values ranging between 0.61 and 0.66. The cores of phenocrysts have similar compositions to microphenocrysts located in ocelli and segregations, showing the highest TiO₂ contents (~8 % wt.), whereas the rims of phenocrysts have similar compositions to mesophenocrysts and microphenocrysts with TiO₂ about 4 % (Fig. 8).

Table 5: Representative electron microprobe analyses of mica minerals from the Rio Grande subvolcanic rocks

	Phenocrysts		Microphenocrysts			Incl	Amp rim	In segregations					Ocelli			
	Rim		Core				bt-rich			cb-rich						
	1	2	1	2	3	4	5	6	7	Rim	Core	Core	10	11	13	
										8						
SiO ₂	37.61	35.80	36.57	36.17	37.47	39.27	38.24	38.89	37.41	35.49	34.25	34.25	34.67	33.61	36.96	
TiO ₂	3.24	6.06	5.23	6.29	3.19	2.68	2.87	3.13	3.16	6.82	7.88	7.85	5.76	5.59	3.68	
Al ₂ O ₃	14.99	15.24	14.83	15.27	14.30	14.76	14.06	14.16	14.86	14.79	15.51	15.33	15.52	14.51	10.11	
FeO	7.06	13.62	13.21	13.82	7.42	6.61	7.29	11.90	7.16	11.43	10.23	11.90	12.94	12.39	22.30	
Fe ₂ O ₃	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	1.06	2.08	1.49	0.48	2.19	4.83	
Cr ₂ O ₃	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	
MnO	0.22	0.10	0.04	0.11	0.15	0.19	0.17	0.12	0.16	0.10	0.10	0.06	0.11	0.25	0.43	
MgO	20.67	14.62	15.29	14.30	20.80	21.21	21.12	18.85	20.97	14.96	14.86	13.38	15.63	14.14	9.37	
CaO	0.04	0.00	0.03	0.01	0.04	0.11	0.08	0.01	0.07	0.11	0.05	0.12	0.08	0.29	0.21	
Na ₂ O	0.40	0.60	0.82	0.72	0.68	0.89	0.61	1.36	0.74	0.38	0.46	0.75	0.35	0.27	0.00	
K ₂ O	9.12	9.16	9.23	9.19	9.26	9.04	9.37	8.31	8.92	8.64	8.71	8.47	9.22	9.10	9.49	
BaO	1.14	0.00	0.28	0.00	1.03	0.74	1.17	0.00	1.57	0.98	2.00	1.39	0.43	2.08	0.00	
Cl	0.00	0.02	0.03	0.02	0.01	0.00	0.00	0.02	0.00	0.02	0.00	0.02	0.02	0.01	0.02	
H ₂ O	4.09	4.01	4.02	4.03	4.06	4.16	4.09	4.14	4.01	3.99	4.02	3.96	3.99	3.88	3.86	
Total	98.57	99.24	99.58	99.92	98.38	99.67	99.07	100.90	99.03	98.75	100.14	98.95	99.18	98.29	101.25	
Structural formula based on 22 oxygens																
Si	5.51	5.34	5.44	5.36	5.52	5.65	5.59	5.62	5.48	5.33	5.13	5.19	5.21	5.22	5.80	
Al ^{IV}	2.49	2.66	2.56	2.64	2.48	2.35	2.41	2.38	2.52	2.62	2.74	2.74	2.75	2.65	1.87	
Fe ³⁺	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.05	0.14	0.08	0.05	0.13	0.33	
Sum T	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	
Al ^{VI}	0.10	0.03	0.04	0.03	0.00	0.15	0.01	0.03	0.04	0.00	0.00	0.00	0.00	0.00	0.00	
Fe ²⁺ ₁	0.86	1.70	1.64	1.71	0.91	0.79	0.89	1.44	0.88	1.38	1.14	1.43	1.58	1.48	2.59	
Fe ³⁺ ₂	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.07	0.10	0.09	0.01	0.12	0.24	
Ti	0.36	0.68	0.59	0.70	0.35	0.29	0.32	0.34	0.35	0.77	0.89	0.89	0.65	0.65	0.43	
Mg	4.51	3.25	3.39	3.16	4.57	4.55	4.60	4.06	4.57	3.35	3.32	3.02	3.50	3.27	2.19	
Mn	0.03	0.01	0.01	0.01	0.02	0.02	0.01	0.01	0.02	0.01	0.01	0.01	0.01	0.03	0.06	
Sum O	5.86	5.67	5.66	5.62	5.85	5.80	5.84	5.88	5.86	5.58	5.46	5.44	5.75	5.56	5.51	
Ca	0.01	0.00	0.00	0.00	0.01	0.02	0.01	0.00	0.01	0.02	0.01	0.02	0.01	0.05	0.04	
Na	0.11	0.17	0.24	0.21	0.19	0.25	0.17	0.38	0.21	0.11	0.13	0.22	0.10	0.08	0.00	
K	0.85	0.87	0.88	0.87	0.87	0.83	0.87	0.77	0.83	1.66	1.66	1.64	1.77	1.80	1.90	
Ba	0.13	0.00	0.03	0.00	0.12	0.08	0.13	0.00	0.18	0.06	0.12	0.08	0.03	0.13	0.00	
Sum A	1.10	1.05	1.15	1.08	1.19	1.18	1.19	1.15	1.23	1.78	1.80	1.88	1.88	1.93	1.93	
Cations	14.96	14.72	14.81	14.70	15.04	14.98	15.04	15.03	15.09	15.36	15.26	15.32	15.63	15.49	15.45	
Mg/(Mg+Fe ²⁺)	0.84	0.66	0.67	0.65	0.83	0.85	0.84	0.74	0.84	0.69	0.71	0.66	0.68	0.66	0.41	
Fe ²⁺ /(Fe ²⁺ +Mg)	0.16	0.34	0.33	0.35	0.17	0.15	0.16	0.26	0.16	0.31	0.29	0.34	0.32	0.34	0.59	

Obs: H₂O recalculated by stoichiometry. Fe₂O₃ recalculated for samples with deficiency in tetrahedral site

The compositions of groundmass phlogopite and some phenocrysts and microphenocrysts rims is marked by higher Mg concentrations compared to cores of phenocrysts and microphenocrysts. This variation is accompanied by decrease in TiO_2 and Al_2O_3 contents. On the other hand, biotite crystals are characterized by lower Mg and higher TiO_2 and Al_2O_3 contents (Fig. 9 A, B) The biotite crystals in segregations, display higher TiO_2 (from 5.59 to 7.88 wt. %) and BaO (from 0.43 to 2.08 wt. %) values (Fig. 9 B).

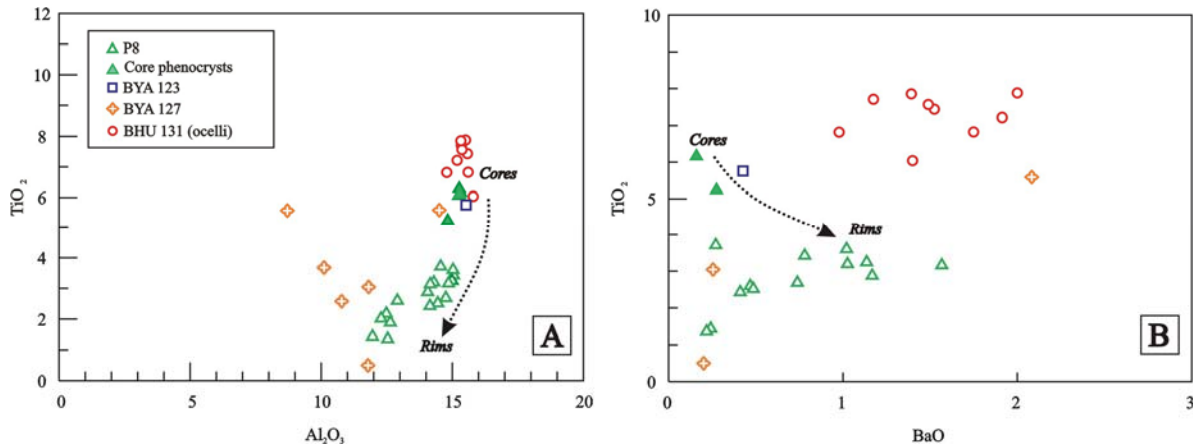


Figure 9: A) Al_2O_3 vs. TiO_2 diagram and B) BaO vs. TiO_2 diagram for micas. The mica from ocelli (samples BHU 131) and cores of phenocrysts are richer in TiO_2 (~7.5 wt. %) and BaO (~2 wt. %) compared with the rest. Note the core-rim trend (arrows) followed by singular grains.

Apatite

Chemical analyses for two apatite crystals are given in Table 6. This phase was only recognized in sample P8 and one corresponds to a phenocryst and the other corresponds to a crystal from a *clinopyroxene phenocrysts with apatite inclusions*. The phenocryst is characterized by the higher SrO content (1.075 wt. %), whereas the apatite in a nodule have SrO values between 0.30 to 0.43 wt. %.

Analcime

Chemical analyses of analcime are given in Table 7. This constitutes a subordinate phase in the groundmass and an important mineral in ocelli and segregations. They are compositionally identical and have 57 wt. % SiO_2 , 24 wt. % Al_2O_3 and 13 wt. % Na_2O and very low CaO and K_2O contents. In the ocelli, analcime crystals show slight increases in Ba content from 0 to 0.11 wt. % BaO from core to rim.

Table 6: Representative analyses microprobe for apatite minerals from sample P8, Rio Grande subvolcanic rocks

	<i>Phenocryst</i>		<i>In augite phenocrysts</i>	
	<i>1</i>	<i>2</i>	<i>3</i>	
<i>SiO₂</i>	0.00	0.13	0.12	
<i>Al₂O₃</i>	0.00	0.00	0.01	
<i>MgO</i>	0.14	0.38	0.34	
<i>Fe₂O₃</i>	0.34	0.34	0.32	
<i>CaO</i>	53.31	53.83	54.50	
<i>Na₂O</i>	0.35	0.06	0.12	
<i>K₂O</i>	0.00	0.01	0.03	
<i>P₂O₅</i>	41.67	40.14	40.51	
<i>Cl</i>	0.24	0.13	0.11	
<i>SrO</i>	1.08	0.43	0.31	
<i>BaO</i>	0.01	0.00	0.00	
<i>H₂O</i>	1.70	1.69	1.72	
<i>Total</i>	98.83	97.14	98.07	
Cations per fórmula unit				
Na	0.12	0.02	0.04	
Si	0.00	0.02	0.02	
Al	0.00	0.00	0.00	
Mg	0.04	0.10	0.09	
K	0.00	0.00	0.01	
P	6.02	5.90	5.90	
Ca	9.74	10.02	10.04	
Cl	-0.05	-0.03	-0.02	
Sr	0.11	0.04	0.03	
Fe	0.05	0.05	0.05	
Ba	0.00	0.00	0.00	
H	0.56	0.57	0.57	
<i>Cations</i>	16.57	16.69	16.71	

Carbonate

This mineral phase has been analyzed in smaller crystals in the groundmass, in xenoliths, in globular structures (ocelli and segregations) and in carbonate rims of one amphibole phenocryst (sample P8). Their results are shown in Table 8.

Dolomite is the principal composition (Fig. 10 A) observed in Río Grande lamprophyres. Only one carbonate mineral, which occurs in segregations, is breunnerite in composition. The breunnerite mineral seems to occur only in ultramafic lamprophyres (UML) and lamproites (LL) (Rock, 1987).

Table 7: Representative electron microprobe analyses of analcime minerals from the Rio Grande subvolcanic rocks

	<i>Rim to amphibole</i>		<i>in Ocelli</i>						<i>in Bt-rich segreg</i>	<i>Microphenocrysts</i>		
	1	2	3	4	5	6	7	8	9	10	11	12
<i>SiO₂</i>	55.51	57.23	53.20	55.76	52.18	54.66	59.20	61.92	58.02	62.77	50.61	52.08
<i>Al₂O₃</i>	24.57	24.63	23.73	21.17	24.92	25.46	24.85	22.71	24.91	22.47	22.72	23.19
<i>FeO</i>	0.14	0.01	0.07	0.09	0.63	0.51	0.13	0.18	0.16	0.26	0.04	1.23
<i>CaO</i>	0.16	0.00	0.30	0.11	0.63	0.69	0.05	0.01	0.13	0.08	0.29	0.32
<i>Na₂O</i>	12.94	13.26	12.93	13.96	10.92	10.46	13.58	12.18	12.24	10.91	11.46	11.85
<i>K₂O</i>	0.12	0.04	0.25	0.12	0.34	0.20	0.08	0.08	0.05	0.09	0.27	0.55
<i>Total</i>	93.43	95.28	90.49	91.30	89.62	91.99	97.90	97.08	95.50	96.69	85.42	89.23
Recalculate cations in basis of 7 (O)												
Si	2.20	2.39	2.23	2.33	2.18	2.29	2.48	2.59	2.43	2.63	2.12	2.18
Al	1.15	1.21	1.17	1.04	1.23	1.26	1.22	1.12	1.23	1.11	1.12	1.14
Sum T	3.35	3.61	3.39	3.38	3.41	3.54	3.70	3.71	3.65	3.73	3.24	3.32
Fe⁺²_t	0.00	0.00	0.00	0.00	0.02	0.02	0.00	0.01	0.01	0.01	0.00	0.04
Ca	0.01	0.00	0.01	0.01	0.03	0.03	0.00	0.00	0.01	0.00	0.01	0.01
Na	1.00	1.08	1.05	1.13	0.89	0.85	1.10	0.99	0.99	0.88	0.93	0.96
K	0.01	0.00	0.01	0.01	0.02	0.01	0.00	0.00	0.00	0.00	0.01	0.03
Sum A	1.01	1.08	1.08	1.15	0.95	0.91	1.11	1.00	1.01	0.90	0.96	1.05
Cations	4.36	4.69	4.47	4.52	4.37	4.45	4.81	4.71	4.66	4.64	4.20	4.37

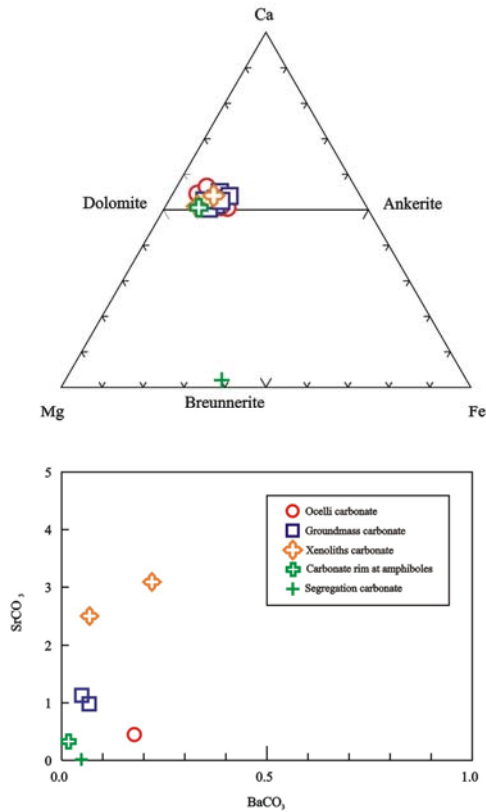


Figure 10: Carbonate compositions of Río Grande lamprophyres. A) Ca-Mg-Fe system (moles %). All types of carbonate are dolomite in composition except for the carbonate segregations that is breunnerite. B) BaCO₃-SrCO₃ covariation diagram. Note that xenoliths carbonate have the highest contents in SrCO₃.

The dolomite compositions from the clinopyroxene-bearing carbonate xenolith are Sr rich (2.5-3.0 wt. % SrCO₃) and moderately Ba rich (0.2 wt. % BaCO₃) (Fig. 10 B).

5. Analytical Procedures

Samples for geochemical analyses were washed and weathered and veined surfaces were cut off. The rocks were crushed and milled to a very fine powder. Major elements and Ni and Sc (ppm) were analyzed by ICP-OES and trace element analyses were carried out by ICP-MS at ACME Laboratories in Canada.

Table 8: Representative electron microprobe analyses of carbonate minerals from the Rio Grande subvolcanic rocks

	In Type-I ocelli							In segregations (Breunnerite)	Groundmass					Cpx-bearing carbonate xenoliths		Rim to amphibole	
	1	2	3	4	5	6	7		1	2	3	4	5	1	2		
<i>FeCO₃</i>	12.70	9.69	7.08	12.12	8.62	15.77	17.70	15.87	45.15	13.02	16.01	14.05	12.85	14.68	11.77	9.21	9.84
<i>MnCO₃</i>	0.56	0.68	0.75	0.57	1.12	0.91	0.96	1.16	0.13	1.09	1.36	1.10	0.73	1.82	0.22	0.37	0.80
<i>MgCO₃</i>	30.73	34.43	33.91	33.86	30.92	30.07	28.63	30.19	51.40	28.35	26.32	30.73	33.19	28.74	29.77	35.41	35.26
<i>CaCO₃</i>	51.20	52.15	55.63	52.19	57.30	50.81	50.39	51.67	2.07	54.75	52.64	51.35	51.11	51.77	53.32	51.89	51.58
<i>SrCO₃</i>	2.06	0.80	0.45	0.15	0.04	0.45	0.47	0.62	0.02	0.96	0.71	0.48	0.39	1.13	3.09	2.50	0.33
<i>BaCO₃</i>	0.00	0.00	0.18	0.00	0.00	0.00	0.00	0.00	0.05	0.00	0.00	0.00	0.00	0.05	0.22	0.07	0.02
<i>TOTAL</i>	98.57	99.10	99.34	100.25	99.35	99.35	99.47	100.88	100.16	99.50	98.35	99.05	99.61	99.52	99.72	100.80	99.18
Cations per fórmula unit																	
Ca	1.02	1.02	1.08	0.95	1.07	1.00	1.00	1.01	0.04	1.05	1.05	0.52	1.00	1.03	1.05	1.00	0.97
Mg	0.73	0.80	0.78	0.73	0.68	0.71	0.67	0.70	1.09	0.65	0.63	0.37	0.77	0.68	0.70	0.81	0.79
Fe	0.22	0.16	0.12	0.19	0.14	0.27	0.30	0.27	0.70	0.22	0.28	0.12	0.22	0.25	0.20	0.15	0.16
Ba	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Sr	0.03	0.01	0.01	0.00	0.00	0.01	0.01	0.01	0.00	0.01	0.01	0.00	0.01	0.02	0.04	0.03	0.00
Mn	0.01	0.01	0.01	0.01	0.02	0.02	0.02	0.02	0.00	0.02	0.02	0.01	0.01	0.03	0.00	0.01	0.01
<i>Cations</i>	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00

Sm-Nd and Rb-Sr isotopic date analyses followed the method described by Gioia and Pimentel (2000) and were carried out at the Geochronology Laboratory of the University of Brasilia, Brazil. Whole-rock powers (~100 mg) were mixed with ^{149}Sm - ^{150}Nd spike solution and dissolved in savillex capsules. Sr, Sm and Nd extraction of whole rock samples followed conventional cation exchange techniques, with Teflon columns containing LN-Spec resin (HDEHP-diethylhexil phosphoric acid supported on PTFE powder). Sr, Sm and Nd samples were loaded on Re evaporation filaments of double filament assemblies, and the isotopic measurements were carried out on a multicollector Finnigan MAT 262 mass spectrometer in static mode. Uncertainties of $^{87}\text{Sr}/^{86}\text{Sr}$ ratios were better than $\pm 0.01\%$ (σ_2) and of Sm/Nd and $^{143}\text{Nd}/^{144}\text{Nd}$ ratios were better than $\pm 0.2\%$ (σ_2) and $\pm 0.005\%$ (σ_2), respectively, according to repeated analyses of international rock standards BHVO-1 and BCR-1. $^{143}\text{Nd}/^{144}\text{Nd}$ ratios were normalized to $^{146}\text{Nd}/^{144}\text{Nd}$ of 0.7219, and the decay constant used was $6.54 \times 10^{-12} \text{ y}^{-1}$. The T_{DM} values were calculated using DePaolo's model (1981).

6. Geochemistry

Major and trace elements analyses were carried out on eleven selected samples (Table 9). In all bulk chemical analyses, the loss on ignition (LOI) values are high, varying from 5.9 to 22.8 %. The analyzed samples show high TiO_2 (up to 4.83 %) and MgO (~ 10 %) in contrast, the variable concentrations of Na_2O (0.09-3.65) and K_2O (0.87-1.9 wt. %) are probably due to subsequent alteration processes.

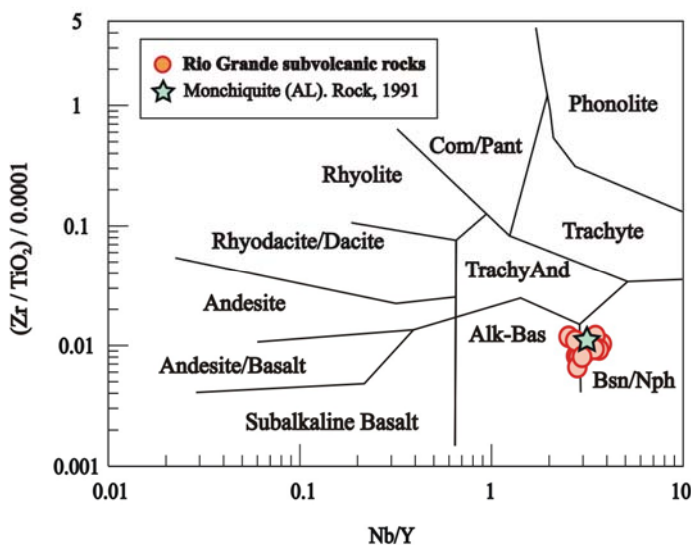


Figure 11: Rock classification diagram using Zr, Ti, Nb and Y immobile elements (Winchester and Floyd, 1977) for the Río Grande rocks.

The high LOI values observed are the result of alteration; although high volatile content is a common feature of lamprophyres (up to 5-6%, Rock, 1991). If we consider that the LOI represent in part magmatic CaCO_2 , the major elements cannot be used in discrimination diagrams; whereas if we assume that the LOI represent CaCO_2 produced by later alteration, recalculation of the major element contents to an anhydrous basis is justifiable. The petrographic and mineral chemistry indicate that some xenoliths and ocelli have dolomite with features that resemble carbonatitic rocks. Therefore the carbonate in the groundmass of these rocks has probably a similar origin, suggesting that the high LOI represent original magmatic volatiles.

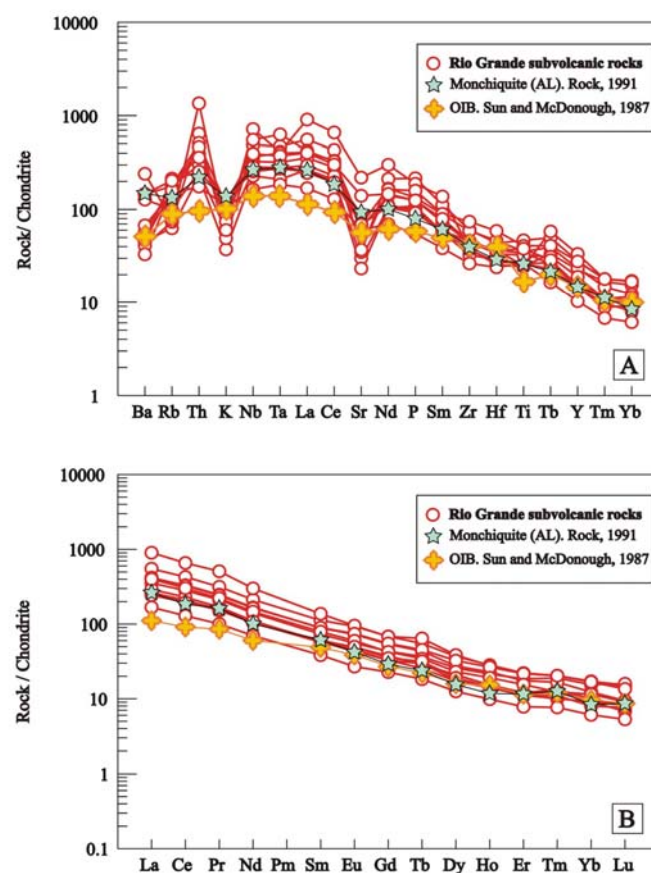


Figure 12: A) Chondrite normalized abundance spidergrams (Thompson 1982) and B) REE normalized diagram after Nakamura 1974, for the Rio Grande rocks. For comparison monchiquite-alkaline lamprophyres after Rock (1991) and average ocean island basalt (OIB) after Sun and McDonough (1989) are shown.

High contents in compatible trace elements (164-253.3 ppm for Cr, 152.4-423 ppm for Ni, 17-24 ppm for Sc, 44.4-94.8 ppm for Co and 202.2-400 ppm for V); indicate moderate degrees of fractionation (Table 9). The wide range in some incompatible trace elements (229-1666 ppm Ba, 356-2590 ppm Sr and 22-75 ppm Rb, 57-251 ppm Nb) may also be due to alteration, although similarly high concentrations and wide variations of LIL elements have been described for several lamprophyres (e.g. Rock, 1991, Azambre et al., 1992, Szabo et al., 1993, Dostal, Owen, 1998, Nedli and Tóth, 2007). Sample P8 displays the highest contents in Ba, Sr and Nb elements.

The Nb/Y vs. Zr/TiO₂ diagram (Fig. 11) show that the R o Grande rocks are transitional between alkali basalts and Basanites/Nephelinites. For comparison, a monchiquite (alkaline lamprophyre) sample is also shown (Rock, 1991). The analyzed samples normalized to the chondrite values from Thompson, (1982) (Fig. 12 A) show enrichments in LILE and LREE with values from 50 to approximately 800 time greater than the chondrite values, whereas the HREE and HFSE contents are lower. Evident negative K and Sr anomalies together with a small negative Ti anomaly are observed. A monchiquite sample (Rock, 1991) and an ocean island basalt (Sun and McDonough, 1987) are also shown for comparison.

Total REE concentrations range from 376 to 798 ppm (Table 9). Chondrite-normalized trace element patterns (Fig. 12 B) are characterized by strong (LREE) enrichment and distinctive REE fractionation with (La/Yb)_n ratio varying from 23.1 to 36.8 for all samples, except for sample P8, which has a higher (La/Yb)_n ratio (53.08). The (La/Sm)_n and (Gd/Yb)_n values from 4.15 to 5.06 and from 2.79 to 4.36 respectively, indicate that garnet was not an important phase in the source. The lack of Eu anomaly suggests that plagioclase fractionation (either by crystal settling or crustal melting) was not important during the origin and differentiation of the original magma. Normally negative Sr and Eu anomalies indicate plagioclase fractionation, but as in the present case no Eu anomaly is observed, the observed negative Sr anomaly is interpreted to indicate source characteristics that could be related to the presence of residual amphibole and clinopyroxene crystals in the source. The depletion of mantle source in Sr during a previous phase of melt extraction is unlikely.

7. Geochronology

7.1. K-Ar data

Some of the samples investigated in the present study have biotite macrocrysts (up to 2 cm). Biotite was separated from two samples (BYA 126 and P8) for K-Ar geochronology at the Center of Geochronological Research of the University of S o Paulo. The analyses resulted

(Table 10) in K-Ar ages of 190 ± 11 Ma (sample BYA 126) and 163 ± 9 Ma (sample P8). Due to the older age corresponds with the altered sample, we prefer to use the 163 ± 9 Ma as representative age.

Table 10: K-Ar biotite age data for the Río Grande subvolcanic rocks

<i>Sample</i>	Rock type	Mineral	Age (Ma)	Max. Error	Ar ⁴⁰ rad ccSTP/g (*10 ⁻⁶)	Ar ⁴⁰ atm (%)	% K	K error (%)
BYA 126	alkaline- lamprophyres	bt	190	11	44.60	15.01	5.73	3.93
P8	alkaline- lamprophyres	bt	163	9	33.05	17.16	4.99	3.10

7.2. Sr and Nd Isotopic Data

Eight whole rock samples were selected for Nd isotopic analyses and five samples were selected for Sr isotope analyses. Sm and Nd concentration were calculated by isotopic dilution, whereas Rb and Sr contents were taken from ICP-MS analyses. The results are listed in Table 11. Initial Sr and Nd isotopic ratios have been re-calculated to the 163 Ma K-Ar age.

The Río Grande rocks display, in general, moderately high initial $^{87}\text{Sr}/^{86}\text{Sr}$ isotopic ratios, ranging from 0.70724 to 0.70781. Samples P8 and BYA 127, however, present much lower Sr initial ratios, of 0.70377 and 0.70455, respectively. $\epsilon_{\text{Nd}}(t)$ values are positive, ranging between + 1.5 (BHU 131) and + 5.6 (BYA 124a). T_{DM} model ages vary between 0.25-0.64 Ga (Table 11).

Table 11: Rb-Sr and Sm-Nd isotopic data for the Río Grande subvolcanic rocks.

Locality	Sample	Rb (**) (ppm)	Sr (**) (ppm)	⁸⁷ Rb/ ⁸⁶ Sr	(⁸⁷ Sr/ ⁸⁶ Sr) _{meas}	(⁸⁷ Sr/ ⁸⁶ Sr) _t	Sm (ppm)	Nd (ppm)	¹⁴⁷ Sm/ ¹⁴⁴ Nd	(¹⁴³ Nd/ ¹⁴⁴ Nd) _{meas}	(¹⁴³ Nd/ ¹⁴⁴ Nd) _t	ε _{Nd(t)}	T _{DM} (Ga)
Coraya	BCO 49	22.00	356.00	0.01	0.70756 (±02)	0.70715	13.38	66.10	0.12	0.512752 (±06)	0.512621	3.77	0.51
	BCO 52						25.38	136.21	0.11	0.512753 (±06)	0.512633	4.00	0.46
	P8*	8.87*	825.60*	0.03*	*0.70373 (±02)	0.70377	25.07*	159.24*	0.09*	0.512635 (±05)*	0.512533	2.06*	0.54
	BYA 123						12.73	74.40	0.10	0.512752 (±06)	0.512642	4.17	0.42
Yacoraite	BYA 124a	52.10	712.30	0.36	0.70829 (±02)	0.70781	7.94	41.21	0.12	0.512716 (±11)	0.512716	5.62	0.25
	BYA 125	57.10	828.40	0.54	0.70783 (±01)	0.70737	18.24	100.84	0.11	0.512747 (±16)	0.512630	3.95	0.45
	BYA 126						18.23	100.73	0.11	0.512743 (±15)	0.512626	3.87	0.46
	BYA 127	71.20	1657.40	0.31	0.70484 (±01)	0.70455	15.78	90.19	0.11	0.512735 (±07)	0.512622	3.79	0.46
Huichaira	BHU 131	35.20	812.40	0.13	0.70753 (±03)	0.70724	13.21	71.68	0.11	0.512625 (±06)	0.512506	1.52	0.64

Obs:

Errors on measured isotope ratios are given in parenthesis as uncertainty in last digits (2s)

ε_{Nd}(T) calculated relative to CHUR with present-day value of ¹⁴³Nd/¹⁴⁴Nd= 0.512638. T_{DM} values were calculated using De Paolo' model (1981).

⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd and ε_{Nd(t)} calculate with an age of 163 Ma.

(*) Data from *Omarini et al., 2005*.

(**) Elemental abundances of Rb and Sr are taken from ICP-MS analyses

8. Discussion and Conclusions

The Río Grande subvolcanic rocks have been classified previously as ankaramites, piroxenitic hornblendites, picrites or basalts with alkaline affinities (Toselli and Aceñolaza, 1984, Manca et al., 1987 and Chayle et al., 1987). The field relationships and the new petrochemical, geochemical, and isotope data, permitted us to classify these rocks as alkaline lamprophyres. The presence of (i) porphyritic panidiomorphic textures, (ii) leucocratic carbonate-analcime rich ocelli and (iii) volatile mineralogical compositions (amphibole, phlogopite) are important characteristics for distinguish lamprophyres rocks from others basic rocks and these characteristics are present in the Rio Grande subvolcanic rocks. According to Rock (1991), the alkaline lamprophyres (AL) are volatile ($H_2O \pm CO_2$)-rich rocks, characterized by high LOI values. On the other hand there is a complete gradation (Rock, 1991) between them and alkali basalts, one can consider AL as volatile-enriched equivalents of alkali basalts (Wimmenauer 1973b).

We plotted the pyroxene and amphibole crystals compositions of the Rio Grande lamprophyres in the classification diagrams proposed by Rock, 1987 for lamprophyres (Fig. 13). In the pyroxene diagram (Fig. 13 A), the diopside and augite crystals fall in the Alkaline lamprophyres (AL) field, close to straddling the boundary separating Calc-alkaline (CAL) and Alkaline lamprophyres (AL) from Ultramafic lamprophyres (UML). In the amphibole diagram (Al_2O_3/TiO_2 vs. CaO/Na_2O) the titanian-pargasite crystals (Fig. 13 B) are plotted at the intersection of Ultramafic lamprophyres-Alkaline lamprophyres with Calcio-alkaline lamprophyres (UML-AL from CAL lamprophyres) fields.

Among the alkaline lamprophyres three major varieties, camptonites, monchiquites and sannaites, are distinguished. On the bases of paragenesis, the Río Grande rocks can be classified as monchiquite, a feldspar-free alkaline lamprophyre composed of diopside, Ti-rich biotite/phlogopite, Ti-pargasite, forsteritic olivine phenocrysts in a groundmass made of the same minerals except olivine, with analcime or glass and abundant globular structures. Since analcime is an important phase in the ocelli, the rock is identified here as *ocellar analcime monchiquite*.

The origin of globular structures is much debated (Rock, 1991). They are believed to represent either: (i) magmatic crystallization involving liquid immiscibility (Philpotts 1976, Eby, 1980) or (ii) segregations of late stage liquids (Cooper, 1979) or amygdales. Those originated by liquid immiscibility have high TiO_2 contents (Philpotts, 1976) and are characterized by distinctive chemistry and different rim compositions between the minerals of globules and groundmass (Foley, 1984), whereas those originated by late stage segregations

have low TiO_2 contents (Philpotts, 1976) and are zoned, with inner analcime-carbonate zone and outer ground-mass mineral zones (Foley, 1984). The similar chemistry between the ocelli and groundmass minerals and the lack of early differentiates (as olivine, clinopyroxene, etc.) within the structures suggest that the ocelli in the Río Grande lamprophyres can be explained by late stage-segregations.

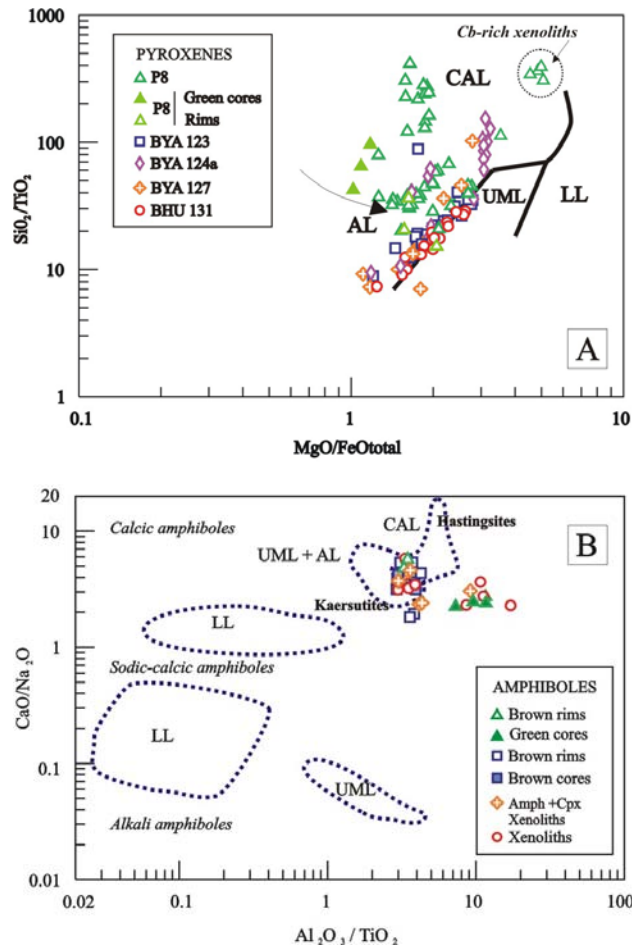


Figure 13: A) Discriminate diagram showing compositional field on pyroxene crystals to the different lamprophyres branches. The Rio Grande pyroxene crystals are plotted in the limit between AL/CAL-UML fields. B) $\text{Al}_2\text{O}_3/\text{TiO}_2$ vs. $\text{CaO}/\text{Na}_2\text{O}$ discriminate diagram for amphibole compositions, after Rock, 1987, 1991. The Rio Grande amphibole crystals are plotted in the UML-AL/CAL limits fields. **CAL:** calc-alkaline lamprophyres, **AL:** alkaline lamprophyres, **UML:** ultramafic lamprophyres and **LL:** lamproites.

The 163 Ma K-Ar ages allow to associate the Río Grande monchiquites with the pre-rift stage of the Salta rift (Jurassic-Eocene) evolution. They were placed in the centre of the future rift structure. For the pre-rift phase the older rocks known were emplaced at ca. 154 Ma (Cristiani et al. 2004, Cristiani et al., 2008) as anorogenic plutonic complexes (Castro Tolay-

Tusaquillas, Abra-Laite-Aguilar and Fundiciones, plutonic complexes) along the Cobres-Salinas Grandes and Tres Cruces branches. The basic magmas of Castro Tolay and Aguilar intrusives have OIB affinity indicating that they were generated in the asthenospheric mantle (Cristiani et al., 2005). At 100 Ma, during the syn-rift phase, rocks with basanite (Matancillas, Finca El Rodeo and Las Conchas, Lucassen et al., 2007) and phonotephrite (Cadillal) compositions have been recognized. For comparison rocks from the different stages of the Salta Rift are plotted in the spiderdiagram of figure 14. The Rio Grande rocks are characterized by the highest contents of LILE, HFSE and REE and by a pronounced K and Sr negative anomaly.

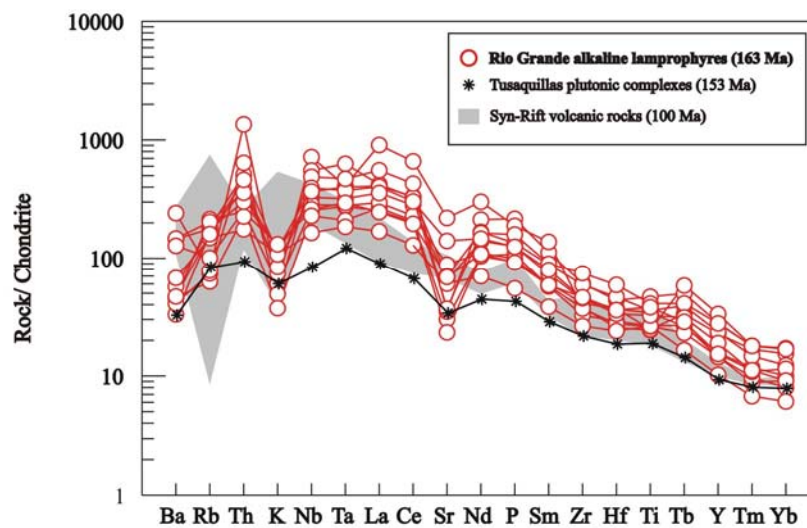


Figure 14: Trace elements pattern of pré-rift stage and syn-rift stage rocks of Salta Rift. The Rio Grande lamprophyres and Tusaquillas plutonic complex belong to pre-rift stage. Note the high concentrations in LILE and HFSE of the Rio Grande lamprophyres compared with the Tusaquillas complex. The syn rift volcanic rocks selected are Las Conchas, El Cadillal, Finca El Rodeo and Matancillas basanites. Date of Tusaquillas complex, from Cristiani et al., 2004. Data of Syn-Rift volcanic rocks from Lucassen et al., 2007. Normalized values from Thompson, 1982.

Lucassen et al., (2005) propose that the lithospheric mantle composition under the Central Andes has been modified periodically by subduction processes since the Paleozoic. Therefore, according to these authors, at present day there are three possible mantle sources under the Central Andes: (i) an early Paleozoic mantle, which formed, or was modified, during the late Cambrian early Ordovician Famatinian orogeny, (ii) a Proterozoic subcontinental lithospheric mantle (SCLM), belonging to the Brazilian shield, which is thought to be progressively underthrust beneath the present Central Andes during Miocene shortening

(e.g., McQuarrie et al., 2005) and (iii) mantle modified by magmatic activity during Mesozoic extension (e.g., Viramonte et al., 1999; Sempere et al., 2002).

In order to constrain isotopically their mantle source(s), the Río Grande monchiquites are plotted in the $^{87}\text{Sr}/^{86}\text{Sr}$ vs. $^{143}\text{Nd}/^{144}\text{Nd}$ isotope diagram (Fig. 15). The samples plot within the high $^{87}\text{Sr}/^{86}\text{Sr}$ peridotite field and close to two Sr-enriched basanite shown by Lucassen et al. (2005). In order to explain the presence of high $^{87}\text{Sr}/^{86}\text{Sr}$ peridotite and the Sr-enriched basanite in this area, these authors propose, the separation of Rb-bearing hydrous minerals, such as phlogopite and amphibole, or the long-term separation of the respective parent/daughter element pairs in the pre-cretaceous metasomatized mantle. The isotopic data point to a metasomatized mantle as a source also for the Río Grande lamprophyres.

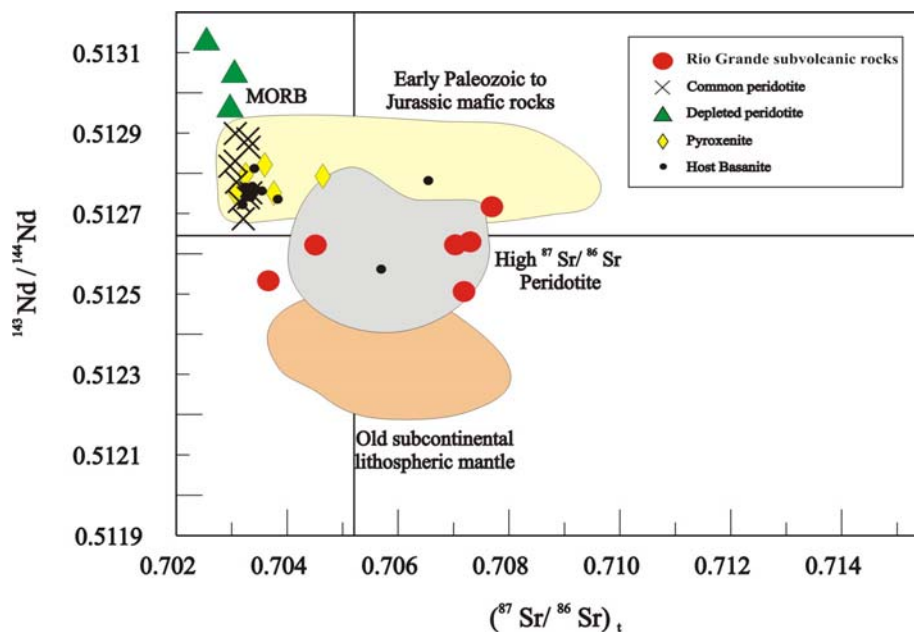


Figure 15: $^{87}\text{Sr}/^{86}\text{Sr}$ vs. $^{143}\text{Nd}/^{144}\text{Nd}$ diagram for the Río Grande lamprophyres. The isotopic compositions of xenoliths and mantle-derived rocks from the Central Andes are also shown. The data are recalculated to 160 Ma (Lucassen et al., 2005).

The Nd and Sr isotopic characteristics also suggest a heterogeneous mantle source, with one end-member formed by high $^{87}\text{Sr}/^{86}\text{Sr}$ ratios and other which show low values for this ratio.

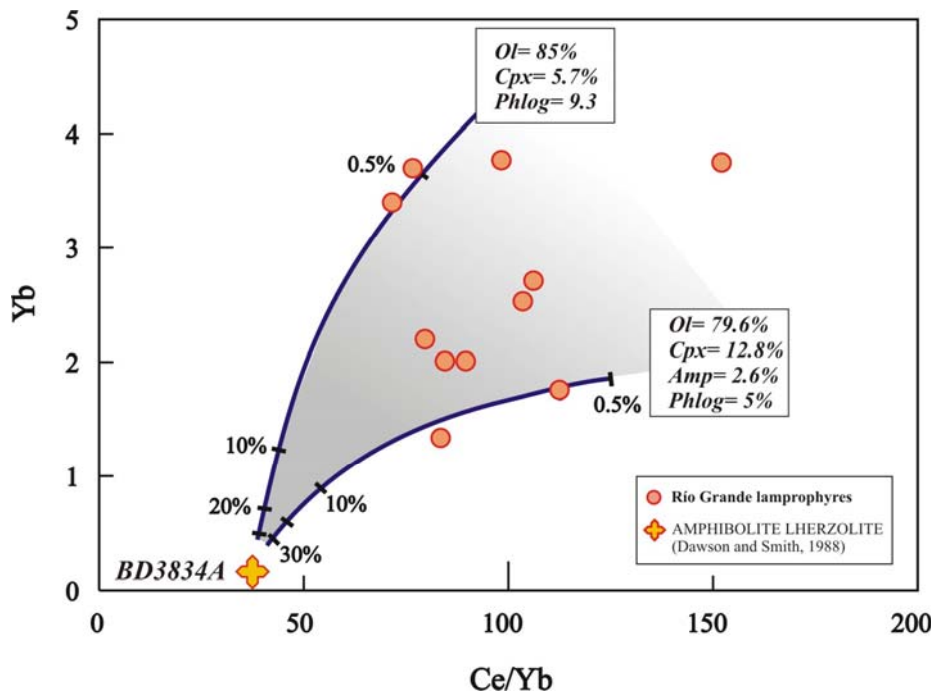


Figure 16: Ce/Yb vs. Ce diagram for the Rio Grande lamprophyres. Melting curves for amphibolite lherzolite of sample BD3438A from a metasomatised mantle of Pello Hill, Tanzania, are shown. Degree of partial melting (%) is indicated along the melting lines.

In figure 16 partial melting modeling of a metasomatized mantle is shown. Modal back-melting calculation have been carried out on a metasomatized mantle. We select an amphibolite lherzolite (BD3848A) from the Pello Hill in Tanzania (Dawson and Smith, 1988). It is characterized by 6.0 ppm of Ce and 0.16 ppm of Yb. For the clinopyroxene and phlogopite we used the partition coefficient proposed by Foley, 1996 for alkaline lamprophyres. For the rest of minerals we used the partition coefficient shown in Rollinson, 1993 for basaltic liquids. Partial melting trajectories suggest that the Rio Grande lamprophyres could be generated from an enriched amphibolite lherzolite mantle source by less than 2 % partial melting (Fig. 16). The different lamprophyres were formed by nearly the same degree of melting of the same source with slightly different parageneses. The BYA123 and 124, are explained by a partial melting of a amphibolite lherzolite with Ol (79.6%), Cpx (12.8%), Amp (2.6%) and Phlo (5%), whereas samples BYA123 and 122 could be explained by a partial melting around 0.5% of a amphibolite lherzolite with Ol (85%), Cpx (5.7%) and Phlo (9.3%). The rest of samples could be modeled using a source with intermediate mineral compositions.

Several types of metasomatism have been identified in mantle xenoliths around the world on the basis of chemical, isotopic and mineralogical compositions. Proposed metasomatic agents include alkali-rich silicate and $\text{CO}_2 \pm \text{H}_2\text{O}$ -rich fluids or carbonatite melts (e.g. Frey and Green, 1974; Roden et al., 1984, Coltorti et al., 1999). The REE patterns of the studied rocks show strong enrichment in LREE, while HREE are depleted. The presumed introduction of a CO_2 rich fluid in the depleted mantle source might have increased LREE with respect to HREE (Barbieri et al., 1997). The anomalous SrCO_3 content of carbonate (e.g. ocelli, cpx-bearing carbonate xenoliths and groundmass in sample P8) (Table 8), may suggest the involvement of carbonatitic liquids during the evolution of the Río Grande monchiquites and the presence of apatite may also be suggestive of metasomatism by carbonatite melts or CO_2 rich fluids. The hypothesis that the mantle where the Rio Grande lamprophyres derived is metasomatized by carbonatite fluids is also supported by the occurrence of carbonatite rocks associated with the Salta rift in NW Argentina such as the Santa Julia Complex (Zappettini, 1998) and the Ayopaya carbonatite complex (Schultz et al., 2004, Lucassen et al., 2007)

Considering that the Río Grande lamprophyre original magmas derived from to the metasomatized mantle, their T_{DM} model ages might record a metasomatic event at approximately 0.45 Ga, coincident with the Ordovician Famatinian Orogeny (Lucassen et al., 2005).

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CAPÍTULO III

Este Capítulo apresenta os resultados petrológicos, geoquímicos e isotópicos, referentes às rochas subvulcânicas da Serra de El Niño Muerto e do Vale de Río Blanco. As informações serão apresentadas na forma de artigo completo, que será submetido ao periódico *Journal of South American Earth Sciences*, com o título:

LA-ICPMS U-PB GEOCHRONOLOGY AND PETROLOGY OF VOLCANIC ROCKS OF THE NEOPROTEROZOIC TO LOWER PALEOZOIC BASEMENT OF THE CENTRAL ANDES, NW ARGENTINA: INSIGHTS ON THE EVOLUTION OF WESTERN GONDWANA.”

LA-ICPMS U-PB GEOCHRONOLOGY AND PETROLOGY OF VOLCANIC ROCKS OF THE NEOPROTEROZOIC TO LOWER PALEOZOIC BASEMENT OF THE CENTRAL ANDES, NW ARGENTINA: INSIGHTS ON THE EVOLUTION OF WESTERN GONDWANA.”

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Abstract

In situ LA-ICPMS U-Pb zircon analyses and Sr-Nd isotopic data combined with major, trace and rare earth elements (REE) contents for the El Niño Muerto and Rio Blanco volcanic rocks of the Neoproterozoic to Lower Paleozoic basement of the Central Andes in NW Argentina are discussed. They are associated with sediments which have been traditionally attributed to the Puncoviscana Formation. The new data helped to investigate the tectonic setting of emplacement of the original magmas, and contribute to the better understanding of the tectonic evolution of Western Gondwana during the Early Paleozoic.

The El Niño Muerto metadacites have high-K calc-alkaline nature characterized by intermediate SiO₂ contents, high Al₂O₃ and low MgO, Cr and Ni. They are peraluminous with up to 5 % normative corundum. Contents of large ion lithophile elements (LILE) are high and high field strength element (HFSE) contents are comparatively low, typical of arc rocks. Σ REE abundances vary from 186 to 206 ppm and chondrite-normalized REE patterns are moderately fractionated, with (La/Yb)_n ratios between 5.9 to 8.6. The Rio Blanco basalts have geochemical characteristics which are similar to E-MORB, with Zr/Nb ratios from 6.8 to 9.5. Contents

of large ion lithophile elements (LILE) are variable (10-100 ppm) and high field strength element (HFSE) are low (8-50 ppm). Σ REE vary from 76 to 196 ppm and chondrite-normalized REE patterns are weakly fractionated, with (La/Yb)_n ratios between 2.6 to 4.5. Both rock units were dated using laser ablation ICPMS on zircon crystals. The data yielded the ages of 495 ± 4 Ma for the El Niño Muerto metadacites and of 501 ± 9 Ma for the Rio Blanco basalts. The El Niño Muerto metadacites have high initial ⁸⁷Sr/⁸⁶Sr (0.71107-0.71180) and low initial ¹⁴³Nd/¹⁴⁴Nd (0.511506-0.511697) corresponding to negative ε_{Nd}(t) values, between -9.7 to -5.9, and T_{DM} model ages between 1.62-1.79 Ga. The Rio Blanco basalts have moderately high ⁸⁷Sr/⁸⁶Sr initial ratios (0.70713-0.70942), positive ε_{Nd}(t) values (+2.47 to +4.46) and T_{DM} model ages between 0.84 to 1.12 Ga.

The geochemical and isotopic data presented in this study indicate that the El Niño Muerto metadacites formed in the Famatinian calc-alkaline magmatic arc and that the original magmas are strongly contaminated with Precambrian (Mesoproterozoic) continental crust. The tectonic setting of formation of the Rio Blanco basalts remain unclear and their geochemical characteristics indicate that they may correspond either to an E-MORB or to within-plate anorogenic basalts. Trace element contents show compositions which are intermediate between typical Pacific E-MORB and OIB. The geochemical characteristics and geological setting of the Rio Blanco

basalts suggest that they may also be interpreted as a remnant of a back-arc basin developed during the Late Cambrian /Early Ordovician in NW Argentina.

1. Introduction

The rock units comprising the basement of the Central Andes, between 22° and 33° S record at least two Paleozoic orogenic episodes: the Pampean, at ca. 520 Ma (Aceñolaza and Toselli 1973) and the Famatinian, at ca. 460 Ma (Aceñolaza and Toselli 1973).

The pampean basement in NW Argentina includes mainly Neoproterozoic to Lower Paleozoic metasedimentary sequences; the Puncoviscana Formation (Turner 1960, Omarini, 1983, Omarini, 1983, Aceñolaza et al., 1988, Jezek, 1985, 1986, Willner et al., 1990, Do Campo and Guevara, 2005, Zimmermann, 2005) is the main and oldest representative of the basement in the Eastern Cordillera. It is characterized by a thick sequence (> 2000 m) of very low to low grade metapelite and turbidites with few intercalations of carbonates and volcanic rocks (Jezek 1990, Omarini et al., 1999). Do Campo et al., (1998), Do Campo, (1999b) and Do Campo et al., (1999), on the basis of K-Ar analyses on clay minerals, defined a Late Neoproterozoic age for the Puncoviscana Formation. Its ichnofauna, whose chronostratigraphic significance has been recently reinterpreted (Buatois and Mángano, 2003), indicates a Vendian-Nemakitian-Daldynian age.

Volcanic rocks of uncertain age intercalated with the Puncoviscana Formation in NW Argentina have been recognized in several localities of the Eastern Cordillera and easternmost Puna (Toselli and Aceñolaza, 1984; Coira et al., 1990, Manca et al., 1987, Chayle et al., 1987; Omarini y Alonso, 1987, Omarini et al., 1999, Omarini et al., 2005). Due to the lack of detailed field data as well as accurate geochronology, little is known about their tectonic setting of emplacement. Two of these volcanic rocks occurrences are the subject of this study.

In this paper we present new field, petrographic and geochemical data for volcanic rocks intercalated in the Puncoviscana Formation, exposed in the El Niño Muerto Hill and in the Rio Blanco Valley. We also show, for the first time, in situ LA-ICPMS U-Pb zircon ages and Sr-Nd isotopic data for these rocks. This study attempts to investigate the relationships between the volcanic rocks and the Paleozoic orogenic cycles, to date the sedimentary sequences which are associated with them, as well as to identify their original tectonic setting. The new results are discussed, in combination with data available from the literature, in order to clarify the geodynamic significance of these volcanic rocks in the tectonic evolution of the western margin of Gondwana in NW Argentina during the Late Neoproterozoic and Early Paleozoic.

2. Geological Setting

The basement of NW Argentina records at least two main orogenic cycles between the break-up of the Rodinia supercontinent (~750 Ma) and the continental amalgamation of Pangea (Upper Carboniferous): the Pampean and the Famatinian orogenies.

The *Pampean Cycle* (Aceñolaza and Toselli 1973), occurred between ca. 550 and 520 Ma (Rapela, et al., 2001a, Sollner et al., 2000b, Ramos and Basei, 1997). It has been documented in the Eastern Cordillera (EC), Puna and more extensively in northwestern Sierras Pampeanas Orientales (Aceñolaza et al., 1990, Toselli, 1990, Rapela, et al., 1992, Kraemer et al., 1995,

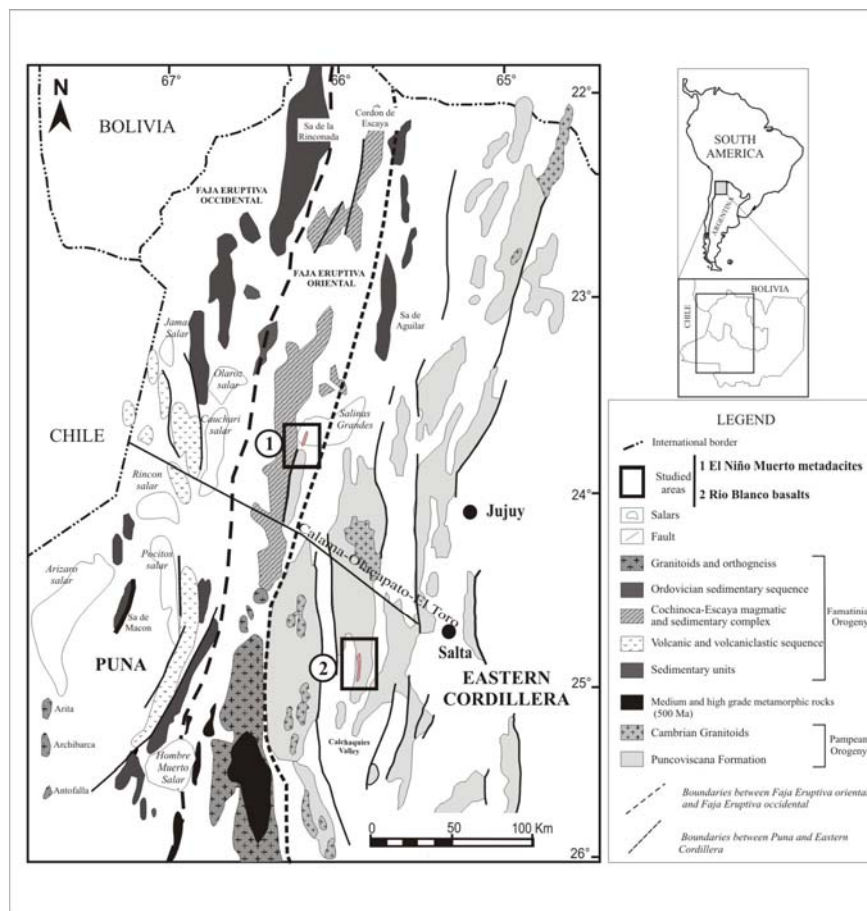


Figure. 1: Geological map showing the distribution of Neoproterozoic to lower Paleozoic basement rock units of NW Argentina (modified after Do Campo and Guevara, 2005 and Viramonte et al., 2007).

Llambías et al., 1998). It was characterized by an early subduction-related magmatic arc followed by the Early Cambrian (535-520 Ma) accretion of an exotic terrane, identified either as the Pampia block (Rapela et al., 2001a) or as the Arequipa-Belen-Antofalla Craton (Ramos 1988). After this accretion, the protomargin of Gondwana became a passive margin between 520 and 499 Ma (Keppie and Bahlburg, 1999, Rapela et al., 2001) until the Ordovician, when the Famatinian orogeny took place.

In NW Argentina (EC and Puna), the Pampean Cycle is represented by a passive margin sequence composed of a low- to very low-grade metamorphic sedimentary succession (Hongn et al., 1985, Mon and Hongn, 1991) made of thick turbiditic deposits with scarce associated limestones and volcanic rocks (Jezek 1990, Omarini et al., 1999).

This succession named by Turner (1960) as the Puncoviscana Formation has Neoproterozoic to lower Cambrian age, according to paleontological and geochronological data (Aceñolaza et al., 1982, Kumpa & Sanchez, 1988, Durand and Aceñolaza, 1990, Adams et al., 1990, Lork et al., 1990, Rossi et al., 1992, Bordonaro 1992, Durand 1996, Do Campo et al., 1999, Do Campo, 1999b, Buatois and Mángano 2003, Adams et al., 2006). Several volcanic rocks of uncertain age are interbedded in the sequence in different localities.

The *Famatinian Cycle* (Aceñolaza and Toselli 1973) took place between 490 to 400 Ma and records the subduction of oceanic lithosphere along the western margin of Gondwana during the Early Paleozoic. The tectonic reconstructions are supported by interpretation of the sedimentary faunal record, the tectonic-magmatic evolution of the basins, and paleomagnetic evidence (Aceñolaza and Toselli 1973, Dalziel, 1997, Rapela et al., 1992, Kraemer et al., 1995, Llambías et al., 1998, Rapela et al., 2001b). The final stage of this cycle was marked by the accretion of an exotic terrane, the Precordillera.

In NW Argentina, the *Famatinian Cycle* (490-400 Ma) is represented by: (i) platformal siliciclastic, shallow-marine sedimentary succession in the Eastern Cordillera and deeper marine sediments to the west, in the Puna plateau (Turner 1960, Salfity et al., 1975, Gohrbrandt, 1992, Moya et al., 1994, Balbhurg 1998), and (ii) basic and acid lavas and associated volcanoclastic turbidites exposed in the eastern Puna extending southwards to the Famatina magmatic arc. Two parallel N-S trending ordovician arcs are recognized in the Puna. The eastern arc known as “Faja Eruptiva de la Puna Oriental” (Mendez et al., 1973), and the western arc known as “Faja Eruptiva de la Puna Occidental” (Palma et al., 1986) (Fig. 1). The eastern arc is characterized by the alignment of small intrusion of Al-rich trondhjemites, tonalites and granodiorites as well as metadacites and metarhyolites (Omarini et al., 1984, Viramonte et al., 1984, Lorck and Balbhurg 1993, Palma et al., 1986, Blasco et al., 1996, Koukharsky et al., 2002, Kleine et al., 2004, Viramonte et al., 2007), with ages ranging from 499 to 478 Ma (Rapela et al., 2001a). The main features of the western arc is the presence of Al-poor granodioritic and monzogranitic cordilleran I-type bodies emplaced between 485 and 465 Ma (Rapela et al., 1998b, Pankhurst et al., 2000). At 460 Ma, the collision of an exotic terrane with Laurentian affinities, the Precordillera terrane, marks the final stage of this orogenic cycle (Rapela et al., 2001a).

Several studies recognized intercalations of volcanic rocks in the Neoproterozoic sedimentary successions of NW Argentina in three main localities: the Rio Blanco Valley, the El Niño Muerto Hill and the Rio Grande Valley (Toselli and Aceñolaza, 1984; Chayle and Coira, 1987; Manca et al., 1987; Do Campo and Riveiro Guevara 2005; Omarini and Alonso, 1987). In this paper we focus on the El Niño Muerto Hill and the Rio Blanco valley outcrops (Fig. 1).

The El Niño Muerto Hill is located at the eastern portion of the Puna, close to the limit with the Eastern Cordillera, between 23° 46' 34.73" and 23° 44' 07.68" S and 66° 14' 50.52" and 66° 12' 48.88" W (Fig. 1), at the southern margin of Salinas Grandes salar. A previous study of the El Niño Muerto Hill investigated the geochemistry of Puncoviscana Formation (Do Campo and Riveiro Guevara 2005) but did not focus on the associated subvolcanic rocks. The authors have recognized a conspicuous slaty cleavage (s1) in both the volcanic and sedimentary rocks (Fig. 2). Kubler index (0.36; 0.24° 20) values for clay mineral as well as TEM (Transmission Electron Microscopy) and SEM (Scanner Electron Microscopy) data indicate a high-anchizonal to epizonal grade for the metamorphism of the Puncoviscana slates in that locality (Do Campo et al., 1998; Do Campo et al., 1999, Do Campo, 1999a, b, Do Campo and Nieto, 2003).

The Rio Blanco Valley is located in the Eastern Cordillera, approximately 60 Km west of Salta, between 24° 58' 08.73" and 24° 57' 05.43" S and from 65° 58' 03.41" to 65° 57' 35.71" W. In this area, submarine basaltic lavas and hialoclastites are associated with the Puncoviscana red pelites and turbidites (Omarini and Alonso, 1987) (Fig. 3).

2.1 Field relationships

2.1.1. El Niño Muerto Hill (ENM)

The ENM consists of a NE-SW hill formed entirely by the Puncoviscana metasedimentary sequence. Acid subvolcanic rocks constitute a sill intruded into the sequence (Fig. 2). The Puncoviscana Formation consists of fine-grained turbidites with alternating centimetric greywacke and pelite beds (Fig. 2 A) which commonly display cross-bedding and parallel stratification.

A felsic subvolcanic body forms a large (~ 6 kilometers long and 30 m thick) NE-SW tabular intrusion, which is intensely foliated in some localities (Fig. 2 A, B, and C). Enclaves of the host metasedimentary rocks, (Fig. 2 D) and thermal aureoles (Fig. 2 C) recognized by bleaching of the metasediments are evidence of the intrusive relationship.

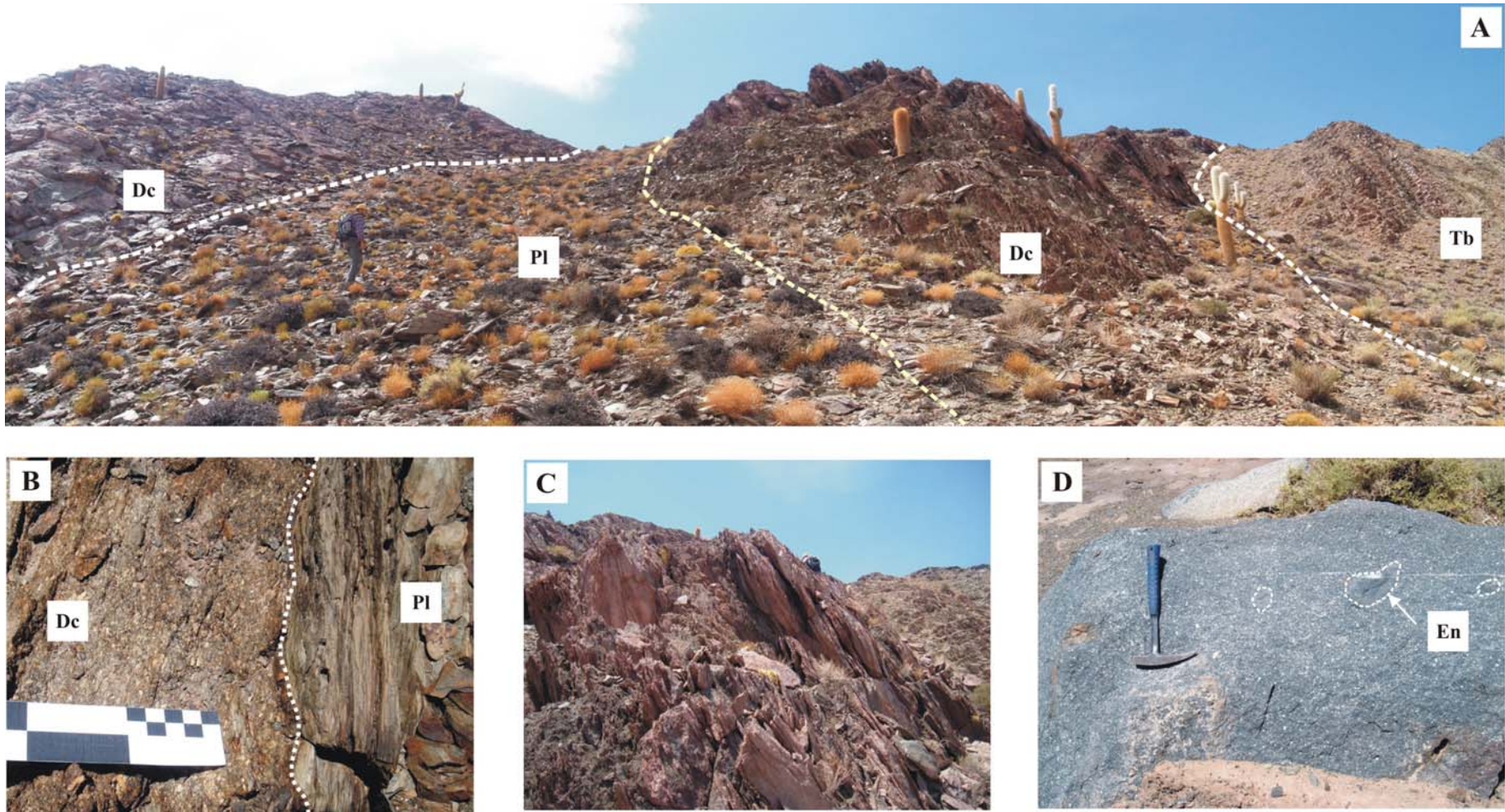


Figure 2: A) Panoramic view of the El Niño Muerto sill (Dc) north sector. The dacites are intruded in the turbidite sequence (Tb) of the Puncoviscana Formatio. B) The contact, is marked by thermometamorphosed and bleached pelites (Pl) indicating the intrusive character. Note that sediment and dacite are intensely foliated. C) View of sill outcrop with prominent foliation. D) Aspect of dacites in fresh cut. They are characterized by porphyritic textures with plagioclase phenocrysts embedded in a gray groundmass of biotite, quartz and plagioclase crystals and abundant pelite enclaves (En). The scale has 25 centimeters long.

The El Niño Muerto metadacites are dark grey, very homogeneous, and fine- to medium-grained. They are characterized by porphyritic textures with abundant plagioclase, K-feldspar and quartz phenocrysts up to 1 cm in size (Fig. 2 D).

2.1.2. Río Blanco Valley

The sedimentary succession in this area consists of a 2000 m-thick turbidite sequence, with a 200 m-thick intercalation of hemipelagic sediments represented by red pelites and minor pink limestone (Fig. 3). Excellent exposures of the Rio Blanco mafic volcanic rocks are associated with the hemipelagic level.

Sedimentary sequence

The turbidite sequence (Tb) is characterized by dominant pale-grey greywacke with minor gray pelites that pass to red pelites (Fig. 3). The greywacke beds exhibit typical sedimentary structures of turbidite, such as flute casts, current waves, cross bedding, erosion surfaces and water-escape structures.

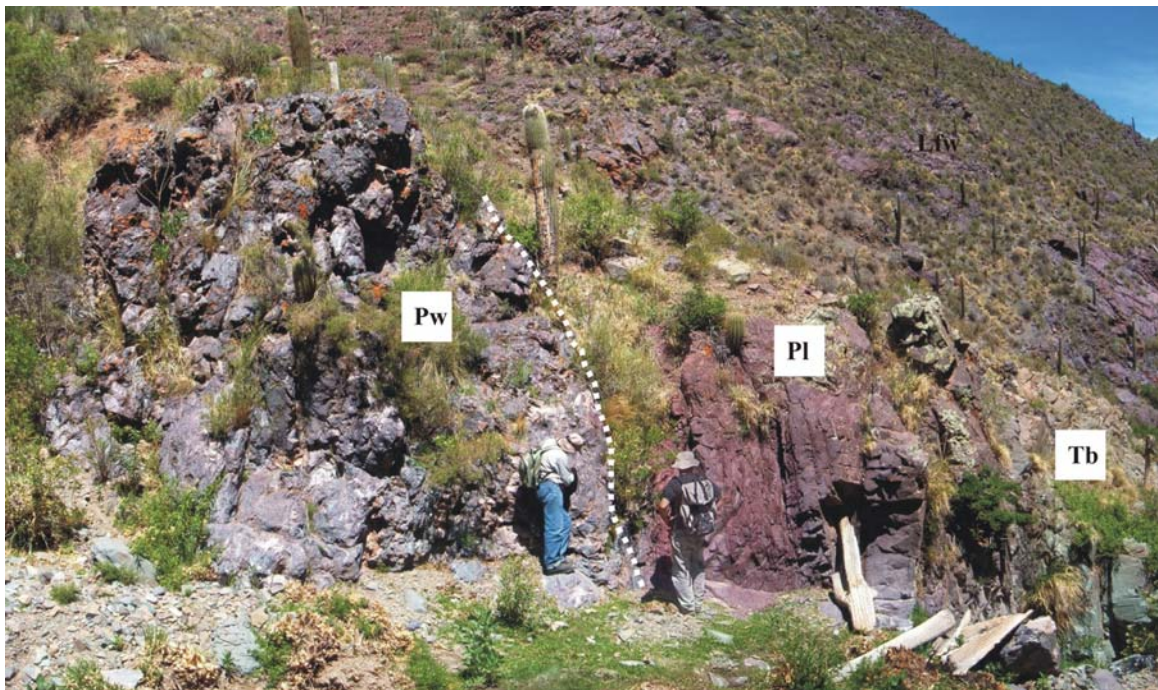


Figure 3: Contact between pillow lavas (Pw), red pelites (PI) and the turbidite sequence (Tb) of the Neoproterozoic to Lower Paleozoic basement in the Rio Blanco Valley.

The turbidite sequence is interrupted by a hemipelagic sequence constituted by yellowish pink limestone (Lm) interbedded with red pelites (Pl). Parallel lamination and slump structures are observed in the pelites.

Río Blanco volcanic rocks

The volcanic rocks are exposed in a tributary valley of the Río Blanco river forming a 5-km long N-S trending structure. They occur as hydrothermally altered basalt interbedded with red pelites and limestones (Fig. 4 A), including both subvolcanic and extrusive bodies. The subvolcanic bodies are characterized by massive to flow banded lavas (SbLv), whereas the extrusive bodies are formed essentially by pillowed lava flows (Pw). Important peperite deposits are associated with the subvolcanic bodies that intrude limestone and red pelites, indicating the synchronicity between the magmatic event and the deposition of these sediments.

Reddish to gray subvolcanic rocks outcrop extensively in the study area (Fig. 4 A) reaching thicknesses of up to 100 meters. They are massive to flow banded bodies characterized by porphyritic textures with variable amounts of plagioclase phenocrysts (up to 6 mm) immersed in an aphanitic greenish to red groundmass with abundant vesicles which have been partially filled with carbonate. Abundant centimetric to decimetric yellow limestone enclaves are found. These enclaves, in the banded lava flows, are strongly deformed parallel to the lava flow displaying lensoidal shape (Fig. 4 C) indicating high velocity and low viscosity of lava flow and suggesting that the limestone was not consolidated at the time of extrusion.

Pillowed lavas are exposed in several outcrops (Fig. 4 A, B) forming beds with variable thicknesses. They are found mainly associated with limestone and red pelites. The pillow structures are commonly tightly packed; when this is not the case, limestone or crushed glass fill the spaces between the pillows. The individual pillows show chilled margins up to 1 cm thick and radial joints up to 10 cm in length, are typically 0.5-1 m across and in their interior plagioclase phenocrysts (up to 3 mm) and carbonate-filled vesicles are common.

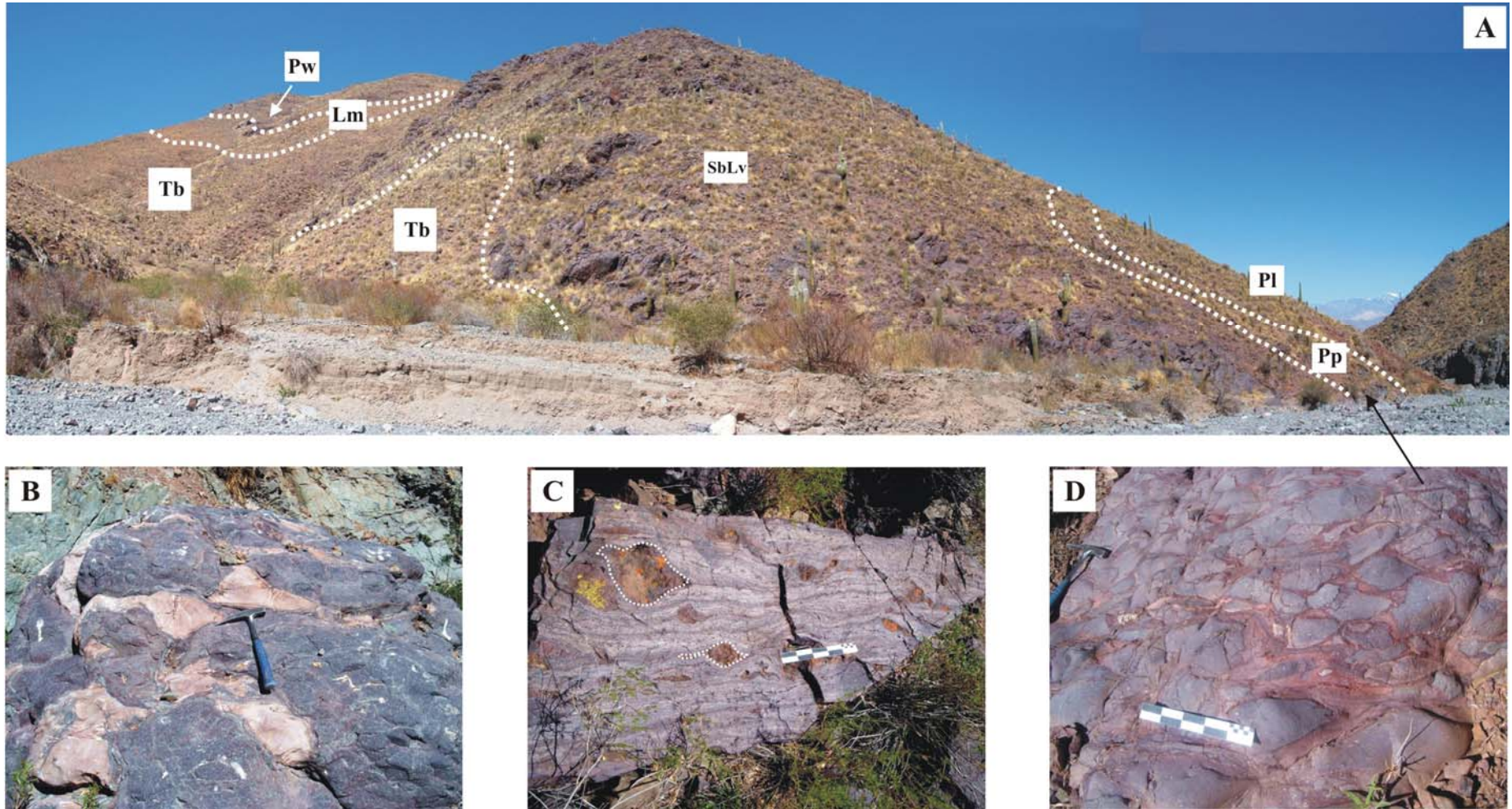


Figure 4: A) Panoramic view of the Rio Blanco volcanic rocks. They are intruded in a hemipelagic? sequence made of red pelite (Pl) and limestone (Lm), that are enclosed by the turbidite sequence (Tb). B) Detail of pillow-like lava (Pw) intruded in yellowish pink limestone (Lm). Note the chilled margin and connection among the individual pillows. C) Aspect of the banded lava flow with abundant lensoidal limestone clasts. D) View of globular peperite deposit formed in the contact between lava flow and red pelite (Pl). The scale is 25 centimeters long.

Peperite deposits

Reddish-gray peperite forming approximately 20 m-thick deposits (Pp) are located where subvolcanic bodies intruded the red pelites (Fig. 4 A, D). Peperite grades from closely packed, (high juvenile clasts/sediment ratios), close to the intrusion contact, to disperse peperites, (low juvenile clasts/sediment ratios) away from the contact with the igneous rocks (Hanson and Wilson, 1993). According to Busby-Spera and White (1987), they were divided into two textural end-members: blocky peperite and globular peperites. The juvenile clasts in both types have compositions and textures similar to the mafic lavas. Angular to pillow-like juvenile clasts characterize the blocky peperite (Fig. 4 D) whereas elongated lobe-like fluidal clasts with complex outlines are the main feature of the globular peperite

3. Petrography

3.1. Subvolcanic rocks of the El Niño Muerto Hill

The El Niño Muerto rocks are characterized by porphyritic textures with approximately 40% phenocrysts. These are quartz (~30%), corundum (~8%), alkali feldspar (~22%), anortite ($\pm 1\%$), albite (~19%), enstatite (~5%) ferrosilite (~4%) and hypersthene (~12%) crystals (Table 2A) set in a fine-grained altered groundmass made of quartz, biotite and sericite/muscovite (Fig. 5). Fe-Ti oxides (~1.6%) and abundant zircon grains are important accessory phases.

The schistose fabric marked by the alternation of quartz- and mica-rich bands is the common feature of the groundmass. Polycrystalline aggregates of quartz grains elongated parallel to the regional cleavage characterize the quartz-rich bands, whereas fine-grained biotite and muscovite crystals elongated along the cleavage characterize the others bands (Fig. 5 A).

Plagioclase (albite-oligoclase An_{10-30}) phenocrysts and microphenocrysts commonly exhibit subhedral outlines and some crystals have disequilibrium features, such as sieve textures and rounded edges (Fig. 5 B). They show incipient alteration to white mica. Subhedral K-feldspar crystals are also observed but not common.

Quartz phenocrysts display rounded outlines and frequent disequilibrium textures such as embayments (Fig. 5 C). Eye-shaped deformation of groundmass fabrics is more common around quartz phenocrysts (Fig. 5 D). Some phenocrysts present poikilitic textures formed by

inclusions of biotite. Quartz in the fine groundmass has anhedral outlines and is strongly elongated.

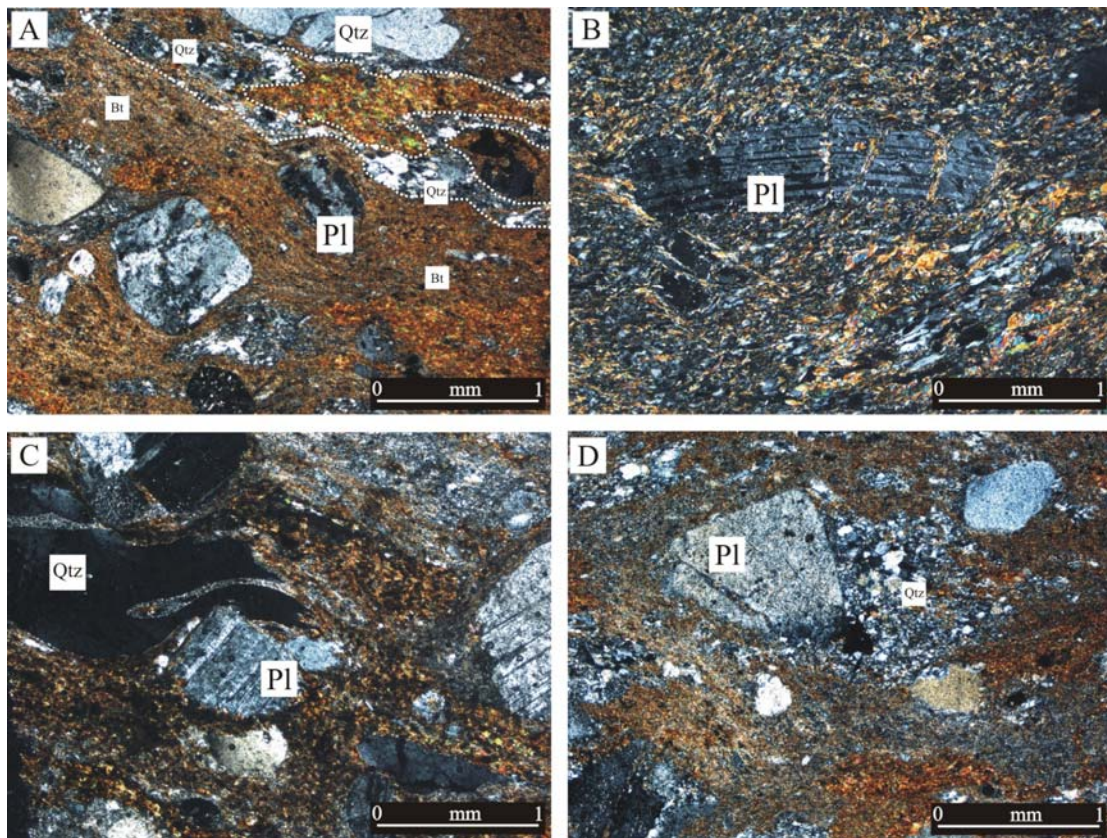


Figure 5: Photomicrographs of representative textures of the El Niño Muerto metadacites. A) Porphyritic metadacite with plagioclase, quartz and K-feldspar phenocrysts. The matrix is made of alternating quartz-rich and biotite-rich bands. B) Plagioclase crystal. The fractures perpendicular to the main stress are filled with mica. C) Quartz crystal showing embayment and D) plagioclase crystals showing pressure shadows. *Abbreviations:* Pl, plagioclase; Qtz, quartz; Kfs, feldspar; Ep, epidote.

Biotite is commonly present either as microphenocrysts or in the groundmass forming the dark-coloured layers. Muscovite is found as very small crystals in the groundmass together with biotite. (Fig. 5 A, C, D). The fine grained sericite is the most common alteration mineral in these rocks; they are found as small crystals in the groundmass as well as inside the plagioclase and sanidine crystals.

3.2. Subvolcanic and Volcanic rocks of the Rio Blanco Valley

Porphyritic to intergranular textures characterize the basaltic subvolcanic bodies and pillowed lava flows of the Rio Blanco Valley. Albite-oligoclase phenocrysts (up to 2 cm

across) and microphenocrysts, are set in a groundmass comprising albite, Fe-Ti oxides (hematite and ilmenite), calcite and quartz (Fig. 6 A). A less common textural type, characterized by large albite phenocrysts partially altered to epidote, is also recognized (Fig. 6 B). Albite phenocrysts (Fig. 6 C) and microphenocrysts display subhedral outlines and commonly show sieve textures. The groundmass albite is generally more altered, compared with the phenocrysts. Their major alteration products include sericite, calcite and epidote. Calcite commonly fills vesicles and also replaces plagioclase. Epidote recognized in samples BRB 64 and 65 (Fig. 6 B) occur as alteration products of calcic plagioclase, whereas chlorite constitutes a minor phase filling vesicles together with carbonate and hematite. Distinguishing features of the Rio Blanco volcanic rocks are the abundant vesicles (Fig. 6 D), present in different sizes and shapes being normally filled with hematite, calcite and minor chlorite

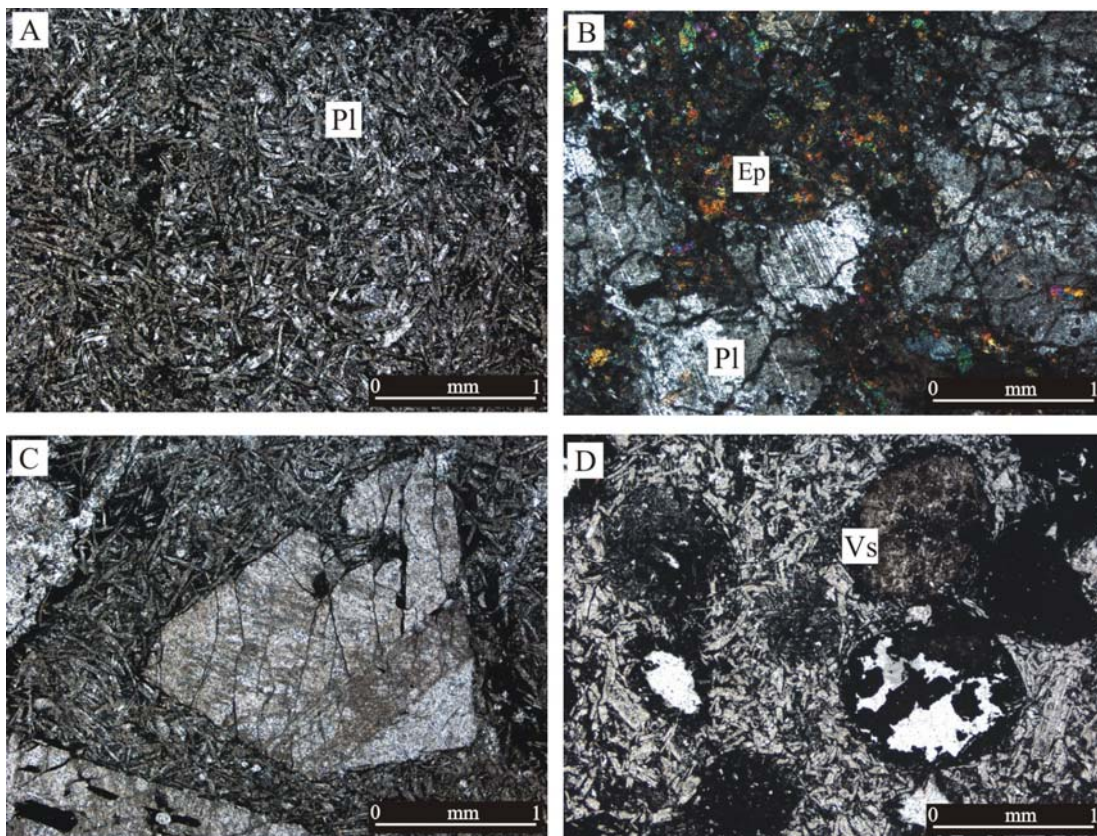


Figure 6: Photomicrographs of the Rio Blanco volcanic rocks. A) A plagioclase laths with intergranular opaque oxides is the most common textural variety of these basaltic rocks B) A less common textural variety is characterized by plagioclase phenocrysts (up to 2 mm) partially altered to epidote. C) Example of porphyritic intergranular texture. Albite phenocrysts are set in a groundmass composed totally by microcrystals of the same composition. D) Carbonate- and hematite-filled vesicles. *Abbreviations:* Pl, plagioclase; Ep, epidote; Vs, vesicles.

4. Analytical Procedures

Samples selected for geochemical analyses were washed and had weathered and veined portions cut off. The rocks were crushed and milled in a videa mill to a very fine powder. Major elements and Ni and Sc (ppm) were analyzed by ICP-OES and trace element analyses were carried out by ICP-MS at the ACME Laboratories in Canada.

Zircon separation and analyses were carried out at the Geochronology Laboratory of the Universidade de Brasilia. Heavy mineral concentrates were obtained using conventional gravimetric and magnetic techniques (panning, Densitest^R separator and Franz isodynamic separator). Final purification was achieved by hand picking using a binocular microscope and selected zircon grains were mounted on a polished epoxy section. Two parts of Resin (Struer, Epo Fix Resin) and one part of Hardener (Struer, Epo Fix Hardener) were poured into the cast with zircons inside and let to rest for twelve hours. The grain mounts containing the samples and standards were polished and cleaned in HNO₃ 3% during five minutes immediately prior to analysis to remove surface Pb contamination. Finally they were dried and mounted together with the standard in a Laser Ablation cell. The operating data are listed in Table 1. During analytical session GJ-1 standard zircon was used for correcting isotopic ratios using a standard-sample-standard “bracketing” technique.

Sm-Nd and Rb-Sr isotopic analyses followed the method described in Gioia and Pimentel (2000) and were carried out at the Geochronology Laboratory of the University of Brasilia. Whole-rock powders (~100 mg) were mixed with a ¹⁴⁹Sm-¹⁵⁰Nd spike solution and dissolved in Savillex^R capsules. Sr, Sm and Nd extraction from whole-rock samples followed conventional cation exchange techniques, with Teflon columns containing LN-Spec

Instrument	MC-ICP-MS Neptune
Scan Mode	Static
Scanned masses	238, 208, 207, 206, 204, 202
Laser	UP213 New Wave, 213 nm, Nd:YAG
Integration time	1 secs
Ablation time	40 secs
Spot size	30 µm
Laser settings	10 Hz, ~2J/cm ²
Auxiliary gas (He)	0.38-0.41 l/min
Sample gas (Ar)	0.903-0.995 l/min

Table 1: Operating data of LA-ICPMS

resin (HDEHP-diethylhexil phosphoric acid supported on PTFE powder). Sr, Sm and Nd samples were loaded onto Re evaporation filaments of double filament assemblies, and the isotopic measurements were carried out on a multicollector Finnigan MAT 262 mass spectrometer in static mode. Uncertainties of $^{87}\text{Sr}/^{86}\text{Sr}$ ratios were better than $\pm 0.01\%$ (2σ) and of Sm/Nd and $^{143}\text{Nd}/^{144}\text{Nd}$ ratios were better than $\pm 0.2\%$ (2σ) and $\pm 0.005\%$ (2σ), respectively, according to repeated analyses of international rock standards BHVO-1 and BCR-1. $^{143}\text{Nd}/^{144}\text{Nd}$ ratios were normalized to $^{146}\text{Nd}/^{144}\text{Nd}$ of 0.7219, and the decay constant used was $6.54 \times 10^{-12} \text{ y}^{-1}$. The T_{DM} values were calculated using DePaolo's model (1981).

5. Geochemistry

5.1. El Niño Muerto metadacites

Major and trace element analyses were carried out on three samples of the ENM dacites (results are in Table 2A, B). They have SiO_2 contents between 65.0 and 67.3 wt.% and

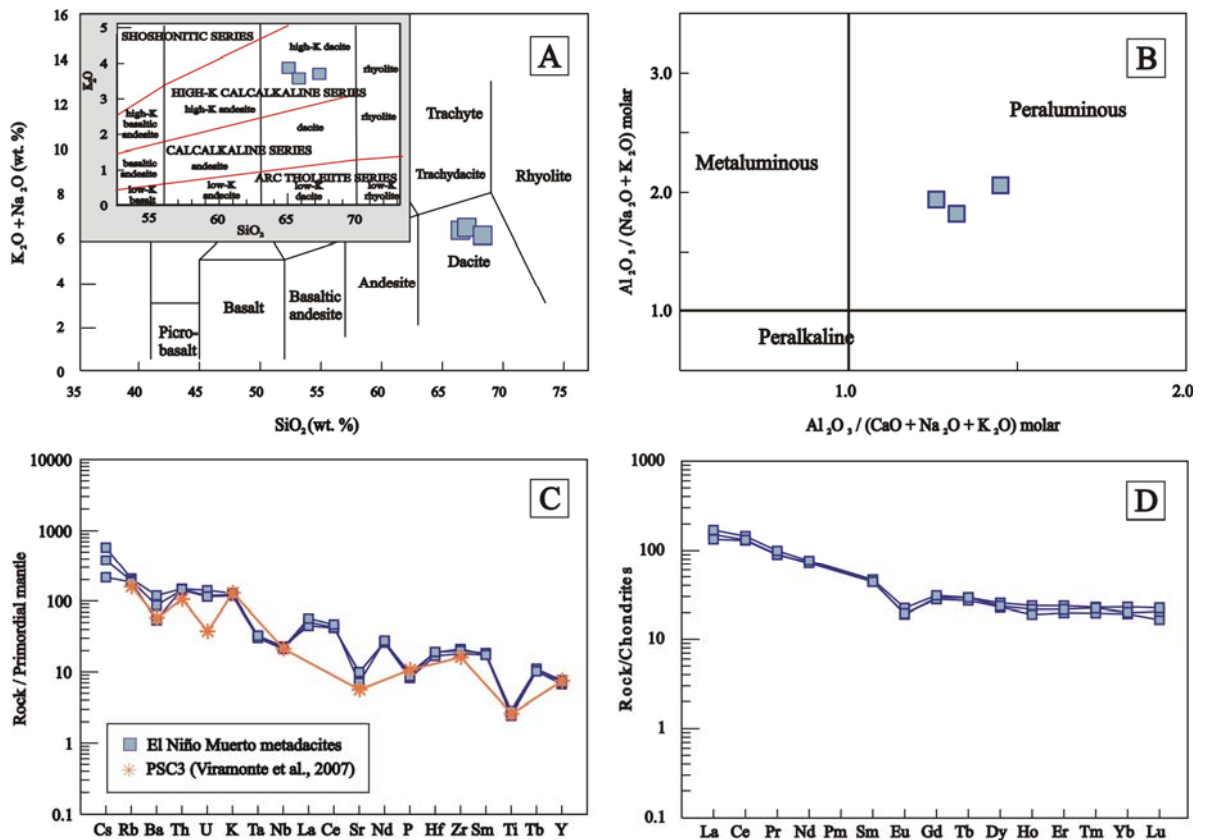


Figure 7: A) Chemical compositions of the El Niño Muerto subvolcanic rocks in the Total alkalis versus Silica diagram (Le Maitre et al., 1989) and in the Peccerilo and Taylor (1976) diagram for subalkaline rocks. B) The $\text{Al}_2\text{O}_3/(\text{CaO} + \text{Na}_2\text{O} + \text{K}_2\text{O})$ vs. $\text{Al}_2\text{O}_3/(\text{Na}_2\text{O} + \text{K}_2\text{O})$ diagram (Shand, 1929) shows the peraluminous nature of the original magmas. C) Spidergrams of Primordial Mantle-normalized trace elements distributions (Wood et al., 1979) for the El Niño Muerto dacites and sample PCS3 of “Faja Eruptiva de la Puna Oriental” (Viramonte et al., 2007). D) Chondrite-normalized diagram for representative samples of the El Niño Muerto metadacites. The normalizing factors are from Sun and McDonough (1989).

Al₂O₃ contents from 15.09 to 16.4 wt. %. In the Total Alkali Silica (TAS) diagram (Fig. 7 A) these metavolcanic rocks fall in the dacite field, and in the Peccerilo and Taylor (1979) diagram for subalkalic rocks, they are classified as high-K dacites (Fig. 7A). They show peraluminous composition (Fig. 7 B) with normative corundum greater than 5 % (Table 2).

The concentrations in compatible elements is low (~ 27 ppm of Cr, ~ 30 ppm of V and ~ 22 ppm of Ni) and these rocks show relative enrichment in Ba (up to 897 ppm of Ba), Sr (161-226 ppm), Rb (up to 177 ppm) and Nb (~ 27 ppm). In the normalized diagram of Wood et al. (1979) (Fig. 7 C) the subvolcanic rocks show enrichment in LILE elements between 100 and 800 times greater than the primordial mantle and depletion in the HFSE (~ 10 times) with negative anomalies in Ba, Ta, Nb, Sr, P and Ti elements. The negative anomalies in Nb and Ta, are characteristic of continental crust (subduction-signature) and depletion in P, Ti and Eu, indicate that apatite, ilmenite, and plagioclase crystallized in early stages of the magma evolution. Negative anomalies in Ba and Sr are consistent with plagioclase removal. Metacites of the El Niño Muerto Hill and a metadacite from the Faja Eruptiva de la Puna Oriental in the Puna (Viramonte et al., 2007) have very similar trace element distribution patterns (Fig. 7 C).

ΣREE (Fig. 7 D) varies from 186 to 206 ppm, with a distinctive enrichment in the LREE (La/Sm_N = 3.2-3.6). The (La/Yb)_n ratio between 5.9 to 8.6 indicates that they are moderately fractionated, with (La/Sm)_n values greater than (Gd/Yb)_n (3.3 and 1.4 respectively). The negative Eu anomaly, from -0.53 to -0.59, indicates plagioclase fractionation.

5.2. Rio Blanco basalts

In all bulk chemical analyses of the Rio Blanco rocks (results are in Table 2A, B) the loss on ignition values are high, varying from 2.5 to 13.8 %, indicating significant alteration.

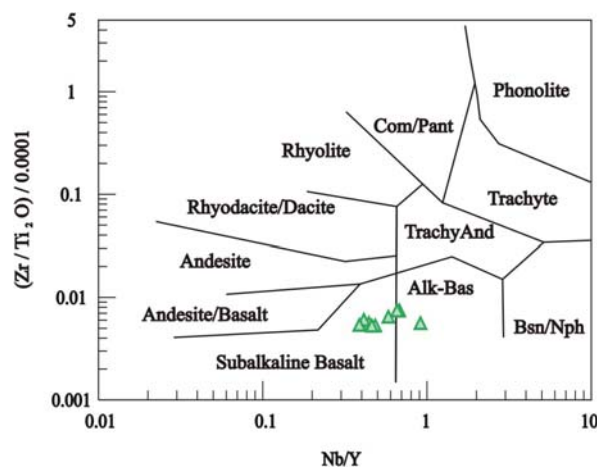


Figure 8: Rock classification diagram using Zr, Ti, Nb and Y immobile elements (Winchester and Floyd, 1977) for the Rio Blanco volcanic rocks.

Table 2A: Representative whole rock analyses of the El Niño Muerto and Río Blanco volcanic rocks.

SAMPLE	<i>El Niño Muerto</i>			<i>Río Blanco</i>								
	BNM 11	BNM 14	BNM 16	BRB 22	BRB 25	BRB 30	BRB 37	BRB 43	BRB 44	BRB 64	BRB 65	BRB 66
<i>Major oxides (wt%)</i>												
<i>SiO₂</i>	65.07	65.81	67.31	41.37	46.94	36.56	38.60	45.78	36.47	45.63	44.75	42.00
<i>TiO₂</i>	0.73	0.62	0.67	3.37	4.20	2.81	2.01	3.98	1.88	2.53	1.98	1.68
<i>Al₂O₃</i>	16.44	15.23	15.09	15.85	17.39	22.53	14.97	16.41	14.55	16.43	19.79	18.01
<i>Fe₂O_{3(t)}</i>	4.97	5.65	4.98	17.49	16.69	14.03	11.62	17.24	9.33	11.08	8.97	8.56
<i>MnO</i>	0.08	0.07	0.08	0.25	0.10	0.06	0.32	0.17	0.17	0.16	0.12	0.11
<i>MgO</i>	2.38	2.58	1.89	3.59	2.35	11.27	4.48	3.15	6.41	7.46	4.52	4.22
<i>CaO</i>	1.83	1.75	2.33	4.75	1.68	1.06	11.69	2.78	12.64	7.37	5.55	8.72
<i>Na₂O</i>	2.29	2.73	2.28	5.06	5.98	0.34	4.96	3.60	3.99	2.23	2.07	5.13
<i>K₂O</i>	3.88	3.58	3.71	1.35	1.24	3.33	0.15	2.23	0.48	0.04	3.31	1.25
<i>P₂O₅</i>	0.20	0.17	0.18	0.44	0.85	0.34	0.21	0.38	0.25	0.33	0.29	0.24
<i>L.O.I</i>	2.00	1.70	1.40	6.40	2.50	7.60	10.90	4.20	13.80	6.70	8.60	9.90
<i>Total</i>	99.87	99.89	99.92	99.92	99.92	99.93	99.91	99.92	99.97	99.96	99.95	99.82
<i>Q</i>	32.10	30.59	34.91	-	-	-	-	-	-	-	-	-
<i>C</i>	8.69	7.02	6.89	-	-	-	-	-	-	-	-	-
<i>Or</i>	23.52	21.64	22.34	-	-	-	-	-	-	-	-	-
<i>Ab</i>	19.88	23.63	19.66	-	-	-	-	-	-	-	-	-
<i>An</i>	0.00	0.00	1.56	-	-	-	-	-	-	-	-	-
<i>En</i>	6.08	6.57	4.80	-	-	-	-	-	-	-	-	-
<i>Fs</i>	4.65	5.59	4.74	-	-	-	-	-	-	-	-	-
<i>Hy</i>	10.73	12.16	9.53	-	-	-	-	-	-	-	-	-
<i>Mt</i>	1.57	1.78	1.56	-	-	-	-	-	-	-	-	-
<i>Il</i>	1.42	1.20	1.30	-	-	-	-	-	-	-	-	-
<i>Ap</i>	0.00	0.00	0.43	-	-	-	-	-	-	-	-	-
<i>Cc</i>	3.35	3.20	3.24	-	-	-	-	-	-	-	-	-
<i>Total</i>	101.27	101.23	101.43	-	-	-	-	-	-	-	-	-

Therefore, major elements oxides and large ion lithophile elements cannot be used to chemically classify or discriminate the tectonic setting of these rocks, immobile elements and REE have been used. Using the diagram of Winchester and Floyd, (1977) the Rio Blanco volcanic rocks fall between subalkaline basalts and alkaline basalts fields (Fig. 8). They show low Cr (mostly between 10 and 147 ppm) and Ni (32-61 ppm) contents and display variable range in incompatible elements (13-364 ppm Ba; 139-445 ppm Sr; 2.8-112 ppm Rb; 13.5-22.5 ppm Nb). The Rio Blanco basalts display Zr/Nb ratios from 6.8 to 9.5 which are characteristic of E-MORB's.

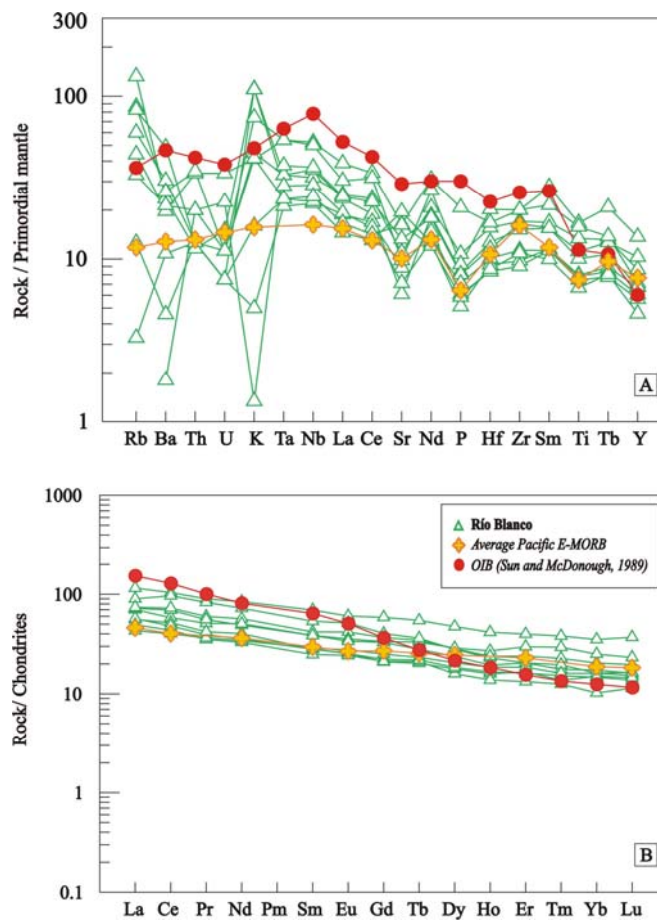


Figure 9: A) Spidergrams of Primordial Mantle-normalized trace elements distributions (Wood et al., 1979) for the Rio Blanco rocks. B) Chondrite-normalized REE abundance diagram for representative samples of the Rio Blanco basalts. The normalizing factors are from Sun and McDonough (1989). E-MORB (average Pacific E-MORB from GEOROC) and OIB (Sun and McDonough 1989) are plotted for comparison.

In the modified Wood et al., (1979) diagram (Fig. 8 A), normalized to the Primordial Mantle, the Rio Blanco (RB) samples show variable enrichment in large-ion lithophile elements (LILE) from 10 to approximately 100 times higher than the Primordial Mantle. It is noteworthy that the Rio Blanco samples do not show Ta and Nb anomalies. Strongly negative or positive anomalies in Th, K, Sr and Ti are observed. Trace elements abundances form a

moderate curvilinear pattern with large ion lithophile element (LILE) enrichment plotting between the OIB and E-MORB typical patterns.

The REE concentrations of the Rio Blanco rocks were normalized to the average C1 chondrite (Sun & McDonough 1989). The Rio Blanco samples exhibit Σ REE values between 76 and 129 ppm. The $(La/Yb)_n$ ratios are between 2.6 and 4.5 indicating slight LREE enrichment (Fig. 8 B). The Eu anomalies (Eu/Eu^*) vary between -0.9 and +0.9, suggesting that crystallization and fractionation of plagioclase was not very important. For comparison, the average Pacific E-MORB (from GEOROC.com) and the OIB of Sun and McDonough (1989) were also plotted in Fig. 9 A and B.

6. Geochronology

6.1. U-Pb data

Samples BNM 11 from the El Niño Muerto Hill and BRB 43 from the Rio Blanco Valley were dated by the U-Pb zircon method using Laser Ablation ICP-MS.

6.1.1. El Niño Muerto metadacites (sample BNM 11)

Two groups of zircons up to 200 μ m are recognized in this sample. One is characterized by needle-shape acicular pink zircon crystals. In agreement with the “Pupin Diagram” this type of zircon is classified as prismatic with greater development of the $\{100\}$ and $\{110\}$ faces (Pupin, 1980). They are very common in rapidly crystallized, porphyritic, subvolcanic intrusions and in high level granites and gabbros (Corfu et al., 2003). The second group exhibits more equant forms with short prismatic and multifaceted pink crystals. They are classified in the Pupin diagram as pyramidal crystals with relatively greater development of the $\{211\}$ and $\{101\}$ face (Pupin, 1980). These are more common in deep-seated, slowly cooled intrusions (Corfu et al., 2003).

Twenty nine zircon grains were analyzed (Table 3, Fig. 10 A) resulting in concordant analyses indicating the Concordia age of 495 ± 4 Ma (MSWD= 0.95) which is considered, in this study, as the best estimate for the crystallization age of the El Niño Muerto dacites. One spot analysis defined the age of ca. 950 Ma indicating inheritance from older crust. On the other hand, IDTIMS analyses of this sample indicated a poorly defined younger age of 462 ± 29 Ma (Table 4, Fig. 10 B) which is probably caused by a combination of inheritance older zircon crystals and recent lead loss. The IDTIMS data, however, reveal strong inheritance

Table 3: LA-ICPMS U-Pb isotopic data from the El Niño Muerto metadacites

Sample:	Shape*	Ratios						Ages						Disc (%) 206/238	rho
		²⁰⁷ Pb/ ²⁰⁶ Pb	2σ	²⁰⁷ Pb/ ²³⁵ U	2σ	²⁰⁶ Pb/ ²³⁸ U	2σ	²⁰⁷ Pb/ ²⁰⁶ Pb	± 2σ	²⁰⁷ Pb/ ²³⁵ U	± 2σ	²⁰⁶ Pb/ ²³⁸ U	± 2σ		
BNM 11															
Grain N°															
1	P	0.0571	0.000157	0.6255	0.017507	0.0800	0.002239	496.35	6	493.32	11	496.10	13	99.95	0.72
2	Pc	0.0581	0.000143	0.6830	0.025244	0.0860	0.003177	531.83	5	528.61	15	531.55	19	99.95	0.73
3	P	0.0576	0.000219	0.6603	0.039000	0.0832	0.004915	515.27	8	514.78	24	515.22	29	99.99	0.73
4	P	0.0569	0.000254	0.6171	0.022046	0.0786	0.002810	488.05	10	488.04	14	488.04	17	100.00	0.86
5	Pl	0.0574	0.000178	0.6454	0.039457	0.0820	0.005014	508.38	7	505.65	24	508.15	30	99.95	0.67
6	Pl	0.0574	0.000200	0.6400	0.028496	0.0817	0.003637	506.55	8	502.32	17	506.20	22	99.93	0.37
7	Pl	0.0576	0.000152	0.6473	0.033731	0.0828	0.004314	513.30	6	506.83	21	512.76	26	99.89	0.78
8	P	0.0574	0.000123	0.6408	0.016493	0.0817	0.002103	506.55	5	502.82	10	506.24	13	99.94	0.50
9	Pl	0.0578	0.000284	0.6711	0.020740	0.0841	0.002598	520.32	11	521.38	13	520.40	15	100.02	0.32
10	Pl	0.0575	0.000125	0.6477	0.017317	0.0825	0.002206	511.48	5	507.08	11	511.11	13	99.93	0.61
11	Pc	0.0577	0.000119	0.6614	0.022739	0.0839	0.002885	519.90	5	515.47	14	519.53	17	99.93	0.67
12	Pc	0.0577	0.000260	0.6609	0.047632	0.0840	0.006051	520.10	10	515.14	29	519.69	36	99.92	0.79
13	Pl	0.0572	0.000184	0.6314	0.062460	0.0804	0.007954	498.65	7	496.96	38	498.52	47	99.97	0.82
14	P	0.0573	0.000152	0.6371	0.028617	0.0811	0.003644	503.10	6	500.50	18	502.89	22	99.96	0.83
15	Pc	0.0569	0.000220	0.6168	0.030734	0.0784	0.003905	486.23	9	487.82	19	486.35	23	100.03	0.54
16	Pc	0.0571	0.000261	0.6211	0.018261	0.0798	0.002346	495.13	10	490.52	11	494.77	14	99.93	0.57
17	P	0.0566	0.000196	0.6011	0.028528	0.0766	0.003637	475.77	8	477.90	18	475.93	22	100.03	0.39
18	Pc	0.0568	0.000256	0.6098	0.029030	0.0778	0.003703	482.82	10	483.43	18	482.87	22	100.01	0.54
19	Pc	0.0571	0.000260	0.6172	0.027456	0.0798	0.003552	495.76	10	488.07	17	495.15	21	99.88	0.59
20	Pc	0.0570	0.000165	0.6181	0.023118	0.0790	0.002955	490.21	6	488.63	14	490.09	18	99.97	0.65
21	Pc	0.0566	0.000193	0.5998	0.046372	0.0766	0.005924	475.81	8	477.09	29	475.91	35	100.02	0.86
22	Pc	0.0566	0.000298	0.6028	0.026690	0.0766	0.003391	475.36	12	478.99	17	475.64	20	100.06	0.50
23	Pl	0.0566	0.000269	0.5963	0.015623	0.0768	0.002012	477.13	11	474.90	10	476.96	12	99.96	0.47
24	Pl	0.0571	0.000228	0.6193	0.017517	0.0795	0.002250	493.78	9	489.42	11	493.43	13	99.93	0.48
25	Pl	0.0572	0.000260	0.6281	0.058307	0.0802	0.007449	497.78	10	494.91	36	497.55	44	99.95	0.49
26	Pc	0.0568	0.000270	0.6222	0.027842	0.0781	0.003496	484.39	10	491.21	17	484.92	21	100.11	0.87
27	Pl	0.0571	0.000142	0.6295	0.023672	0.0802	0.003015	497.26	5	495.79	15	497.14	18	99.98	0.60
28	Pc	0.0572	0.000259	0.6269	0.027953	0.0804	0.003587	499.19	10	494.15	17	498.79	21	99.92	0.33
29	Pl	0.0572	0.000360	0.6277	0.041617	0.0807	0.005352	500.98	14	494.69	26	500.48	32	99.90	0.45

(*) P: prismatic crystals, Pc: short prismatic crystals and Pl: long prismatic crystals. Disc. (%) denotes percentage of discordance

Table 4: Summary of ID-TIMS U-Pb data for metadacite of the El Niño Muerto

Fraction	Ratios										Age				
	Weight (mg)	U ppm	Pb ppm	$^{206}\text{Pb}/^{204}\text{U}$	$^{207}\text{Pb}^*/^{235}\text{U}$	(pct) ^a	$^{206}\text{Pb}/^{238}\text{U}$	(pct) ^b	(rho)	$^{207}\text{Pb}^*/^{206}\text{Pb}$	(pct) ^c	$^{206}\text{Pb}^*/^{238}\text{U}$	$^{207}\text{Pb}^*/^{235}\text{U}$	$^{207}\text{Pb}^*/^{206}\text{Pb}$	(Ma)**
1	0.023	288.29	22.6410	1702.30	0.6201	0.70	0.0773	0.66	0.94	0.058178	0.235	479.97	489.88	536.47	5.1
2	0.025	246.38	21.7820	3839.91	0.7358	0.69	0.0881	0.68	0.99	0.0605692	0.102	544.36	559.98	623.97	2.2
3	0.03	327.96	25.2850	3248.57	0.5995	0.57	0.0756	0.54	0.96	0.0575178	0.165	469.8	476.95	511.44	3.6
4	0.03	314.94	25.5480	1658.77	0.6420	0.92	0.0800	0.87	0.95	0.0581947	0.276	496.19	503.55	537.1	6
5	0.024	322.95	31.8500	1523.70	0.8996	0.76	0.1005	0.72	0.95	0.0649094	0.245	617.46	651.53	771.4	5.2
6	0.063	509.92	41.0610	3382.16	0.6707	0.59	0.0824	0.57	0.98	0.0590506	0.121	510.28	521.13	568.96	2.6
7	0.059	215.66	22.6020	875.44	0.9417	1.22	0.1020	0.99	0.82	0.0669922	0.698	625.84	673.78	837.53	15
8	0.059	316.49	25.6900	1708.17	0.6687	0.63	0.0808	0.61	0.96	0.0600526	0.175	500.62	519.89	605.47	3.8

(*) Common Pb corrected using measured ^{204}Pb (pct)^a= Percentage of error of $^{207}\text{Pb}^*/\text{U}^{235}$; (pct)^b= Percentage of error of $^{206}\text{Pb}^*/^{238}\text{U}$; (pct)^c= Percentage of error of $^{207}\text{Pb}^*/^{206}\text{Pb}$. (rho)= correlation coefficient of error between $^{207}\text{Pb}^*/\text{U}^{235}$ and $^{206}\text{Pb}^*/^{238}\text{U}$. (**) Millions years error for individual fraction.

from Mesoproterozoic ca. 1.16 Ga (Grenville?) sialic basement, which is not obvious from the LA-ICPMS data.

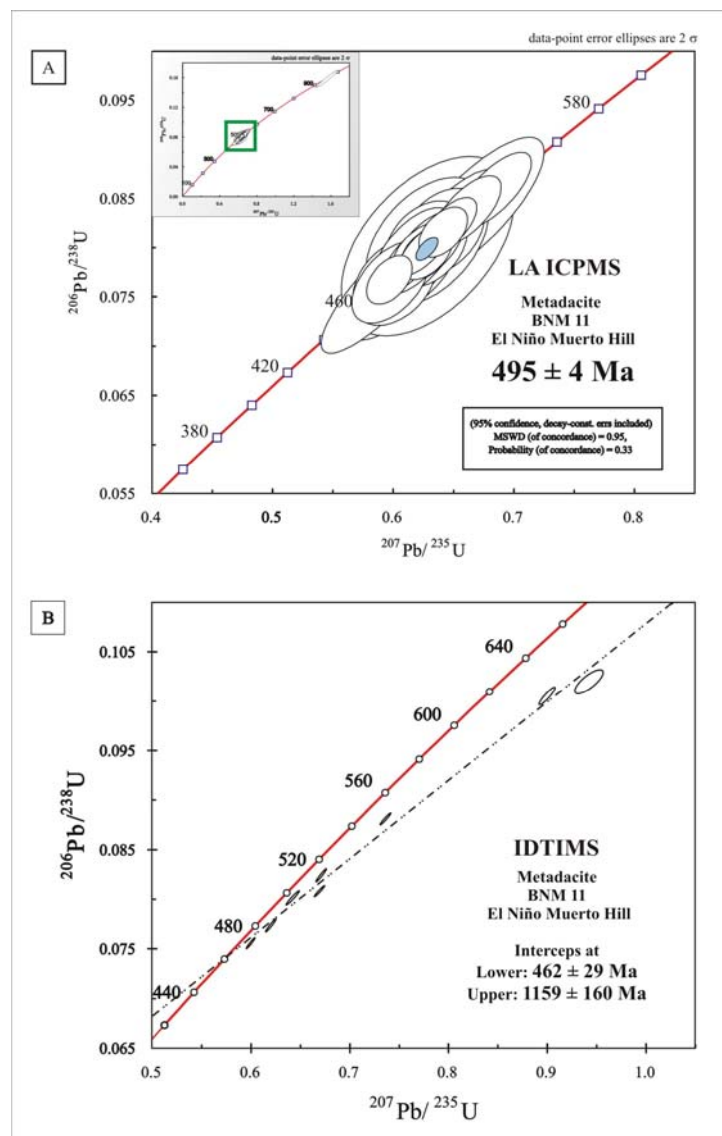


Figure 10 Zircon U-Pb data for the El Niño Muerto Hill. A) LA-ICPMS Concordia Age. B) ID-TIMS.

6.1.2. Río Blanco basalts (sample BRB 43)

Pink, transparent sub-rounded to prismatic zircon crystals up to 100 μm across are identified in this sample.

A total of seventy zircons were analyzed using Laser Ablation ICPMS (Table 5; Fig. 11) resulting in concordant analyses. On the basis of their distribution, two distinct groups of ages are recognized (Fig. 11): a younger age of 501 ± 9 Ma determined for four

Table 5: LA-ICPMS U-Pb isotopic data from the Río Blanco volcanic rocks

Sample:	Shape*	Ratios						Ages						Disc (%) 206/238	rho
		²⁰⁷ Pb/ ²⁰⁶ Pb	2σ	²⁰⁷ Pb/ ²³⁵ U	2σ ind	²⁰⁶ Pb/ ²³⁸ U	2σ	²⁰⁷ Pb/ ²⁰⁶ Pb	± 2σ	²⁰⁷ Pb/ ²³⁵ U	± 2σ	²⁰⁶ Pb/ ²³⁸ U	± 2σ		
BRB 43															
Grain N°															
1	<i>Psr</i>	0.0671	0.000763	1.3545	0.082017	0.1397	0.008458	839.59	23	869.51	35	842.84	48	100.39	0.79
2	<i>Psr</i>	0.0583	0.000244	0.7369	0.021919	0.0880	0.002617	542.15	9	560.62	13	543.71	15	100.29	0.02
3	<i>Psr</i>	0.0586	0.000246	0.7548	0.022448	0.0899	0.002675	553.77	9	570.98	13	555.22	16	100.26	0.02
4	<i>Psr</i>	0.0582	0.000571	0.6926	0.025591	0.0866	0.003200	535.52	22	534.34	15	535.42	19	99.98	0.66
5	<i>Psr</i>	0.0569	0.000407	0.6479	0.043151	0.0790	0.005259	488.43	16	507.19	26	489.93	31	100.31	0.35
6	<i>Psr</i>	0.0575	0.000344	0.6423	0.022619	0.0822	0.002895	509.90	13	503.76	14	509.40	17	99.90	0.64
7	<i>Psr</i>	0.0585	0.000477	0.7100	0.054497	0.0890	0.006832	550.10	18	544.75	32	549.66	40	99.92	0.89
8	<i>Psr</i>	0.0575	0.000341	0.6435	0.034272	0.0824	0.004386	510.69	13	504.46	21	510.19	26	99.90	0.84
9	<i>Psr</i>	0.0586	0.000452	0.7207	0.030073	0.0896	0.003740	553.55	17	551.07	18	553.34	22	99.96	0.82
10	<i>Psr</i>	0.0568	0.000376	0.6432	0.028215	0.0784	0.003437	484.73	14	504.28	17	486.29	21	100.32	0.57
11	<i>Psr</i>	0.0617	0.000095	0.9304	0.031254	0.1085	0.003646	663.86	3	667.83	16	664.22	21	100.05	0.92
12	<i>Psr</i>	0.0592	0.000169	0.7702	0.030992	0.0931	0.003748	573.58	6	579.86	18	574.12	22	100.09	0.60
13	<i>Psr</i>	0.0597	0.000810	0.7837	0.061369	0.0963	0.007537	592.88	30	587.60	34	592.42	44	99.92	0.35
14	<i>Psr</i>	0.0593	0.000108	0.7738	0.028624	0.0939	0.003474	578.28	4	581.95	16	578.59	20	100.05	0.64
15	<i>Psr</i>	0.0598	0.000284	0.7957	0.033493	0.0970	0.004081	596.75	10	594.42	19	596.55	24	99.97	0.29
16	<i>Psr</i>	0.0591	0.000170	0.7553	0.031977	0.0927	0.003923	571.33	6	571.29	18	571.33	23	100.00	0.61
17	<i>Psr</i>	0.0614	0.000114	0.8879	0.024587	0.1066	0.002953	653.93	4	645.23	13	653.14	17	99.88	0.90
18	<i>Psr</i>	0.0589	0.000750	0.7775	0.217442	0.0917	0.025653	564.05	27	584.08	117	565.76	150	100.30	0.76
19	<i>Psr</i>	0.0565	0.001260	7.1717	0.307598	0.1309	0.005614	472.37	34	2133.00	38	792.93	32	167.86	0.95
20	<i>Psr</i>	0.0602	0.000195	0.8133	0.063000	0.0994	0.007699	611.50	7	604.33	35	610.87	45	99.90	0.92
21	<i>Psr</i>	0.0614	0.000947	1.2505	0.040131	0.1093	0.003507	651.60	30	823.61	18	668.61	20	102.61	0.60
22	<i>Psr</i>	0.0597	0.000114	0.7899	0.022773	0.0961	0.002769	591.34	4	591.13	13	591.32	16	100.00	0.74
23	<i>Psr</i>	0.0612	0.000148	0.8763	0.034411	0.1052	0.004131	645.44	5	638.97	18	644.85	24	99.91	0.36
24	<i>Psr</i>	0.0776	0.000230	2.4873	0.101772	0.1963	0.008032	1136.30	6	1268.35	29	1155.43	43	101.68	0.84
25	<i>Psr</i>	0.0595	0.000218	0.7937	0.024875	0.0951	0.002981	585.04	8	593.27	14	585.75	18	100.12	0.57
26	<i>Psr</i>	0.0598	0.000219	0.8067	0.025281	0.0972	0.003045	597.54	8	600.58	14	597.80	18	100.04	0.57
27	<i>Psr</i>	0.0599	0.000396	0.8004	0.058406	0.0976	0.007121	600.56	14	597.06	32	600.25	42	99.95	0.37

(*) *Psr*: prismatic subrounded crystals

concordant spot analysis and an older age of 581 ± 8 Ma for nine concordant zircon analyses. The age values ranging from ca. 580 to ca. 1200 Ma observed for the Rio Blanco basalts represent inherited zircons and agree with part of the age spectrum for the detrital zircons of the Puncoviscana Formation (Adams et al., 2006), suggesting contamination of the original magmas with these sediments. Adams et al. (2006) recognized, in Puncoviscana rocks, detrital zircons ages ranging from 580 to 1100 Ma, with two significant age groups at ca. 860-1100 and 580-760 Ma. Therefore zircon grains with ages between ca. 580 and 1200 Ma in the Rio Blanco volcanics may be interpreted as inherited zircons from the Punillica phase (Adams et al., 2006) of the Pampean Ranges. The younger age of **501 ± 9 Ma** is interpreted in the present study, therefore, as the crystallization of the original basalts.

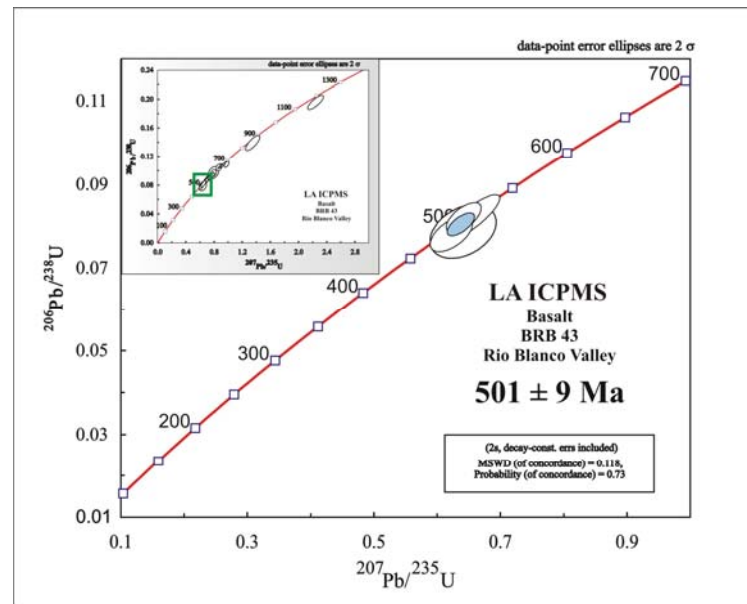


Figure 11: LA ICPMS Zircon U-Pb data for the Rio Blanco basalt

The presence of peperite deposits in the RB Valley indicates that the deposition of the turbidite sequence was contemporaneous with the extrusion of the RB Basalts, at approximately 500 Ma. Therefore, the sedimentary succession investigated here is

younger than the Puncoviscana Formation which has a minimum age of 530 Ma, being intruded by the Tastil granite at ca. 535 Ma (Bachmann et al., 1987). This is in agreement with the 527 Ma U-Pb age on detrital zircon found by Lork et al. (1990) in a similar turbidite sequence 30 km to the SW of the study area, indicating that this whole sedimentary succession can not be correlated with the Puncoviscana Formation.

6.2. Nd and Sr Isotopic Data

Fifteen representative samples of the El Niño Muerto Hill and the Rio Blanco igneous rocks were selected for whole-rock Nd isotope analyses. Sm and Nd elemental abundances were measured by isotope dilution following the method described by Gioia and Pimentel (2000). Only five whole-rock samples were selected for Sr isotope analyses. Rb and Sr concentrations were derived from ICPMS analyses. The isotopic data are listed in Table 6.

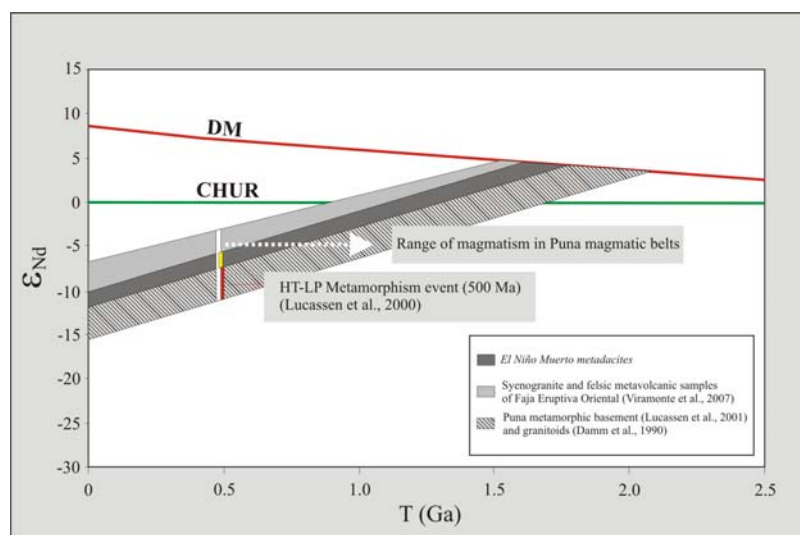


Figure 12: ϵ_{Nd} versus age diagram for the El Niño Muerto metadacites. The Nd isotopic compositions of these rocks are similar to the felsic metavolcanic samples of Faja Eruptiva Oriental and the metamorphic and granitoids rocks of the basement of Puna.

Nd and Sr isotopic ratios for the El Niño Muerto Hill were re-calculated to **495** Ma. The analyzed samples have fairly homogeneous high initial $^{87}\text{Sr}/^{86}\text{Sr}$ ratios ranging between 0.71107 and 0.71180. All samples have negative $\epsilon_{Nd}(t)$ values between -9.7 (BNM 16) and -5.9 (BNM 10) and their T_{DM} model ages range between 1.62 and 1.79 Ga (Table 6). Nd and Sr isotopic data for the Rio Blanco basalts were corrected to **501** Ma. The initial $^{87}\text{Sr}/^{86}\text{Sr}$ isotopic ratios of three Rio Blanco samples vary between 0.70713 and 0.70942 (Table 5). All samples have positive initial $\epsilon_{Nd}(t)$ values, between +2.5 (BRB 30) and +4.5 (BLC 95) with only one sample (BRB 25) showing a slightly negative $\epsilon_{Nd}(t)$ value of -0.9. The T_{DM} model ages for the Rio Blanco basalts vary from 0.84 to 1.12 Ga.

Table 6: Sr-Nd isotopic data for volcanic rocks from the El Niño Muerto metadacites and Río Blanco basalts.

Locality	Sample	Rb(***) (ppm)	Sr (***) (ppm)	$^{87}\text{Rb}/^{86}\text{Sr}$	$(^{87}\text{Sr}/^{86}\text{Sr})_{\text{meas}}$	$(^{87}\text{Sr}/^{86}\text{Sr})_{\text{t}}$	Sm (ppm)	Nd (ppm)	$^{147}\text{Sm}/^{144}\text{Nd}$	$(^{143}\text{Nd}/^{144}\text{Nd})_{\text{meas}}$	$(^{143}\text{Nd}/^{144}\text{Nd})_{\text{t}}$	$\epsilon_{\text{Nd}}(t)$	T_{DM}
El Niño Muerto (*)	BMN 10						7.79	37.17	0.1267	0.512108 (± 10)	0.511697	-5.92	1.63
	BMN 11	177.50	226.80	2.2690	0.72776 (± 02)	0.71179	6.17	28.15	0.1325	0.512086 (± 13)	0.511656	-6.71	1.79
	BMN 12						6.189	30.10	0.1243	0.512091 (± 06)	0.511688	-6.10	1.62
	BMN 14						6.53	32.10	0.1230	0.512068 (± 11)	0.511669	-6.46	1.63
	BMN 15						6.452	30.71	0.1270	0.512107 (± 15)	0.511695	-5.96	1.64
	BNM 16	158.40	161.70	2.8407	0.73111 (± 05)	0.71107	9.29	33.17	0.1693	0.512055 (± 14)	0.511506	-9.65	-
Río Blanco (**)	BRB 22						6.75	26.29	0.1552	0.512719 (± 03)	0.512210	4.24	0.85
	BRB25	37.50	373.80	0.2903	0.71054 (± 02)	0.70847	13.15	39.20	0.2028	0.512612 (± 09)	0.511946	-0.90	-
	BRB 30						6.13	24.38	0.1521	0.512618 (± 15)	0.512119	2.47	1.05
	BRB 37	2.80	445.40	0.0182	0.70726 (± 02)	0.70714	8.26	30.44	0.1641	0.512696 (± 10)	0.512158	3.23	1.06
	BRB 43	73.40	223.60	0.9505	0.71621 (± 03)	0.70942	8.362	34.31	0.1473	0.512677 (± 15)	0.512193	3.93	0.84
	BRB 44						4.06	15.20	0.1615	0.512715 (± 12)	0.512185	3.76	0.95
	BRB 64						6.88	25.92	0.1606	0.512649 (± 18)	0.512122	2.54	1.12
	BRB 65						4.11	16.79	0.1480	0.512639 (± 14)	0.512153	3.14	0.94
	BRB 66						4.25	17.28	0.1485	0.512665 (± 13)	0.512177	3.60	0.88

Obs:

Errors on measured isotope ratios are given in parenthesis as uncertainty in last digits (2s)

$\epsilon_{\text{Nd}}(T)$ calculated relative to CHUR with present-day value of $^{143}\text{Nd}/^{144}\text{Nd} = 0.512638$. T_{DM} values were calculated using DePaolo model (1981).

(*) $^{87}\text{Sr}/^{86}\text{Sr}$ and $^{143}\text{Nd}/^{144}\text{Nd}$ and $\epsilon_{\text{Nd}(t)}$ calculate with an age of 495 ± 4 Ma

(**) $^{87}\text{Sr}/^{86}\text{Sr}$ and $^{143}\text{Nd}/^{144}\text{Nd}$ and $\epsilon_{\text{Nd}(t)}$ calculate with an age of 501.0 Ma

(***) Elemental abundances of Rb and Sr are taken from ICP-MS analyses

In the ϵ_{Nd} vs. time diagram (Fig. 12) the El Niño Muerto Hill dacites are compared with data for the Central Andean metamorphic-magmatic basement (Lucassen et al., 2000, 2001; Viramonte et al., 2007). The basement has a homogeneous distribution of Sm-Nd model ages (T_{DM}) between ca. 2.0 and 1.6 Ga (Lucassen et al., 2001). The El Niño Muerto dacites have very similar Nd isotopic characteristics compared to the metamorphic basement rocks of the Puna Plateau. They are also very similar isotopically and geochronologically to metadacites, metarhyolites and granitoids of the Faja Eruptiva de la Puna Oriental, either in the Salar Centenario area (485 ± 5 Ma; the Oire Igneous-Metamorphic Complex, Viramonte et al. 2007) or in the Calalaste Range (495 ± 3 Ma; Pinheiro et al., 2008), suggesting that the original magmas formed during a Late Cambrian-Early Ordovician magmatic event which reworked Mesoproterozoic or Paleoproterozoic continental crust.

7. Discussion and conclusions

The new petrographic, geochemical and isotopic data of the El Niño Muerto dacites (ENM) and the Rio Blanco basalts (RB) interlayered into the Lower paleozoic sediments of the Andean basement in NW Argentina are discussed below in order to constrain the magmatic evolution and tectonic setting of the study area.

7.1 El Niño Muerto metadacites

The new data for the ENM rocks indicate that they are high K calc-alkaline dacites. The calcalkaline, subduction-related, signature is also evident by the strong fractionation between LILE and HFSE and by the Nb-Ta negative anomaly.

The peraluminous character of the ENM rock suggests that they may have been generated by partial melting of sedimentary rocks or that they derived from a more primitive source but acquired the peraluminous nature by intense assimilation of high Al_2O_3 crustal material. Their high initial $^{87}\text{Sr}/^{86}\text{Sr}$ isotopic ratios and negative ϵ_{Nd} (t) values (-5.2 to -9) confirm the involvement of older crustal material during the genesis of the original magmas. T_{DM} model ages between 1.63 to 1.79 Ga of ENM rocks and the zircon inheritance pattern suggest participation of Mesoproterozoic continental crust which is in agreement with other basement data for the Central Andes (Lucassen et al., 2001, Viramonte et al., 2007).

7.2. Rio Blanco basalts

The RB submarine volcanic rocks have been affected by intense hydrothermal alteration that **modifies** the distribution of LIL elements and, for that reason, the application of major element oxides and large ion lithophile elements is not useful for identifying their paleotectonic setting. Therefore, the classification based on immobile elements (Winchester and Floyd 1977) is needed, and it shows that the RB rocks are transitional between subalkaline and alkali basalts (Fig. 8).

Concentrations of Th, Hf and Ta of the RB basalts are similar to either E-MORB or Within Plate Basalts (Fig. 13 A; Wood, 1980). Similarly, Zr, Ti and Y concentrations of most of the RB rocks are similar to those of within plate basalts (Fig. 13 B, Pearce and Cann, 1973). In the Zr-Nb-Y diagram (Meschede, 1986) (Figure 13 C) rock samples of the RB basalts seem to be transitional between AII (within plate alkali basalts) and B (plume-MORB). The RB basalts have low Zr/Nb ratios (~ 7), which is characteristic of rocks with E-MORB affinity or of oceanic island tholeiites (OIT) (Wood et al., 1979). They show moderate enrichment in the LREE and the lack of Eu anomaly suggests that contamination with older continental crust was not important, which is also supported by the generally positive values of $\epsilon_{\text{Nd}}(t)$, although inherited zircon grains have been recognized. Negative Nb anomalies are not observed either. The slightly high Sr isotopic initial ratio, the enrichment in LILE elements and some of the Nd isotopic characteristics may reflect original features of the mantle source, combined with some degree of crustal assimilation.

The high $^{87}\text{Sr}/^{86}\text{Sr}$ isotopic ratio (0.70714-0.70942) and the positive $\epsilon_{\text{Nd}}(t)$ values (from +2.47 to +4.24) suggest that these rocks are similar to OIB or E-MORB basalts, although the high $^{87}\text{Sr}/^{86}\text{Sr}$ isotopic ratio observed in many E-MORB or OIB are attributed to seawater-alteration effects.

7.3. Age and tectonic evolution

The 495 ± 4 Ma zircon age of the El Niño Muerto dacites places them into the initial stages of the Famatinian Cycle (Rapela et al., 1998). The ENM dacites, on the basis of the new field, geochemical, isotopic and geochronological data, are part of the Faja Eruptiva Oriental (Fig. 1) which is characterized by the presence of volcano-sedimentary sequences composed by metadacites and metarhyolites of similar age (ca. 485 to 475 Ma) and geochemical nature (e.g. Viramonte et al., 2007).

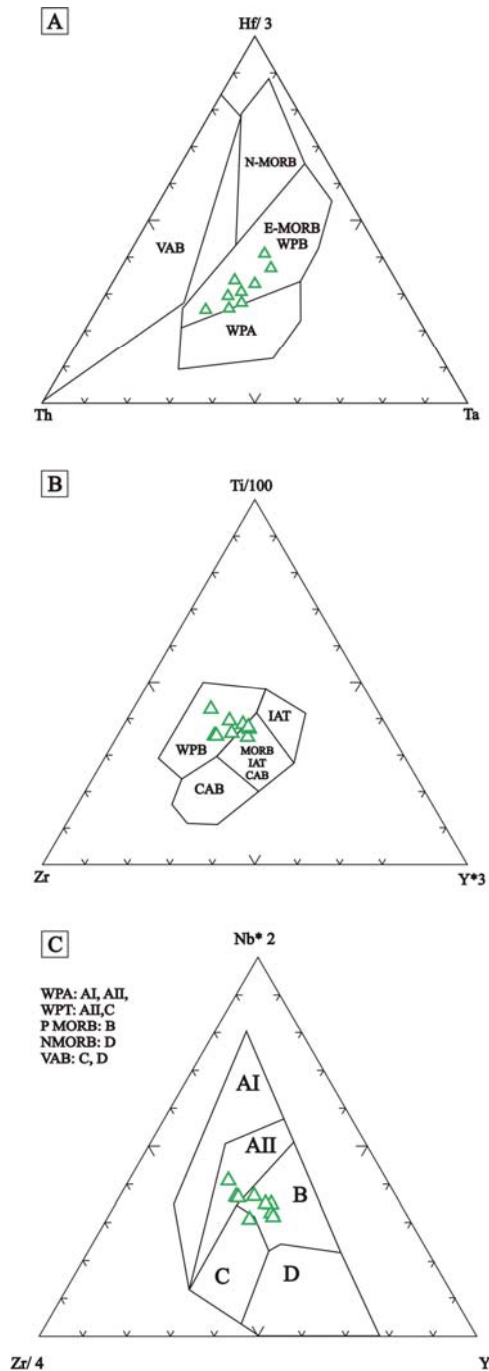


Figure 13: Tectonic discriminant diagrams for the Rio Blanco basalts. A) Triangular discriminate Th-Hf/3-Ta diagram after Wood 1980. Abbreviations: N-MORB= normal mid-ocean ridge basalt, E-MORB and WPB= enriched mid-ocean ridge basalt and within plate basalt respectively, WPA= within plate alkali, VAB= volcanic arc basalt. B) Triangular discriminant diagram of Pearce and Cann (1973). Abbreviations: IAT= island-arc tholeiite, MORB= mid-ocean ridge basalt, CAB=calc-alkali basalt, WPB= within plate basalt. C) Triangular discriminant diagram of Meschede, 1986.

The 501 ± 9 Ma U-Pb age of the RB basalts indicate that they are roughly synchronous with ENM arc dacites. Based on the field relationships as well as on the geochemical and isotopic E-MORB signature, the RB basalts may be interpreted as a remnant of a back-arc basin related with the Ordovician arc in NW Argentina. This model is in

agreement with that proposed by Miller and Sollner (2005) for the evolution of the Famatinian magmatic arc during the Ordovician.

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CAPÍTULO IV

4.1. Conclusões

Os novos dados de campo, geoquímicos e isotópicos obtidos para as rochas vulcânicas do embasamento do Noroeste Argentino permitiram as seguintes conclusões.

Setor Norte: Vale do Río Grande

1) Com base na composição dos minerais e na geoquímica de rocha total, os diques alojados na Formação Puncoviscana no Vale do Río Grande, são re-classificados como *ocellar analcima monchiquites*. Estes são uma variedade de lamprófiro alcalino sem feldspato ou plagioclásio, caracterizados pela presença de diopsídio, biotita/flogopita, estas últimas ricas em Ti, pargasita titanífera e olivina forsterítica como fenocristais em uma matriz de composição semelhante, porém sem olivina. A nomenclatura dessas rochas foi baseada na presença de analcima e carbonato nas abundantes estruturas globulares presentes.

2) A idade K-Ar de 163 ± 8 Ma obtida em biotitas e os diagramas discriminantes de ambientes que indicam ambiente de intraplaca para essas rochas permitiram associar os lamprófiros do Vale de Río Grande com a etapa de pré-rift do Rift de Salta, desenvolvido durante o Cretáceo no noroeste argentino.

3) As altas razões iniciais de $^{87}\text{Sr}/^{86}\text{Sr}$ e os valores positivos de $\epsilon_{\text{Nd}(t)}$, em conjunto com as características observadas durante o estudo petroquímico sugerem que o manto do qual estas rochas derivam é um manto metassomatizado. Altos conteúdos em SrO no carbonato presente nas estruturas globulares, na matriz e em alguns xenólitos indicam que ele é de origem magmática. Isto pode ser compatível com a participação de fluidos carbonatíticos durante metassomatismo.

4) As T_{DM} destas rochas, entre 0.64 e 0.25 Ga, indicam, bem como os dados da literatura, que o evento de metassomatismo possivelmente ocorreu durante a Orogenia Famatiniana (460 Ma).

5) Pode-se concluir que as rochas deste setor não estão associadas aos orógenos paleozóicos da Plataforma Sulamericana.

1) As rochas félicas da Serra de El Niño Muerto são classificadas como dacitos calcialcalinos de alto-K e são relacionadas a ambientes de subducção. Seu caráter peraluminoso indica que elas possivelmente derivam da fusão parcial de rochas sedimentares ou que rochas sedimentares foram assimiladas pelos magmas mais primitivos. O alto fracionamento de LILE e HFSE, as anomalias negativas em Nb-Ta, as altas relações iniciais de $^{87}\text{Sr}/^{86}\text{Sr}$ e os valores negativos de $\epsilon_{\text{Nd}(t)}$ sugerem que os dacitos de El Niño Muerto foram alojados em um ambiente de arco continental. Em sua gênese houve participação de crosta continental possivelmente mesoproterozóica, como é indicam as idades modelo em torno 1.70 Ga.

2) Com base em elementos imóveis, as rochas de Río Blanco possuem características intermediárias entre basaltos subalcalinos e basaltos alcalinos. Os diagramas discriminantes indicam que elas foram alojadas em ambiente de intraplaca, ou seja, em ambiente de E-MORB. Razões iniciais de $^{87}\text{Sr}/^{86}\text{Sr}$ intermediárias e $\epsilon_{\text{Nd}(t)}$ positivos indicam que a fonte dos magmas era mantélica e levemente enriquecida.

3) As análises U-Pb dos dacitos de El Niño Muerto fornecem idade de **495 ± 4 Ma**. Assim, eles estão relacionados aos estágios iniciais do magmatismo do Ciclo Famatiniano, desenvolvido durante o Ordoviciano no centro-noroeste da Argentina. Pela localização, estes pertencem ao arco mais interno, conhecido na literatura como Faixa Eruptiva da Puna Oriental.

4) As análises U-Pb dos basaltos de Río Blanco fornecem idade de **501 ± 8 Ma**, o que indicaria que estes são contemporâneos aos dacitos de El Niño Muerto. As características de campo, geoquímicas e isotópicas mostram afinidades E-MORB e permitem interpretar os basaltos de Río Blanco como possíveis remanescentes de um back-arc desenvolvido a leste do arco Famatiniano, no noroeste de Argentina.

Os resultados obtidos com base em dados geológicos do noroeste da Argentina permitem acrescentar e esclarecer os processos ocorridos durante o Paleozóico Inferior. Por outro lado levantam novos problemas geológicos e novos pontos de vista sobre a evolução da margem oeste de Gondwana requerendo, portanto, estudos mais detalhados neste setor dos Andes Centrais.

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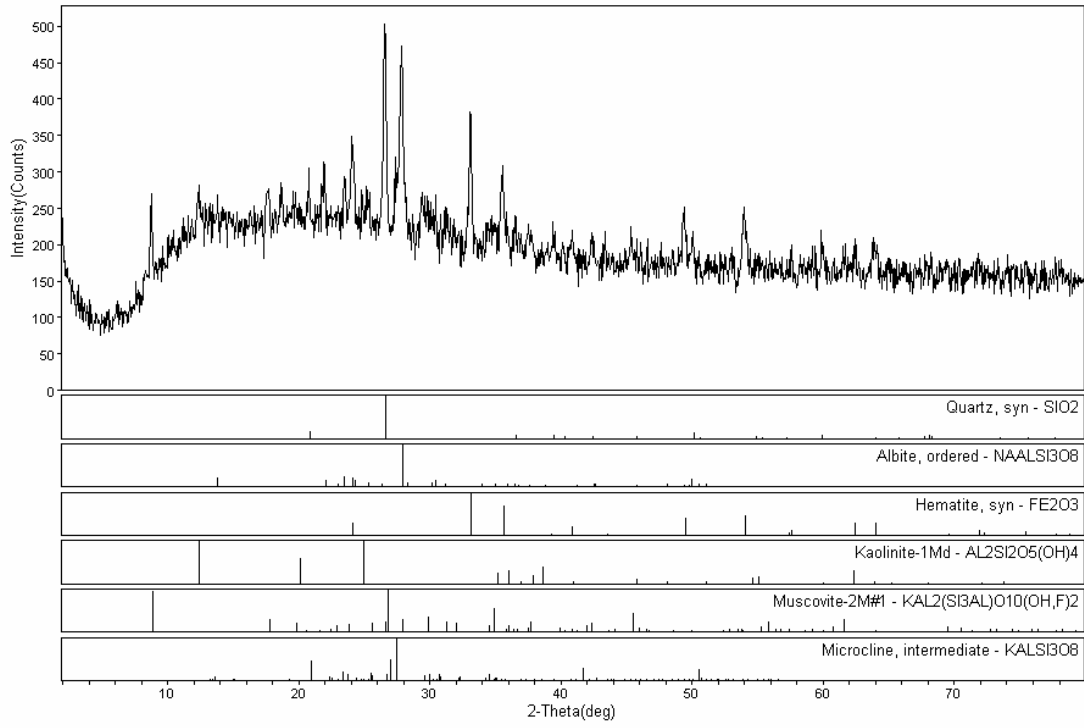
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ANEXO I:
DIFRATOMETRIA DE RAIOS-X

Amostra BRB 43



**ANEXO II:
ANALISES DE MICROSONDA**



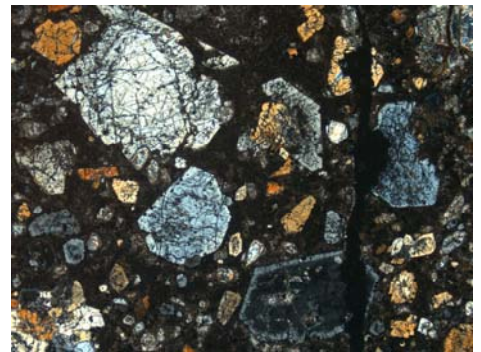
OLIVINA

ANEXO: Análises de microsonda de cristais de olivina da amostra BYA 124a, lamprofiros de Ríó Grande

<i>AMOSTRA BYA 124A</i>					
	Phenocristais			Xenolitos	
	1	2	3	4	5
<i>Amostra</i>	BYA124A.10.OLV.214	BYA124A.7.OLV.266	BYA124A.5.OLV.280	BYA124A.4.OLVXEN.236	BYA124A.4.OLVXEN.238
<i>SiO₂</i>	40.37	38.06	39.49	39.98	39.60
<i>TiO₂</i>	0.00	0.03	0.02	0.04	0.03
<i>Al₂O₃</i>	0.03	0.02	0.01	0.04	0.07
<i>FeO_(t)</i>	17.40	21.96	17.04	17.80	14.28
<i>MnO</i>	0.13	0.31	0.22	0.20	0.19
<i>MgO</i>	42.51	39.05	42.86	42.45	44.50
<i>CaO</i>	0.19	0.13	0.31	0.19	0.21
<i>Cr₂O₃</i>	0.03	0.01	0.01	0.03	0.03
<i>NiO</i>	0.12	0.18	0.29	0.21	0.30
<i>K₂O</i>	0.01	0.00	0.01	0.01	0.00
<i>Total</i>	100.77	99.75	100.25	100.95	99.20
<i>FO</i>	81.42	75.89	81.92	80.99	84.82
<i>FA</i>	18.70	23.94	18.28	19.05	15.27
<i>mg#</i>	0.71	0.64	0.72	0.70	0.76

ANEXO: Análises de microsonda de cristais de olivina da amostra BYA 127, lamprofiros de Ró Grande

		AMOSTR BYA 127																		
		Bordas							Intermedios				Núcleos							
		1	2	3	4	5	6	7	8	2	4	1	2	3	4	5	6	7	8	9
Amostra		BYA127.5 .OLVRIM 478	BYA127.3 B.OLVRIM .481	BYA127.1 A.OLVRIM .501	BYA127.1 B.OLVRI M.505	BYA127.2 B.OLVRI M.508	BYA127.2 A.OLVRIM .519	BYA127.6 A.OLVRIM .545	BYA127.6 .OLVRIM 557	BYA127.3B .OLVINTER .482	BYA127.1 B.OLVCOR E.503	BYA127.5 OLVCORE .479	BYA127.3B OLVINTCOR E.483	BYA127. 1A.OLV. 500	BYA127.1B.O LVMORECOR E.504	BYA127.2B .OLVCORE .509	BYA127.2 A.OLVCOR E.518	BYA127.6 A.OLVCOR E.543	BYA127.6 OLVCORE .556	BYA127.1. OLVCORE 506
<i>SiO₂</i>		39.52	40.32	40.66	39.47	40.11	39.95	39.52	39.44	40.21	39.59	39.38	40.22	40.22	39.66	40.43	40.37	39.64	39.87	39.01
<i>TiO₂</i>		0.00	0.05	0.02	0.07	0.04	0.05	0.06	0.08	0.01	0.01	0.05	0.02	0.00	0.01	0.03	0.03	0.07	0.00	0.01
<i>Al₂O₃</i>		0.02	0.08	0.02	0.01	0.05	0.53	0.03	0.04	0.02	0.03	0.06	0.05	0.06	0.05	0.07	0.05	0.02	0.03	0.07
<i>FeO_(t)</i>		15.03	15.00	12.60	15.08	14.60	14.93	14.84	14.93	13.05	14.11	19.14	13.20	12.90	13.10	13.36	12.96	13.40	12.81	13.87
<i>MnO</i>		0.26	0.22	0.14	0.25	0.24	0.17	0.27	0.27	0.15	0.12	0.24	0.18	0.15	0.12	0.14	0.11	0.12	0.11	0.15
<i>MgO</i>		43.70	44.02	46.84	44.12	44.50	43.98	44.16	43.92	45.56	44.79	41.16	45.21	45.41	45.68	46.17	45.88	44.88	45.83	45.20
<i>CaO</i>		0.38	0.42	0.19	0.40	0.30	0.34	0.43	0.41	0.21	0.26	0.11	0.24	0.18	0.23	0.22	0.24	0.22	0.16	0.23
<i>Cr₂O₃</i>		0.05	0.08	0.07	0.09	0.02	0.02	0.09	0.07	0.04	0.04	0.01	0.03	0.03	0.07	0.00	0.07	0.01	0.03	0.04
<i>NiO</i>		0.20	0.19	0.37	0.15	0.20	0.15	0.15	0.17	0.32	0.21	0.16	0.31	0.26	0.20	0.23	0.42	0.20	0.46	0.29
<i>K₂O</i>		0.01	0.00	0.01	0.00	0.02	0.09	0.00	0.00	0.01	0.01	0.00	0.01	0.00	0.01	0.00	0.02	0.00	0.00	0.00
Total		99.16	100.38	100.91	99.63	100.06	100.21	99.53	99.32	99.58	99.17	100.30	99.46	99.21	99.13	100.65	100.15	98.56	99.30	98.87
FO		84.03	84.24	86.98	84.14	84.58	84.24	84.39	84.22	86.27	85.18	79.22	86.04	86.32	86.30	86.16	86.49	85.80	86.53	85.43
FA		16.21	16.11	13.13	16.13	15.57	16.04	15.91	16.06	13.86	15.05	20.67	14.10	13.76	13.88	13.99	13.71	14.37	13.57	14.71
mg#		0.74	0.75	0.79	0.75	0.75	0.75	0.75	0.75	0.78	0.76	0.68	0.77	0.78	0.78	0.78	0.78	0.77	0.78	0.77



CLINOPIROXÊNIO

ANEXO: Análises de microsonda de cristais de clinopiroxênio da amostra P8, lamprofiros de Río Grande

<i>Fenocristais brancos</i>														
Amostra	Bordas							Núcleos						
	1	2	3	4	5	6	7	1	2	3	4	5	6	7
	P8A.5.CPX BORDE.A	P8A.13.CPX RIM.18	P8A.13.CPX RIM.25	P8G.4.CPX RIM.712	P8G.9.CPX RIM.719	P8H.7.CPX RIMXENOX .668	P8H.4.CPX RIM.617	P8A.5.CPX CORE.B	P8A.13.CPX CORE.17	P8A.13.CPX CORE.24	P8G.4.CPX CORE.711	P8G.9.CPX CORE.718	P8H.7.CPX. XENOX.667	P8H.4.CPX. CORE.616
<i>SiO₂</i>	46.85	47.85	50.13	48.26	49.32	49.47	49.69	50.81	50.04	49.19	52.10	49.67	48.36	49.03
<i>TiO₂</i>	2.34	1.50	1.42	1.60	1.54	1.31	1.27	1.09	1.12	1.45	0.87	1.34	1.19	1.36
<i>Al₂O₃</i>	6.73	7.08	5.25	4.80	4.82	5.29	5.23	5.68	5.14	5.37	3.96	5.07	5.51	5.98
<i>FeO_(t)</i>	7.82	5.57	8.34	7.79	7.93	7.86	7.41	6.73	5.36	8.07	6.80	9.39	7.70	7.03
<i>MnO</i>	0.20	0.09	0.16	0.16	0.14	0.13	0.12	0.12	0.09	0.19	0.15	0.23	0.10	0.16
<i>MgO</i>	11.95	13.30	11.87	12.73	13.27	12.65	12.84	13.53	14.85	11.93	14.15	11.84	12.74	12.45
<i>CaO</i>	22.66	21.28	20.67	22.64	22.58	20.68	21.42	20.71	21.66	20.99	20.94	19.96	20.70	21.54
<i>Cr₂O₃</i>	0.05	0.67	0.02	0.01	0.00	0.03	0.10	0.30	0.48	0.02	0.55	0.00	0.08	0.12
<i>Na₂O</i>	0.89	0.80	1.66	0.53	0.80	1.49	1.49	1.37	0.35	1.60	1.24	2.07	1.32	1.46
<i>K₂O</i>	0.02	0.02	0.01	0.00	0.00	0.00	0.03	0.02	0.02	0.01	0.00	0.04	0.02	0.01
<i>Total</i>	100.07	98.41	99.82	98.83	100.93	99.31	100.08	100.58	99.15	99.32	100.95	100.28	98.21	99.61
Fórmula estrutural baseada em 6 oxigênios														
Si	1.75	1.79	1.86	1.82	1.82	1.84	1.84	1.86	1.85	1.84	1.90	1.84	1.82	1.82
Al^{IV}	0.25	0.21	0.14	0.18	0.18	0.16	0.16	0.14	0.15	0.16	0.10	0.16	0.18	0.18
Fe³⁺	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Sítio T	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00
Al^{VI}	0.05	0.10	0.09	0.03	0.03	0.08	0.07	0.10	0.08	0.08	0.07	0.06	0.07	0.08
Ti	0.07	0.04	0.04	0.05	0.04	0.04	0.04	0.03	0.03	0.04	0.02	0.04	0.03	0.04
Fe³⁺	0.14	0.06	0.08	0.09	0.12	0.11	0.13	0.07	0.02	0.12	0.05	0.17	0.14	0.12
Fe²⁺	0.11	0.12	0.18	0.15	0.12	0.13	0.10	0.14	0.15	0.14	0.16	0.12	0.11	0.10
Mn	0.01	0.00	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.01	0.00	0.01	0.00	0.00
Mg	0.67	0.74	0.66	0.72	0.73	0.70	0.71	0.74	0.82	0.67	0.77	0.66	0.72	0.69
Cr	0.00	0.02	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.00	0.02	0.00	0.00	0.00
Sítio O	1.03	1.09	1.06	1.05	1.05	1.07	1.04	1.09	1.11	1.04	1.09	1.06	1.07	1.04
Ca	0.91	0.85	0.82	0.92	0.89	0.83	0.85	0.81	0.86	0.84	0.82	0.79	0.84	0.86
Na	0.06	0.06	0.12	0.04	0.06	0.11	0.11	0.10	0.03	0.12	0.09	0.15	0.10	0.10
K	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Sítio C	0.97	0.91	0.94	0.95	0.95	0.93	0.96	0.91	0.89	0.96	0.91	0.94	0.93	0.96
Wo	49.92	48.21	47.29	48.76	47.81	46.56	47.53	46.23	46.57	47.83	45.58	45.60	46.59	48.55
En	36.63	41.95	37.81	38.15	39.08	39.63	39.65	42.04	44.44	37.83	42.86	37.66	39.89	39.07
Fs	13.45	9.84	14.90	13.09	13.10	13.81	12.83	11.72	8.99	14.35	11.56	16.74	13.52	12.38
mg##	73.14	80.99	71.73	74.45	74.89	74.16	75.55	78.20	83.18	72.50	78.76	69.22	74.68	75.94

*mg# = 100*Mg/(Mg+Fetot)

ANEXO: Análises de microsonda de cristais de clinopiroxênio da amostra P8, lamprofiros de Ríó Grande

<i>Meso e microfenocristais brancos</i>													
	8	9	10	11	12	13	14	15	16	17	18	19	
<i>Amostra</i>	P8A.5.CPXCORE.12	P8A.13.CPX.RIM.19	P8D.2.CPX.432	P8D.4.CPX.440	P8D.3.CPX.443	P8D.3.CPX.444	P8E.1A.CPX.300	P8E.3.CPX.312	P8E.7.CPX.CORE.344	P8E.13.CPX.CORE.350	P8G.1Y2.CPX.698	P8G.1Y2.CPX.699	
<i>SiO₂</i>	46.59	49.55	49.54	50.23	50.59	50.22	53.49	51.26	51.86	48.32	53.27	51.41	
<i>TiO₂</i>	2.25	1.26	1.55	1.32	1.14	0.86	0.18	1.08	0.66	1.70	0.25	0.42	
<i>Al₂O₃</i>	8.05	5.33	6.80	5.71	4.78	3.25	0.98	4.63	3.14	8.30	1.36	4.54	
<i>FeO_(t)</i>	5.99	5.41	8.25	7.29	6.96	6.71	8.78	6.77	8.98	6.38	7.63	7.86	
<i>MnO</i>	0.02	0.04	0.11	0.14	0.11	0.18	0.33	0.16	0.24	0.12	0.21	0.16	
<i>MgO</i>	12.57	14.55	11.67	13.07	12.95	13.66	13.88	13.47	11.32	12.73	13.43	12.65	
<i>CaO</i>	21.65	21.65	20.64	21.17	21.35	22.97	20.43	21.35	19.99	20.31	20.90	21.60	
<i>Cr₂O₃</i>	0.51	0.33	0.05	0.08	0.16	0.00	0.01	0.12	0.13	0.21	0.20	0.00	
<i>Na₂O</i>	0.80	0.63	1.74	1.25	1.41	0.66	1.67	1.02	2.52	1.38	2.06	1.15	
<i>K₂O</i>	0.00	0.01	0.02	0.03	0.00	0.01	0.01	0.00	0.00	0.00	0.02	0.00	
<i>Total</i>	98.66	98.96	100.70	100.58	99.78	98.84	100.15	100.42	99.27	99.69	99.81	100.02	
Fórmula estrutural baseada em 6 oxigênios													
<i>Si</i>	1.75	1.84	1.82	1.85	1.87	1.88	1.98	1.89	1.94	1.78	1.97	1.90	
<i>Al^{IV}</i>	0.25	0.16	0.18	0.15	0.13	0.12	0.02	0.11	0.06	0.22	0.03	0.10	
<i>Fe³⁺</i>	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	
<i>Sítio T</i>	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	
<i>Al^{VI}</i>	0.10	0.07	0.12	0.09	0.08	0.03	0.02	0.09	0.08	0.14	0.03	0.10	
<i>Ti</i>	0.06	0.04	0.04	0.04	0.03	0.02	0.00	0.03	0.02	0.05	0.01	0.01	
<i>Fe³⁺</i>	0.07	0.05	0.10	0.08	0.08	0.09	0.11	0.04	0.13	0.07	0.13	0.05	
<i>Fe²⁺</i>	0.12	0.11	0.16	0.15	0.14	0.12	0.16	0.17	0.15	0.13	0.10	0.19	
<i>Mn</i>	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.00	0.01	0.00	
<i>Mg</i>	0.70	0.81	0.64	0.72	0.71	0.76	0.76	0.74	0.63	0.70	0.74	0.70	
<i>Cr</i>	0.02	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.00	
<i>Sítio O</i>	1.07	1.09	1.06	1.08	1.05	1.03	1.07	1.07	1.02	1.10	1.02	1.06	
<i>Ca</i>	0.87	0.86	0.81	0.83	0.85	0.92	0.81	0.86	0.80	0.80	0.83	0.86	
<i>Na</i>	0.06	0.05	0.12	0.09	0.10	0.05	0.12	0.07	0.18	0.10	0.15	0.08	
<i>K</i>	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	
<i>Sítio C</i>	0.93	0.91	0.94	0.92	0.95	0.97	0.93	0.93	0.98	0.90	0.98	0.94	
<i>Wo</i>	49.41	46.94	47.65	47.00	47.65	48.65	43.85	47.50	46.76	47.23	45.89	47.65	
<i>En</i>	39.92	43.90	37.48	40.37	40.22	40.25	41.44	40.96	36.84	41.19	41.04	38.82	
<i>Fs</i>	10.67	9.16	14.87	12.63	12.13	11.09	14.70	11.54	16.40	11.58	13.08	13.53	
<i>mg#*</i>	78.91	82.74	71.60	76.17	76.83	78.40	73.81	78.01	69.19	78.05	75.83	74.15	

*mg# = 100 * Mg / (Mg + Fetot)

ANEXO: Análises de microsonda de cristais de clinopiroxênio verdes da amostra P8, lamprofiros de Rfo Granc

Fenocristais verdes

Amostra	Bordas			Núcleos		
	1	2	3	1	2	3
	P8D.7.CPX RIM.437	P8E.4.CPX RIM.316	P8E.10.CPX RIM.362	P8D.CPX CORE.438	P8E.4.CPX. CORE315	P8E.10.CPX CORE.363
SiO ₂	47.36	46.44	48.00	50.72	51.34	51.91
TiO ₂	2.31	3.06	1.52	1.20	0.80	0.54
Al ₂ O ₃	6.68	6.84	5.50	4.20	3.21	3.27
FeO _(v)	7.75	6.37	7.67	10.71	10.24	9.53
MnO	0.11	0.06	0.14	0.17	0.22	0.19
MgO	12.14	13.13	12.46	10.89	11.12	11.21
CaO	22.64	23.92	22.81	19.47	20.30	20.06
Cr ₂ O ₃	0.00	0.00	0.02	0.03	0.03	0.00
Na ₂ O	1.09	0.44	0.82	2.38	2.35	2.51
K ₂ O	0.00	0.01	0.00	0.01	0.02	0.00
Total	100.62	100.77	99.41	100.29	100.19	99.80
Fórmula estrutural baseada em 6 oxigênios						
Si	1.75	1.72	1.80	1.89	1.91	1.93
Al ^{IV}	0.25	0.28	0.20	0.11	0.09	0.07
Fe ³⁺	0.00	0.00	0.00	0.00	0.00	0.00
Sítio T	2.00	2.00	2.00	2.00	2.00	2.00
Al ^{VI}	0.05	0.01	0.04	0.07	0.05	0.08
Ti	0.06	0.09	0.04	0.03	0.02	0.02
Fe ³⁺	0.15	0.13	0.13	0.14	0.16	0.14
Fe ²⁺	0.09	0.07	0.11	0.19	0.16	0.16
Mn	0.00	0.00	0.00	0.01	0.01	0.01
Mg	0.67	0.72	0.70	0.60	0.62	0.62
Cr	0.00	0.00	0.00	0.00	0.00	0.00
Sítio O	1.02	1.02	1.03	1.05	1.02	1.02
Ca	0.90	0.95	0.92	0.78	0.81	0.80
Na	0.08	0.03	0.06	0.17	0.17	0.18
K	0.00	0.00	0.00	0.00	0.00	0.00
Sítio C	0.98	0.98	0.97	0.95	0.98	0.98
Wo	49.66	50.72	49.44	45.29	46.38	46.54
En	37.06	38.74	37.59	35.26	35.37	36.19
Fs	13.27	10.55	12.97	19.45	18.26	17.27
mg#*	73.63	78.60	74.35	64.45	65.95	67.70

*mg# = 100*Mg/(Mg+Fetot)

ANEXO: Análises de microsonda de cristais e megacristais de clinopiroxênio da amostra BYA 123, lamprofiros de Rio Grande

Microfenocristais					
	Cristais na matriz				Inclusão
	1	2	3	4	5
<i>Amostra</i>	BYA123.1.CPXMZ. 391	BYA123.2.CPXMZ. 397	BYA123.2.CPXMZ. ACIC.403	BYA123.2.CPXMZ. ACI.404	BYA123.1.INCLIN CPXGRANDE.377
<i>SiO₂</i>	48.09	40.91	44.14	46.86	45.26
<i>TiO₂</i>	2.17	4.62	3.00	2.44	2.77
<i>Al₂O₃</i>	5.38	10.10	9.69	5.50	8.24
<i>FeO_(t)</i>	6.62	8.60	8.51	7.29	6.47
<i>MnO</i>	0.11	0.10	0.07	0.05	0.07
<i>MgO</i>	13.61	10.36	12.36	12.77	12.06
<i>CaO</i>	23.33	23.11	20.64	23.22	22.98
<i>Cr₂O₃</i>	0.17	0.00	0.00	0.00	0.47
<i>Na₂O</i>	0.41	0.35	0.68	0.51	0.53
<i>K₂O</i>	0.02	0.04	0.47	0.03	0.02
<i>Total</i>	100.32	98.84	100.31	99.16	99.38
Fórmula estrutural baseada em 6 oxigênios					
Si	1.78	1.57	1.64	1.76	1.70
Al^{IV}	0.22	0.43	0.36	0.24	0.30
Fe³⁺	0.00	0.00	0.00	0.00	0.00
Sítio T	2.00	2.00	2.00	2.00	2.00
Al^{VI}	0.02	0.02	0.07	0.01	0.06
Ti	0.06	0.13	0.08	0.07	0.08
Fe³⁺	0.10	0.18	0.19	0.13	0.11
Fe²⁺	0.10	0.10	0.07	0.10	0.10
Mn	0.00	0.00	0.00	0.00	0.00
Mg	0.75	0.59	0.69	0.72	0.67
Cr	0.00	0.00	0.00	0.00	0.01
Sítio O	1.04	1.02	1.11	1.02	1.04
Ca	0.93	0.95	0.82	0.94	0.92
Na	0.03	0.03	0.05	0.04	0.04
K	0.00	0.00	0.02	0.00	0.00
Sítio C	0.96	0.98	0.89	0.98	0.96
Wo	49.19	52.24	46.40	49.75	51.29
En	39.92	32.58	38.67	38.06	37.44
Fs	10.89	15.18	14.93	12.18	11.27
mg#*	78.57	68.22	72.14	75.75	76.87

*mg# =100*Mg/(Mg+Fetot)

ANEXO: Análises de microsonda de cristais e megacristais de clinopiroxênio da amostra BYA 124a, lamprofiros de Ríó Grande

	<i>Mesofenocristais</i>			
	1	2	3	4
<i>Amostra</i>	BYA124A.1.CPX.201	BYA124A.2.CPX.207	BYA124A.4.CPX.231	BYA124A.4.CPX.239
<i>SiO₂</i>	50.83	46.12	51.40	49.39
<i>TiO₂</i>	1.25	3.29	0.85	1.38
<i>Al₂O₃</i>	6.16	7.48	5.69	7.22
<i>FeO_(t)</i>	8.05	7.09	5.17	5.14
<i>MnO</i>	0.06	0.08	0.07	0.03
<i>MgO</i>	13.45	12.54	15.87	14.51
<i>CaO</i>	18.88	23.30	18.92	20.50
<i>Cr₂O₃</i>	0.04	0.03	0.72	0.36
<i>Na₂O</i>	1.51	0.58	0.94	1.14
<i>K₂O</i>	0.01	0.02	0.02	0.00
<i>Total</i>	100.43	100.97	99.70	100.03
Fórmula estrutural baseada em 6 oxigênios				
Si	1.86	1.70	1.88	1.81
Al^{IV}	0.14	0.30	0.12	0.19
Fe³⁺	0.00	0.00	0.00	0.00
Sítio T	2.00	2.00	2.00	2.00
Al^{VI}	0.13	0.03	0.12	0.12
Ti	0.03	0.09	0.02	0.04
Fe³⁺	0.04	0.12	0.00	0.07
Fe²⁺	0.20	0.09	0.16	0.09
Mn	0.00	0.00	0.00	0.00
Mg	0.74	0.69	0.86	0.79
Cr	0.00	0.00	0.02	0.01
Sítio O	1.15	1.04	1.19	1.12
Ca	0.74	0.92	0.74	0.80
Na	0.11	0.04	0.07	0.08
K	0.00	0.00	0.00	0.00
Sítio C	0.85	0.96	0.81	0.88
Wo	43.03	50.34	42.00	45.87
En	42.65	37.71	49.05	45.16
Fs	14.32	11.95	8.96	8.97
mg#*	74.86	75.93	84.56	83.43

*mg# = 100*Mg/(Mg+Fetot)

ANEXO: Análises de microsonda de cristais de clinopiroxênio da amostra BHU 131, lamprofiros de Ríó Grande

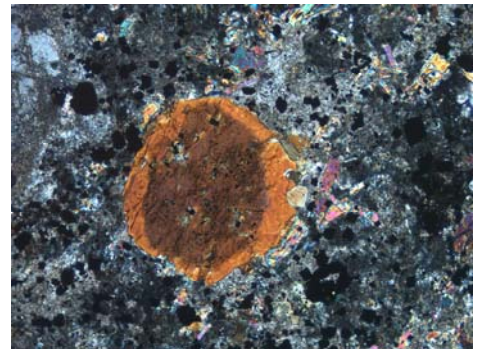
Mesofenocristais														
<i>Amostra</i>	Bordas						Intermedios		Núcleos					
	1	2	3	4	5	6	4	5	1	2	3	4	5	6
	BHU131.5A.CPX RIM.65	BHU131.5A.CPX RIM.67	BHU131.12.CPX RIM.74	BHU131.12.CP XRIM.75	BHU131.1.CPX MZRIM.105	BHU131.9.CPX RIM.128	BHU131.12.CPX INTER.76	BHU131.1.CP XINTER.106	BHU131.5A. CPXCORE.64	BHU131.5A.CP XCORE.68	BHU131.12.CP XCORE.73	BHU131.12.C PXCORE.77	BHU131.1.CP XCORE.108	BHU131.9.CPX .127
<i>SiO₂</i>	49.03	47.44	44.80	44.21	47.98	40.55	50.03	44.97	46.46	46.89	45.10	49.42	48.17	46.37
<i>TiO₂</i>	1.71	2.60	3.60	4.43	2.78	5.51	1.86	3.49	2.96	2.67	3.44	1.75	2.73	3.20
<i>Al₂O₃</i>	3.35	4.55	7.09	8.04	4.63	9.92	3.36	6.97	6.00	5.64	6.80	3.45	4.59	5.88
<i>FeO_(t)</i>	5.43	6.71	7.43	7.48	6.79	8.46	5.56	7.14	6.33	6.12	6.87	5.91	6.81	6.50
<i>MnO</i>	0.12	0.14	0.11	0.03	0.08	0.09	0.11	0.10	0.07	0.08	0.08	0.07	0.10	0.09
<i>MgO</i>	14.25	13.34	12.21	11.98	13.42	10.51	14.41	12.08	12.78	12.93	12.42	14.36	13.76	12.98
<i>CaO</i>	23.36	23.70	23.46	23.39	23.91	22.75	23.59	23.35	23.15	23.40	22.79	22.94	23.63	23.10
<i>Cr₂O₃</i>	0.39	0.00	0.02	0.34	0.01	0.02	0.47	0.18	0.63	0.60	0.17	0.24	0.13	0.39
<i>Na₂O</i>	0.42	0.48	0.32	0.54	0.72	0.80	0.40	0.25	0.22	0.43	0.34	0.48	0.39	0.42
<i>K₂O</i>	0.00	0.00	0.02	0.02	0.01	0.01	0.00	0.01	0.01	0.00	0.00	0.02	0.01	0.00
<i>Total</i>	98.37	99.45	99.51	100.99	100.90	99.37	100.02	98.88	98.84	99.06	98.31	98.95	100.72	99.26
Fórmula estrutural baseada em 6 oxigênios														
<i>Si</i>	1.84	1.78	1.69	1.64	1.77	1.54	1.85	1.70	1.75	1.76	1.71	1.85	1.78	1.74
<i>Al^{IV}</i>	0.15	0.20	0.31	0.35	0.20	0.44	0.15	0.30	0.25	0.24	0.29	0.15	0.20	0.26
<i>Fe³⁺</i>	0.01	0.02	0.00	0.00	0.03	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.00
<i>Sítio T</i>	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00
<i>Al^{VI}</i>	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.02	0.01	0.02	0.00	0.00	0.00
<i>Ti</i>	0.05	0.07	0.10	0.12	0.08	0.16	0.05	0.10	0.08	0.08	0.10	0.05	0.08	0.09
<i>Fe³⁺</i>	0.08	0.11	0.13	0.14	0.12	0.20	0.06	0.09	0.05	0.09	0.09	0.08	0.09	0.09
<i>Fe²⁺</i>	0.09	0.08	0.10	0.09	0.06	0.05	0.11	0.13	0.15	0.11	0.13	0.10	0.10	0.11
<i>Mn</i>	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
<i>Mg</i>	0.80	0.75	0.69	0.66	0.74	0.60	0.80	0.68	0.72	0.72	0.70	0.80	0.76	0.73
<i>Cr</i>	0.01	0.00	0.00	0.01	0.00	0.00	0.01	0.01	0.02	0.02	0.01	0.01	0.00	0.01
<i>Sítio O</i>	1.03	1.01	1.03	1.03	1.00	1.01	1.04	1.03	1.05	1.03	1.05	1.05	1.03	1.04
<i>Ca</i>	0.94	0.95	0.95	0.93	0.95	0.93	0.94	0.95	0.94	0.94	0.93	0.92	0.94	0.93
<i>Na</i>	0.03	0.03	0.02	0.04	0.05	0.06	0.03	0.02	0.02	0.03	0.02	0.03	0.03	0.03
<i>K</i>	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
<i>Sítio C</i>	0.97	0.99	0.97	0.97	1.00	0.99	0.96	0.97	0.95	0.97	0.95	0.95	0.97	0.96
<i>Wo</i>	49.26	49.89	50.72	50.96	49.93	51.73	49.16	51.06	50.46	50.69	50.15	48.25	49.12	49.95
<i>En</i>	41.81	39.08	36.74	36.33	39.00	33.25	41.79	36.76	38.78	38.97	38.05	42.04	39.82	39.07
<i>Fs</i>	8.93	11.03	12.53	12.71	11.07	15.01	9.05	12.18	10.77	10.34	11.80	9.71	11.06	10.97
<i>mg#*</i>	82.39	77.99	74.57	74.08	77.90	68.90	82.20	75.11	78.27	79.03	76.33	81.24	78.27	78.07

*mg# = 100*Mg/(Mg+Fetot)

ANEXO: Análises de microsonda de microcristais de clinopiroxênio da amostra BHU 131, lamprofiros de Rio Grande.

Amostra	Matriz						Glomerófiro
	5	6	7	8	9	10	11
	BHU131.11.CPXMZ.	BHU131.6.CPXMZ.	BHU131.8.CPXMZ.	BHU131.5A.CPX.	BHU131.1.CPXMZ.	BHU131.5A.CPXRIM.	BHU131.4.CPXGLOME.
	143	72	121	66	104	69	117
<i>SiO</i> ₂	45.36	49.04	47.94	46.80	45.30	44.21	47.62
<i>TiO</i> ₂	2.96	2.09	2.19	2.75	3.64	4.15	2.42
<i>Al</i> ₂ <i>O</i> ₃	6.24	4.17	3.98	4.59	7.22	7.48	5.03
<i>FeO</i> _(t)	6.66	6.24	6.10	6.63	7.57	7.62	6.77
<i>MnO</i>	0.03	0.09	0.11	0.10	0.10	0.11	0.15
<i>MgO</i>	12.33	14.04	13.85	13.28	11.94	12.06	13.39
<i>CaO</i>	23.24	23.41	23.17	23.34	22.74	23.00	22.75
<i>Cr</i> ₂ <i>O</i> ₃	0.38	0.03	0.36	0.27	0.10	0.25	0.11
<i>Na</i> ₂ <i>O</i>	0.42	0.31	0.20	0.51	0.30	0.62	0.49
<i>K</i> ₂ <i>O</i>	0.00	0.01	0.00	0.01	0.00	0.03	0.05
<i>Total</i>	98.01	99.67	98.14	98.77	99.11	100.14	99.16
Fórmula estrutural baseada em 6 oxigênios							
Si	1.73	1.83	1.82	1.77	1.71	1.66	1.79
Al^{IV}	0.27	0.17	0.18	0.20	0.29	0.33	0.21
Fe³⁺	0.00	0.00	0.01	0.03	0.00	0.01	0.00
Sítio T	2.00	2.00	2.00	2.00	2.00	2.00	2.00
Al^{VI}	0.01	0.01	0.00	0.00	0.03	0.00	0.01
Ti	0.08	0.06	0.06	0.08	0.10	0.12	0.07
Fe³⁺	0.11	0.07	0.06	0.11	0.06	0.15	0.10
Fe²⁺	0.10	0.12	0.12	0.07	0.18	0.08	0.11
Mn	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Mg	0.70	0.78	0.78	0.75	0.67	0.67	0.75
Cr	0.01	0.00	0.01	0.01	0.00	0.01	0.00
Sítio O	1.02	1.04	1.05	1.02	1.06	1.03	1.05
Ca	0.95	0.93	0.94	0.94	0.92	0.92	0.91
Na	0.03	0.02	0.01	0.04	0.02	0.05	0.04
K	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Sítio C	0.98	0.96	0.95	0.98	0.94	0.97	0.95
Wo	50.98	48.96	49.09	49.66	50.24	50.31	48.75
En	37.62	40.86	40.83	39.33	36.71	36.69	39.93
Fs	11.40	10.18	10.08	11.02	13.05	13.00	11.33
mg#*	76.75	80.05	80.20	78.12	73.77	73.83	77.90

*mg# = 100*Mg/(Mg+Fetot)



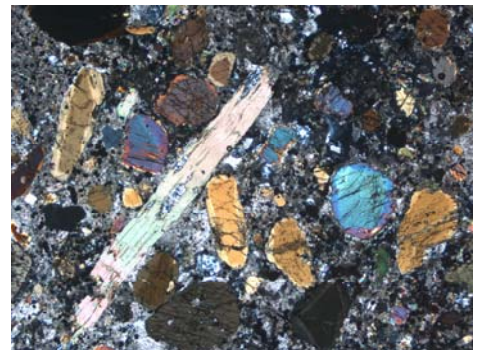
ANFIBÓLIO

Anexo : Análises de microsonda em anfibólios verdes da amostra P8, lamprofiros de Ró Grande.

Amostra	ANFIBÓLIOS VERDES							
	Bordas			Intermedios	Núcleos			Incl
	10	11	12		10	11	12	
	P8A.13.AMPRIM. 16	P8A.SD.AMPRIM. 31	P8A.12.AMPRIM. 53	P8A.SD.AMPIN TER.32	P8A.13.AMPCOR E.15	P8A.SD.AMPCOR E.33	P8A.12.AMPCOR E.54	P8G.8.INCAMP .732
SiO ₂	40.13	41.49	40.67	41.30	41.01	42.77	42.46	39.47
TiO ₂	3.96	3.46	3.50	1.22	1.24	1.40	1.70	3.79
Al ₂ O ₃	12.97	11.28	12.91	14.12	14.45	13.38	12.41	13.24
Fe ₂ O ₃	0.56	0.31	0.59	3.12	3.07	2.52	3.18	1.06
FeO	10.09	13.37	9.54	8.92	10.70	9.46	9.52	8.59
Cr ₂ O ₃	0.01	0.07	0.04	0.01	0.00	0.00	0.02	0.09
MnO	0.14	0.24	0.11	0.23	0.25	0.30	0.25	0.12
MgO	13.69	11.48	13.91	12.54	11.13	12.80	12.73	14.11
CaO	11.94	11.24	12.00	10.14	9.63	9.76	9.57	11.73
Na ₂ O	2.50	3.13	2.60	3.74	4.02	3.98	4.29	2.30
K ₂ O	1.90	1.86	1.99	1.29	1.06	1.33	0.96	2.01
Total	97.89	97.94	97.84	96.60	96.55	97.69	97.08	96.52
Fórmula estrutural baseada em 23 oxigênios								
Si	5.95	6.22	6.01	6.15	6.14	6.28	6.29	5.90
AlIV	2.05	1.78	1.99	1.85	1.86	1.72	1.71	2.10
Sum T	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00
AlVI	0.21	0.22	0.26	0.62	0.69	0.60	0.46	0.24
Fe3+	0.06	0.04	0.07	0.35	0.35	0.28	0.35	0.12
Ti	0.44	0.39	0.39	0.14	0.14	0.15	0.19	0.43
Cr	0.00	0.01	0.00	0.00	0.00	0.00	0.00	0.01
Mg	3.03	2.57	3.07	2.78	2.49	2.80	2.81	3.15
Fe2+	1.25	1.68	1.18	1.11	1.33	1.16	1.18	1.06
Zn	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Mn	0.01	0.03	0.01	0.00	0.00	0.00	0.00	0.00
Sum C	5.00	4.93	4.98	5.00	5.00	5.00	5.00	5.00
Mg	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Fe2+	0.00	0.00	0.00	0.00	0.01	0.00	0.00	0.02
Zn	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Mn	0.01	0.00	0.00	0.03	0.03	0.04	0.03	0.02
Ca	1.90	1.81	1.90	1.62	1.54	1.54	1.52	1.88
Na	0.09	0.19	0.10	0.36	0.42	0.43	0.45	0.09
Sum B	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00
Na	0.63	0.72	0.65	0.72	0.75	0.71	0.78	0.58
K	0.36	0.36	0.38	0.24	0.20	0.25	0.18	0.38
Sum A	0.99	1.07	1.02	0.97	0.95	0.96	0.96	0.96
(Ca+Na)B	1.99	2.00	2.00	1.97	1.96	1.96	1.97	1.97
NaB	0.09	0.19	0.10	0.36	0.42	0.43	0.45	0.09
CaB	1.90	1.81	1.90	1.62	1.54	1.54	1.52	1.88
(Na+K)A	0.99	1.07	1.02	0.97	0.95	0.96	0.96	0.96
Si	5.95	6.22	6.01	6.15	6.14	6.28	6.29	5.90
Mg/(Mg+Fe2+)	0.71	0.60	0.72	0.71	0.65	0.71	0.70	0.75

ANEXO: Análises de microsonda em anfibólios (amostra P8) presentes em xenólitos, lamprofiros de Rio Grand

	P8C.9.CORONA. 157	P8C.11.AMP. 197	P8E.9.INCLBLUE XEN.336	P8E.9.AMP XEN.337	P8D.5.AMPBLUEX EN.339	P8D.5.AMPC ONSEC.340	P8E.7.AMPG REEN.348	P8E.7.AMPH RIM.349	P8E.14.AMP CORE.353	P8H.8.AMPG REEN.604	P8H.5.AMPX EN.642	P8H.6.AMPX EN.662	P8H.6.AMPX EN.664	P8G.1Y2.AMP CORE.701	P8G.1Y2.AMP RIM.702	P8G.8.AMP CORE.727	P8G.8.AMP RIM.728
<i>SiO₂</i>	41.85	41.33	44.01	41.42	40.88	40.98	41.51	41.37	42.44	42.22	43.53	44.24	44.82	41.43	41.01	41.15	40.84
<i>TiO₂</i>	0.01	1.68	2.90	3.80	4.05	4.13	0.89	3.73	1.31	1.47	2.97	2.33	2.22	1.51	3.54	3.97	3.83
<i>Al₂O₃</i>	19.61	17.95	10.19	12.81	12.24	12.34	15.17	12.43	14.71	12.65	11.57	10.11	9.18	13.83	12.79	11.93	13.07
<i>Fe₂O₃</i>	4.35	1.58	1.17	0.00	0.41	0.00	3.35	0.00	2.15	2.58	0.19	2.04	2.74	1.67	1.13	1.23	0.22
<i>FeO</i>	4.23	7.01	8.16	8.91	14.66	15.25	9.49	10.32	9.91	11.48	8.47	10.45	9.62	10.57	9.76	11.68	9.16
<i>Cr₂O₃</i>	0.00	0.00	0.10	0.05	0.02	0.04	0.01	0.04	0.04	0.00	0.47	0.21	0.25	0.04	0.04	0.04	0.09
<i>MnO</i>	0.09	0.09	0.10	0.11	0.21	0.13	0.22	0.11	0.34	0.23	0.10	0.20	0.15	0.20	0.11	0.14	0.11
<i>MgO</i>	14.44	13.24	15.51	14.22	10.38	10.09	11.69	13.95	12.60	11.44	15.27	13.79	14.52	12.19	14.01	12.31	14.35
<i>CaO</i>	10.54	10.98	10.64	11.41	10.43	10.39	9.86	12.43	10.42	9.34	11.39	10.10	10.01	10.55	11.75	10.40	11.84
<i>Na₂O</i>	3.78	3.01	3.30	2.71	2.89	3.32	4.31	2.14	3.78	4.04	3.29	4.20	4.26	3.45	2.56	2.78	2.58
<i>K₂O</i>	0.24	1.04	1.40	1.94	1.84	1.84	1.25	1.88	1.03	1.24	1.25	1.15	1.10	1.37	1.93	2.00	1.99
<i>Total</i>	99.13	97.89	97.46	97.37	98.01	98.51	97.73	98.40	98.72	96.70	98.51	98.82	98.86	96.80	98.64	97.63	98.05
Fórmula estrutural baseada em 23 oxigênios																	
Si	5.89	5.95	6.43	6.10	6.14	6.14	6.12	6.08	6.17	6.33	6.30	6.46	6.53	6.18	6.02	6.14	6.00
AlIV	2.11	2.05	1.57	1.90	1.86	1.86	1.88	1.92	1.83	1.67	1.70	1.54	1.47	1.82	1.98	1.86	2.00
Sum T	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00
AlVI	1.14	1.00	0.19	0.33	0.31	0.32	0.75	0.23	0.69	0.56	0.28	0.20	0.10	0.61	0.23	0.24	0.27
Fe3+	0.46	0.17	0.13	0.00	0.05	0.00	0.37	0.00	0.24	0.29	0.02	0.22	0.30	0.19	0.13	0.14	0.02
Ti	0.00	0.18	0.32	0.42	0.46	0.47	0.10	0.41	0.14	0.17	0.32	0.26	0.24	0.17	0.39	0.45	0.42
Cr	0.00	0.00	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.05	0.02	0.03	0.00	0.00	0.01	0.01
Mg	3.03	2.84	3.38	3.12	2.33	2.25	2.57	3.06	2.73	2.56	3.29	3.00	3.15	2.71	3.07	2.74	3.14
Fe2+	0.38	0.81	0.97	1.10	1.84	1.91	1.17	1.27	1.19	1.43	1.03	1.28	1.17	1.32	1.18	1.44	1.13
Zn	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Mn	0.00	0.00	0.00	0.01	0.01	0.02	0.03	0.01	0.00	0.00	0.01	0.02	0.00	0.00	0.00	0.00	0.01
Sum C	5.00	5.00	5.00	4.99	5.00	4.97	4.99	4.99	5.00	5.00	5.00	5.00	5.00	5.00	5.00	5.00	5.00
Mg	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Fe2+	0.12	0.04	0.02	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.00	0.00	0.00	0.00	0.01	0.02	0.00
Zn	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Mn	0.01	0.01	0.01	0.00	0.01	0.00	0.00	0.00	0.04	0.03	0.01	0.01	0.01	0.02	0.01	0.02	0.01
Ca	1.59	1.69	1.67	1.80	1.68	1.67	1.56	1.96	1.62	1.50	1.77	1.58	1.56	1.69	1.85	1.66	1.86
Na	0.28	0.26	0.30	0.20	0.31	0.33	0.44	0.04	0.32	0.46	0.23	0.41	0.42	0.29	0.12	0.30	0.13
Sum B	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00
Na	0.75	0.58	0.64	0.57	0.53	0.63	0.79	0.57	0.74	0.71	0.70	0.78	0.78	0.71	0.60	0.51	0.61
K	0.04	0.19	0.26	0.36	0.35	0.35	0.23	0.35	0.19	0.24	0.23	0.21	0.20	0.26	0.36	0.38	0.37
Sum A	0.79	0.77	0.90	0.94	0.89	0.99	1.02	0.92	0.93	0.95	0.93	0.99	0.98	0.97	0.97	0.89	0.98
<i>(Ca+Na)B</i>	1.87	1.95	1.96	2.00	1.99	2.00	2.00	2.00	1.95	1.96	1.99	1.99	1.99	1.98	1.97	1.96	1.99
<i>NaB</i>	0.28	0.26	0.30	0.20	0.31	0.33	0.44	0.04	0.32	0.46	0.23	0.41	0.42	0.29	0.12	0.30	0.13
<i>CaB</i>	1.59	1.69	1.67	1.80	1.68	1.67	1.56	1.96	1.62	1.50	1.77	1.58	1.56	1.69	1.85	1.66	1.86
<i>(Na+K)A</i>	0.79	0.77	0.90	0.94	0.89	0.99	1.02	0.92	0.93	0.95	0.93	0.99	0.98	0.97	0.97	0.89	0.98
<i>Si</i>	5.89	5.95	6.43	6.10	6.14	6.14	6.12	6.08	6.17	6.33	6.30	6.46	6.53	6.18	6.02	6.14	6.00
<i>Mg/(Mg+Fe2+)</i>	0.86	0.77	0.77	0.74	0.56	0.54	0.69	0.71	0.69	0.64	0.76	0.70	0.73	0.67	0.72	0.65	0.74



BIOTITE

ANEXO: Análises de microsonda de cristais de biotita da amostra P8, lamprofiros de Río Grande

Amostra	<i>Cristais</i>									
	Bordas				Intermedio		Núcleos			
	1	2	3	4	3	4	1	2	3	4
	P8A.2.RIM.A	P8G.BTRIM.723	P8G.8.BTRIM.724	P8G.8.BTRIM.729	P8G.8.BTINTER.72	P8G.8.BTINTER.730	P8A.2.BTCORE.B	P8G.BTCORE.722	P8G.8.BTCORE.726	P8G.8.BTCORE.731
SiO ₂	37.61	37.39	35.80	41.94	36.17	39.97	36.57	36.69	36.10	40.15
TiO ₂	3.24	3.60	6.06	1.89	6.29	2.17	5.23	3.71	6.16	2.02
Al ₂ O ₃	14.99	15.03	15.24	12.64	15.27	12.49	14.83	14.56	15.31	12.29
FeO	7.06	7.20	13.62	10.69	13.82	11.08	13.21	13.29	13.78	11.16
Fe ₂ O ₃	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Cr ₂ O ₃	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
MnO	0.22	0.23	0.10	0.10	0.11	0.13	0.04	0.17	0.10	0.10
MgO	20.67	21.36	14.62	19.19	14.30	20.32	15.29	15.97	14.71	20.12
CaO	0.04	0.03	0.00	0.07	0.01	0.04	0.03	0.02	0.03	0.02
Na ₂ O	0.40	0.78	0.60	1.01	0.72	1.26	0.82	1.84	0.60	1.32
K ₂ O	9.12	9.09	9.16	7.68	9.19	8.24	9.23	7.98	9.24	8.28
BaO	1.14	1.02	0.00	0.00	0.00	0.00	0.28	0.27	0.16	0.00
Cl	0.00	0.01	0.02	0.01	0.02	0.02	0.03	0.03	0.02	0.02
H ₂ O	4.09	4.13	4.00	4.15	4.03	4.11	4.01		4.04	4.10
Total	98.57	99.87	99.23	99.36	99.92	99.83	99.57	94.53	100.26	99.55
Fórmula estrutural baseada em 22 oxigênios										
Si	5.51	5.42	5.34	6.04	5.36	5.80	5.44	5.50	5.34	5.85
Al ^{IV}	2.49	2.57	2.66	1.96	2.64	2.14	2.56	2.50	2.66	2.11
Fe ⁺³	0.00	0.01	0.00	0.00	0.00	0.06	0.00	0.00	0.00	0.04
Sum T	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00
Al ^{VI}	0.10	0.00	0.03	0.19	0.03	0.00	0.04	0.07	0.01	0.00
Fe ⁺² _t	0.86	0.86	1.70	1.29	1.71	1.29	1.64	1.67	1.71	1.32
Fe ⁺³	4.51	4.61	3.25	4.12	3.16	4.40	3.39	3.57	3.25	4.37
Ti	0.03	0.03	0.01	0.01	0.01	0.02	0.01	0.02	0.01	0.01
Mg	0.36	0.39	0.68	0.20	0.70	0.24	0.59	0.42	0.69	0.22
Mn	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Sum O	5.86	5.89	5.67	5.82	5.62	5.94	5.66	5.74	5.66	5.92
Ca	0.01	0.00	0.00	0.01	0.00	0.01	0.00	0.00	0.00	0.00
Na	0.11	0.22	0.17	0.28	0.21	0.36	0.24	0.53	0.17	0.37
K	0.85	0.84	0.87	0.71	0.87	0.76	0.88	0.76	0.87	0.77
Ba	0.13	0.12	0.00	0.00	0.00	0.00	0.03	0.03	0.02	0.00
Sum A	1.10	1.18	1.05	1.00	1.08	1.12	1.15	1.33	1.07	1.14
Cations	14.96	15.07	14.72	14.82	14.70	15.06	14.81	15.08	14.73	15.06
Mg/(Mg+Fe ⁺²)	0.84	0.84	0.66	0.76	0.65	0.77	0.67	0.68	0.66	0.77
Fe ⁺² /(Fe ⁺² +Mg)	0.16	0.16	0.34	0.24	0.35	0.23	0.33	0.32	0.34	0.23

Obs: H₂O recalculated by stoichiometry. Fe₂O₃ only recalculated for samples with deficiency in tetrahedral site.

ANEXO: Análises de microsonda de cristais de biotita da amostra P8, lamprofiros de Rió Grande

	Mesofenocristais							Matriz		Inclusões			Borda de anfibólio
	1	2	3	4	5	6	7	8	9	10	11	12	13
<i>Amostra</i>	P8G.8.BTCORE .733	P8A.2.MZBT. C	P8A.5.BTMZ. E	P8A.13.BT.2 3	P8A.12.BT.4 7	P8E.1A.BT. 301	P8G.1Y2.BTCORE. 708	P8H.10.BT.68 9	P8H.1.BT.62 1	P8E.10.BTMZ.3 64	P8A.8D.INCLUSIONBT. 34	P8D.BTINCAMP. 442	P8H.10.BTRIM.68 5
<i>SiO₂</i>	36.23	38.24	37.47	39.27	40.81	39.83	40.10	38.79	37.88	36.47	40.77	38.89	37.41
<i>TiO₂</i>	6.10	2.87	3.19	2.68	1.43	2.59	1.35	2.51	2.42	3.42	1.33	3.13	3.16
<i>Al₂O₃</i>	15.28	14.06	14.30	14.76	11.96	12.91	12.54	14.44	14.15	15.05	14.23	14.16	14.86
<i>FeO</i>	13.48	7.29	7.42	6.61	5.61	10.14	7.82	6.83	7.03	10.65	11.84	11.90	7.16
<i>Fe₂O₃</i>	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
<i>Cr₂O₃</i>	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
<i>MnO</i>	0.03	0.17	0.15	0.19	0.21	0.08	0.12	0.21	0.16	0.15	0.23	0.12	0.16
<i>MgO</i>	14.98	21.12	20.80	21.21	24.21	20.32	22.10	22.56	22.44	20.26	13.02	18.85	20.97
<i>CaO</i>	0.01	0.08	0.04	0.11	0.11	0.09	0.04	0.10	0.10	0.12	10.15	0.01	0.07
<i>Na₂O</i>	0.65	0.61	0.68	0.89	1.25	0.95	1.12	0.91	0.40	0.13	2.88	1.36	0.74
<i>K₂O</i>	9.16	9.37	9.26	9.04	8.58	8.45	8.83	9.00	9.10	9.01	2.12	8.31	8.92
<i>BaO</i>	0.00	1.17	1.03	0.74	0.24	0.47	0.22	0.49	0.41	0.78	0.00	0.00	1.57
<i>Cl</i>	0.03	0.00	0.01	0.00	0.00	0.04	0.03	0.01	0.01	0.02	0.00	0.02	0.00
<i>H₂O</i>	4.04	4.09	4.05	4.16	4.15	4.11	4.09	4.17	4.09	4.08	4.17	4.13	4.09
<i>Total</i>	99.98	99.07	98.37	99.67	98.55	99.99	98.36	100.03	98.19	100.14	100.74	100.89	99.11
Fórmula estrutural baseada em 22 oxigênios													
Si	5.36	5.59	5.52	5.65	5.88	5.77	5.85	5.57	5.54	5.34	5.85	5.62	5.48
Al^{IV}	2.64	2.41	2.48	2.35	2.03	2.20	2.15	2.43	2.44	2.60	2.15	2.38	2.52
Fe⁺³	0.00	0.00	0.00	0.00	0.09	0.03	0.00	0.00	0.02	0.06	0.00	0.00	0.00
Sum T	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00
Al^{VI}	0.02	0.01	0.00	0.15	0.00	0.00	0.00	0.01	0.00	0.00	0.26	0.03	0.04
Fe⁺²_t	1.67	0.89	0.91	0.79	0.58	1.20	0.95	0.82	0.84	1.24	1.42	1.44	0.88
Fe⁺³	3.30	4.60	4.57	4.55	5.20	4.39	4.81	4.83	4.90	4.42	2.79	4.06	4.57
Ti	0.00	0.02	0.02	0.02	0.03	0.01	0.01	0.03	0.02	0.02	0.03	0.01	0.02
Mg	0.68	0.32	0.35	0.29	0.15	0.28	0.15	0.27	0.27	0.38	0.14	0.34	0.35
Mn	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Sum O	5.67	5.84	5.85	5.80	5.96	5.88	5.92	5.95	6.02	6.06	4.64	5.88	5.86
Ca	0.00	0.01	0.01	0.02	0.02	0.01	0.01	0.01	0.02	0.02	1.56	0.00	0.01
Na	0.19	0.17	0.19	0.25	0.35	0.27	0.32	0.25	0.11	0.04	0.80	0.38	0.21
K	0.86	0.87	0.87	0.83	0.79	0.78	0.82	0.82	0.85	0.84	0.19	0.77	0.83
Ba	0.00	0.13	0.12	0.08	0.03	0.05	0.02	0.05	0.05	0.09	0.00	0.00	0.18
Sum A	1.05	1.19	1.19	1.18	1.18	1.12	1.17	1.15	1.03	0.99	2.56	1.15	1.23
Cations	14.72	15.04	15.04	14.98	15.14	15.00	15.09	15.10	15.05	15.05	15.20	15.03	15.09
<i>Mg/(Mg+Fe⁺²)</i>	0.66	0.84	0.83	0.85	0.90	0.78	0.83	0.85	0.85	0.78	0.66	0.74	0.84
<i>Fe⁺²/(Fe⁺²+Mg)</i>	0.34	0.16	0.17	0.15	0.10	0.22	0.17	0.15	0.15	0.22	0.34	0.26	0.16

Obs: H₂O recalculated by stoichiometry. Fe₂O₃ only recalculated for samples with deficiency in tetrahedral site.

ANEXO: Análises de microsonda de cristais de biotita em ocelli e segregados (amostra BHU 131), lamprofiros de Ríó Grande

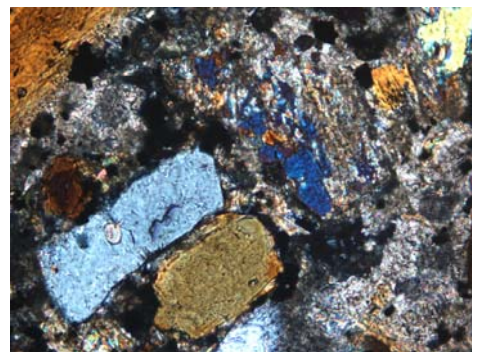
Amostra	em ocelli				em segregados				Matriz
	1	2	3	4	5	6	7	8	9
	BHU131.3.BTANAL.6	BHU131.3.BTANAL.62	BHU131.12.BTCORE.OCELLUS.80	BHU131.12.BTRIM.OCELLUS.81	BHU131.1.BTOCELLA.93	BHU131.11.BTTG.1.40	BHU131.1.BTMZCLARA.96	BHU131.11.BTMZRICAEN.149	BHU131.6.BT.71
SiO_2	34.35	34.01	34.25	35.49	34.29	33.78	34.74	34.25	34.93
TiO_2	7.21	7.43	7.88	6.82	7.71	7.56	6.04	7.85	6.83
Al_2O_3	15.18	15.58	15.51	14.79	15.35	15.37	15.80	15.33	15.61
FeO	13.54	13.62	12.10	12.38	15.29	11.77	13.42	13.23	14.56
Fe_2O_3	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Cr_2O_3	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
MnO	0.15	0.16	0.10	0.10	0.14	0.14	0.17	0.06	0.14
MgO	13.37	13.32	14.86	14.96	13.64	15.01	14.60	13.38	13.85
CaO	0.04	0.09	0.05	0.11	0.18	0.03	0.05	0.12	0.04
Na_2O	0.52	0.61	0.46	0.38	0.60	0.44	0.34	0.75	0.72
K_2O	8.68	8.56	8.71	8.64	8.67	8.51	8.43	8.47	8.35
BaO	1.91	1.52	2.00	0.98	1.18	1.49	1.40	1.39	1.75
Cl	0.01	0.00	0.00	0.02	0.03	0.00	0.01	0.02	0.01
H_2O	3.93	3.94	4.02	3.98	3.99	3.94	3.97	3.95	4.01
Total	98.89	98.83	99.93	98.63	101.05	98.04	98.96	98.79	100.78
Fórmula estrutural baseada em 22 oxigênios									
Si	5.23	5.17	5.13	5.33	5.12	5.13	5.24	5.19	5.21
Al^{IV}	2.72	2.79	2.74	2.62	2.70	2.75	2.76	2.74	2.74
Fe⁺³	0.05	0.05	0.14	0.05	0.18	0.12	0.00	0.08	0.04
Sum T	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00	8.00
Al^{VI}	0.00	0.00	0.00	0.00	0.00	0.00	0.04	0.00	0.00
Fe⁺²	1.67	1.68	1.38	1.50	1.73	1.37	1.69	1.60	1.77
Fe⁺³	3.03	3.02	3.32	3.35	3.04	3.40	3.28	3.02	3.08
Ti	0.02	0.02	0.01	0.01	0.02	0.02	0.02	0.01	0.02
Mg	0.83	0.85	0.89	0.77	0.87	0.86	0.68	0.89	0.77
Mn	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Sum O	5.55	5.57	5.60	5.63	5.65	5.65	5.72	5.52	5.64
Ca	0.01	0.01	0.01	0.02	0.03	0.00	0.01	0.02	0.01
Na	0.15	0.18	0.13	0.11	0.17	0.13	0.10	0.22	0.21
K	0.84	0.83	0.83	0.83	0.83	0.82	0.81	0.82	0.79
Ba	0.23	0.18	0.23	0.12	0.14	0.18	0.17	0.17	0.20
Sum A	1.23	1.20	1.21	1.07	1.16	1.14	1.08	1.22	1.21
Cations	14.78	14.77	14.80	14.70	14.82	14.79	14.81	14.74	14.85
Mg/(Mg+Fe⁺²)	0.64	0.64	0.71	0.69	0.64	0.71	0.66	0.65	0.63
Fe⁺²/(Fe⁺²+Mg)	0.36	0.36	0.29	0.31	0.36	0.29	0.34	0.35	0.37

Obs: H_2O recalculated by stequiometric. Fe_2O_3 only recalculated for samples with deficiency in tetrahedric site.

ANEXO: Análises de microsonda de cristais de biotita em ocelli e segregados (amostras BYA123 e BYA127), lamprofiros de R Grande

	1	2	3	3	3	6
Amostra	BYA123.3D.BTOCE L.417	BYA127.2.BTGLO .527	BYA127.4.BTCOREGL O.577	BYA127.4.BTRIMGL O.578	BYA127.7.BTCARBO 591	BYA127.6A.COREOL V.546
<i>SiO</i> ₂	34.67	34.62	36.96	37.19	33.61	39.48
<i>TiO</i> ₂	5.76	5.57	3.68	2.58	5.59	0.50
<i>Al</i> ₂ <i>O</i> ₃	15.52	8.71	10.11	10.78	14.51	11.79
<i>FeO</i>	13.37	32.04	26.65	22.72	14.36	13.33
<i>Fe</i> ₂ <i>O</i> ₃	0.00	0.00	0.00	0.00	0.00	0.00
<i>Cr</i> ₂ <i>O</i> ₃	0.00	0.00	0.00	0.00	0.00	0.00
<i>MnO</i>	0.11	0.52	0.43	0.50	0.25	0.33
<i>MgO</i>	15.63	4.46	9.37	12.38	14.14	20.21
<i>CaO</i>	0.08	0.24	0.21	0.47	0.29	0.04
<i>Na</i> ₂ <i>O</i>	0.35	0.00	0.00	0.00	0.27	0.81
<i>K</i> ₂ <i>O</i>	9.22	9.03	9.49	9.59	9.10	8.56
<i>BaO</i>	0.43	0.00	0.00	0.00	2.08	0.20
<i>Cl</i>	0.02	0.00	0.02	0.01	0.01	0.01
<i>H</i> ₂ <i>O</i>	3.98	3.61	3.81	3.85	3.85	4.04
<i>Total</i>	99.12	98.81	100.72	100.06	98.04	99.29
Fórmula estrutural baseada em 22 oxigênios						
Si	5.21	5.73	5.80	5.77	5.22	5.85
Al^{IV}	2.75	1.70	1.87	1.97	2.65	2.06
Fe⁺³	0.05	0.57	0.33	0.26	0.13	0.09
Sum T	8.00	8.00	8.00	8.00	8.00	8.00
Al^{VI}	0.00	0.00	0.00	0.00	0.00	0.00
Fe⁺²_t	1.63	3.86	3.16	2.69	1.73	1.56
Fe⁺³	3.50	1.10	2.19	2.86	3.27	4.46
Ti	0.01	0.07	0.06	0.07	0.03	0.04
Mg	0.65	0.69	0.43	0.30	0.65	0.06
Mn	0.00	0.00	0.00	0.00	0.00	0.00
Sum O	5.80	5.73	5.85	5.92	5.69	6.12
Ca	0.01	0.04	0.04	0.08	0.05	0.01
Na	0.10	0.00	0.00	0.00	0.08	0.23
K	0.88	0.95	0.95	0.95	0.90	0.81
Ba	0.05	0.00	0.00	0.00	0.25	0.02
Sum A	1.05	1.00	0.98	1.03	1.28	1.07
Cations	14.84	14.73	14.83	14.94	14.97	15.19
Mg/(Mg+Fe⁺²)	0.68	0.22	0.41	0.52	0.65	0.74
Fe⁺²/(Fe⁺²+Mg)	0.32	0.78	0.59	0.48	0.35	0.26

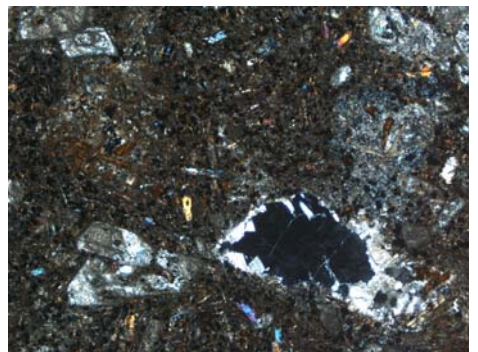
Obs: *H*₂*O* recalculated by stequiometrie. *Fe*₂*O*₃ only recalculated for samples with deficiency in tetraedric site.



ΑΡΑΤΤΑ

ANEXO: Análises de microsonda de cristais de apatita,
lamprofiros de R o Grande

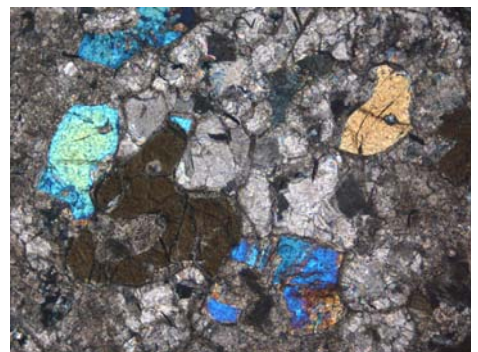
	<i>Fenocristal</i>	<i>em fenocristal de augite</i>	
	<i>1</i>	<i>2</i>	<i>3</i>
<i>SiO₂</i>	0.00	0.13	0.12
<i>Al₂O₃</i>	0.00	0.00	0.01
<i>MgO</i>	0.14	0.38	0.34
<i>Fe₂O₃</i>	0.34	0.34	0.32
<i>CaO</i>	53.31	53.83	54.50
<i>Na₂O</i>	0.35	0.06	0.12
<i>K₂O</i>	0.00	0.01	0.03
<i>P₂O₅</i>	41.67	40.14	40.51
<i>Cl</i>	0.24	0.13	0.11
<i>SrO</i>	1.08	0.43	0.31
<i>BaO</i>	0.01	0.00	0.00
<i>H₂O₃</i>	1.70	1.69	1.72
<i>Total</i>	98.83	97.14	98.07
Cations por f�rmula unidade			
Na	0.12	0.02	0.04
Si	0.00	0.02	0.02
Al	0.00	0.00	0.00
Mg	0.04	0.10	0.09
K	0.00	0.00	0.01
P	6.02	5.90	5.90
Ca	9.74	10.02	10.04
Cl	-0.05	-0.03	-0.02
Sr	0.11	0.04	0.03
Fe	0.05	0.05	0.05
Ba	0.00	0.00	0.00
H	0.56	0.57	0.57
<i>Cations</i>	16.57	16.69	16.71



ANALCIMA

ANEXO: Análises de microsonda de cristais de analcima, lamprofiros de R o Grande

	<i>Borda de anf�b�lio</i>		<i>em Ocelli</i>						<i>em segrados ricos em biotita</i>		<i>Microfenocristais</i>		
	1	2	3	4	5	6	7	8	9	10	11	12	
<i>SiO₂</i>	55.51	57.23	53.20	55.76	52.18	54.66	59.20	61.92	58.02	62.77	50.61	52.08	
<i>Al₂O₃</i>	24.57	24.63	23.73	21.17	24.92	25.46	24.85	22.71	24.91	22.47	22.72	23.19	
<i>FeO</i>	0.14	0.01	0.07	0.09	0.63	0.51	0.13	0.18	0.16	0.26	0.04	1.23	
<i>CaO</i>	0.16	0.00	0.30	0.11	0.63	0.69	0.05	0.01	0.13	0.08	0.29	0.32	
<i>Na₂O</i>	12.94	13.26	12.93	13.96	10.92	10.46	13.58	12.18	12.24	10.91	11.46	11.85	
<i>K₂O</i>	0.12	0.04	0.25	0.12	0.34	0.20	0.08	0.08	0.05	0.09	0.27	0.55	
<i>Total</i>	93.43	95.28	90.49	91.30	89.62	91.99	97.90	97.08	95.50	96.69	85.42	89.23	
Cations recalculados em base a 7 oxig�nios													
Si	2.20	2.39	2.23	2.33	2.18	2.29	2.48	2.59	2.43	2.63	2.12	2.18	
Al	1.15	1.21	1.17	1.04	1.23	1.26	1.22	1.12	1.23	1.11	1.12	1.14	
Sum T	3.35	3.61	3.39	3.38	3.41	3.54	3.70	3.71	3.65	3.73	3.24	3.32	
Fe⁺²_t	0.00	0.00	0.00	0.00	0.02	0.02	0.00	0.01	0.01	0.01	0.00	0.04	
Ca	0.01	0.00	0.01	0.01	0.03	0.03	0.00	0.00	0.01	0.00	0.01	0.01	
Na	1.00	1.08	1.05	1.13	0.89	0.85	1.10	0.99	0.99	0.88	0.93	0.96	
K	0.01	0.00	0.01	0.01	0.02	0.01	0.00	0.00	0.00	0.00	0.01	0.03	
Sum A	1.01	1.08	1.08	1.15	0.95	0.91	1.11	1.00	1.01	0.90	0.96	1.05	
<i>Cations</i>	4.36	4.69	4.47	4.52	4.37	4.45	4.81	4.71	4.66	4.64	4.20	4.37	



CARBONATO

ANEXO: Análises de microsonda de cristais de carbonato da amostra P8, lamproiros de Río Grande

	<i>Em Ocelli tipo-I</i>							<i>Segregados</i>	<i>Matriz</i>				<i>Carbonato do xenólito</i>		<i>Borda de anfibólio</i>		
	<i>1</i>	<i>2</i>	<i>3</i>	<i>4</i>	<i>5</i>	<i>6</i>	<i>7</i>		<i>1</i>	<i>2</i>	<i>3</i>	<i>4</i>	<i>5</i>	<i>6</i>		<i>7</i>	
<i>FeCO₃</i>	12.70	9.69	7.08	12.12	8.62	15.77	17.70	15.87	45.15	13.02	16.01	14.05	12.85	14.68	11.77	9.21	9.84
<i>MnCO₃</i>	0.56	0.68	0.75	0.57	1.12	0.91	0.96	1.16	0.13	1.09	1.36	1.10	0.73	1.82	0.22	0.37	0.80
<i>MgCO₃</i>	30.73	34.43	33.91	33.86	30.92	30.07	28.63	30.19	51.40	28.35	26.32	30.73	33.19	28.74	29.77	35.41	35.26
<i>CaCO₃</i>	51.20	52.15	55.63	52.19	57.30	50.81	50.39	51.67	2.07	54.75	52.64	51.35	51.11	51.77	53.32	51.89	51.58
<i>Sr₂CO₃</i>	2.06	0.80	0.45	0.15	0.04	0.45	0.47	0.62	0.02	0.96	0.71	0.48	0.39	1.13	3.09	2.50	0.33
<i>BaCO₃</i>	0.00	0.00	0.18	0.00	0.00	0.00	0.00	0.00	0.05	0.00	0.00	0.00	0.00	0.05	0.22	0.07	0.02
<i>TOTAL</i>	98.57	99.10	99.34	100.25	99.35	99.35	99.47	100.88	100.16	99.50	98.35	99.05	99.61	99.52	99.72	100.80	99.18
<i>Ca</i>	1.02	1.02	1.08	0.95	1.07	1.00	1.00	1.01	0.04	1.05	1.05	0.52	1.00	1.03	1.05	1.00	0.97
<i>Mg</i>	0.73	0.80	0.78	0.73	0.68	0.71	0.67	0.70	1.09	0.65	0.63	0.37	0.77	0.68	0.70	0.81	0.79
<i>Fe</i>	0.22	0.16	0.12	0.19	0.14	0.27	0.30	0.27	0.70	0.22	0.28	0.12	0.22	0.25	0.20	0.15	0.16
<i>Ba</i>	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
<i>Sr</i>	0.03	0.01	0.01	0.00	0.00	0.01	0.01	0.01	0.00	0.01	0.01	0.00	0.01	0.02	0.04	0.03	0.00
<i>Mn</i>	0.01	0.01	0.01	0.01	0.02	0.02	0.02	0.02	0.00	0.02	0.02	0.01	0.01	0.03	0.00	0.01	0.01
<i>Cations</i>	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00

ANEXO III: COORDENADAS AMOSTRAS

Amostras	Coordenadas geográficas					Amostras	Coordenadas geográficas						
Dacitos de El Niño Muerto													
BNM-08	23 44	16.43	S	66 10	39.34	W	BRB-47B	25 19	23.19	S	66 14	08.12	W
BNM-09	23 44	16.43	S	66 10	39.34	W	BRB-64	24 57	20.51	S	65 58	45.50	W
BNM-10	23 44	16.43	S	66 10	39.34	W	BRB-65	24 57	20.51	S	65 58	45.50	W
BNM-11	23 44	16.48	S	66 12	59.10	W	BRB-66	24 57	20.51	S	65 58	45.50	W
BNM-12	23 44	16.48	S	66 12	59.10	W	BRB-67	24 57	18.29	S	65 40	59.74	W
BNM-13	23 44	16.48	S	66 12	57.79	W	BRB-68	24 57	13.76	S	66 05	04.23	W
BNM-14	23 44	16.48	S	66 12	57.79	W	Lamprofiros de Coraya						
BNM-15	23 44	49.87	S	66 13	18.71	W	BCOY-49A	23 10	03.74	S	65 26	29.89	W
BNM-16	23 44	56.89	S	66 13	17.74	W	BCOY-49B	23 10	03.74	S	65 26	29.89	W
BNM-17	23 50	02.47	S	66 15	14.05	W	BCOY-50	23 10	03.74	S	65 26	29.89	W
BNM-18	24 14	32.06	S	66 20	09.52	W	BCOY-51A	23 10	05.96	S	65 26	32.24	W
Basaltos de Río Blanco													
BRB-19	25 48	41.79	S	65 41	19.42	W	BCOY-51B	23 10	05.96	S	65 26	32.24	W
BRB-20	26 01	22.28	S	65 49	41.24	W	BCOY-52	23 10	09.58	S	65 26	34.40	W
BRB-21	24 57	30.11	S	65 49	46.88	W	BCOY-53	23 10	13.91	S	65 26	37.02	W
BRB-22	24 57	30.11	S	65 49	46.88	W	BCOY-54	23 10	13.98	S	65 26	55.94	W
BRB-23	24 57	23.03	S	65 57	47.79	W	BCOY-55	23 10	13.98	S	65 26	55.94	W
BRB-24	24 57	24.98	S	65 57	30.75	W	Lamprofiros do Vale de Yacoraite						
BRB-25	24 57	24.98	S	65 57	30.75	W	BYA-121A	23 20	13.3	S	65 25	43.4	W
BRB-26	24 57	26.76	S	65 57	29.14	W	BYA-121B	23 20	13.0	S	65 25	43.4	W
BRB-27	24 57	30.37	S	65 57	29.60	W	BYA-122	23 20	13.3	S	65 25	43.4	W
BRB-28	24 57	30.37	S	65 57	29.60	W	BYA-123	23 20	13.3	S	65 25	43.4	W
BRB-29	24 57	33.20	S	65 57	30.67	W	BYA-124A	23 20	13.3	S	65 25	43.4	W
BRB-30	24 57	33.20	S	65 57	30.67	W	BYA-124B	23 20	13.3	S	65 25	43.4	W
BRB-31	24 57	24.62	S	65 57	33.71	W	BYA-125	23 19	58.8	S	65 26	12.5	W
BRB-32	24 57	27.02	S	65 57	29.32	W	BYA-126	23 19	58.8	S	65 26	12.5	W
BRB-33A	24 57	27.02	S	65 57	29.32	W	BYA-127	23 19	55.8	S	65 26	36.1	W
BRB-33B	24 57	27.02	S	65 57	29.32	W	BYA-128	23 19	55.8	S	65 26	36.1	W
BRB-34	24 57	36.77	S	65 57	28.14	W	Lamprofiros do Vale de Huichaira						
BRB-35	24 57	46.79	S	65 57	35.62	W	BHUI-131	23 33	51.1	S	65 27	32.8	W
BRB-36	24 57	50.49	S	65 57	36.59	W	BHUI-132	23 33	51.1	S	65 27	32.8	W
BRB-37	24 57	52.99	S	65 57	33.77	W							
BRB-38	24 57	51.79	S	65 57	31.91	W							
BRB-39	24 57	51.79	S	65 57	31.91	W							
BRB-40	24 57	51.79	S	65 57	31.91	W							
BRB-41	24 57	30.34	S	65 57	29.39	W							
BRB-42	24 57	30.34	S	65 57	29.39	W							
BRB-43	24 57	26.76	S	65 57	29.14	W							
BRB-44	25 08	18.40	S	66 10	05.00	W							
BRB-45	25 08	18.40	S	66 10	05.00	W							
BRB-46	25 08	18.40	S	66 10	05.00	W							
BRB-47	25 19	23.19	S	66 14	08.12	W							